

1 **The migration path of Gondwanian dinosaurs toward Adria: new insights from the**
2 **Cretaceous of NW Sicily (Italy)**

3 Vincenzo Randazzo (1); Pietro Di Stefano (1, 4); Felix Schlagintweit (2)

4 Simona Todaro (1); Simona Cacciatore (3); Giuseppe Zarcone (1)

5 1) University of Palermo, Department of Earth and Marine Sciences, 90123 Palermo, Italy

6 2) Lerchenauerstr. 167, 80935 Munich, Germany

7 3) Eni S.p.A. Upstream and Technical Services, 20097 San Donato Milanese, Italy

8 4) Istituto Nazionale di Geofisica e Vulcanologia (INGV), 90100 Palermo, Italy

9

10 **Abstract**

11 The increasing dinosaur record from Italy questioned classic paleogeographic scenarios for the
12 Central Mediterranean area and suggest the proximity of landmass areas and a geographical
13 connection between Gondwana and Laurasia during Cretaceous times. Besides several track-sites
14 and exceptionally-preserved specimens (e.g. *Scipionyx samniticus*), the Italian dinosaur record also
15 consists of isolated bones, among which the bone fragment of a theropod discovered in north-
16 western Sicily. The bone occurs in a shallow-water carbonate succession (i.e. Pizzo Muletta,
17 Palermo Mountains) pertaining to the Panormide Carbonate Platform (PCP). The bone was
18 previously ascribed to the Cenomanian, strongly supporting the hypothesis of a land bridge
19 connecting Gondwana and Adria via PCP. More recently, new sedimentological and
20 biostratigraphic studies on the Pizzo Muletta succession have been carried out. The obtained results
21 allow to predate the stratigraphic position of the dinosaur bone to the late Aptian–early Albian and
22 to assess a detailed Aptian–Cenomanian evolution of this sector of the PCP. In particular, the
23 karstic overprint of Cenomanian rudist limestones indicate a subaerial exposure of the platform
24 preceding its drowning during latest Cenomanian times. The new assumptions allow to extend the
25 temporal duration of the intermittent land bridge between Gondwana and Laurasia at least from

26 Aptian to Cenomanian times and to add further evidences of the dominant tectonic control affecting
27 the Western Tethys during Cretaceous times.

28 Corresponding author: vincenzo.randazzo04@unipa.it

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30

31 **1. Introduction**

32 The geodynamic history and the Mesozoic palaeogeography of the Mediterranean area were broadly
33 debated during the last four decades (Anderson, 1987; Biju-Duval et al., 1976; Bosellini, 2002;
34 Casero and Roure, 1994; Channell et al., 1979; Finetti, 1982, 2005; Frizon de Lamotte et al., 2011;
35 Masse et al., 1993; Nicosia et al., 2007; Patacca and Scandone, 2004; Rosenbaum et al., 2004;
36 Wortmann et al., 2001; Zarcone et al., 2010). Classic palaeogeographic reconstructions depicted the
37 Western Tethys as punctuated by a number of restricted and isolated shallow-water areas separated
38 from Gondwana and Laurasia by deep seaways (Biju-Duval et al., 1977; Catalano et al., 2001;
39 Dercourt et al., 1986; Dewey et al., 1973; Finetti, 1985; Finetti et al., 1996; Passeri et al., 2005;
40 Patacca and Scandone, 2007; Stampfli and Borel, 2004; Stampfli and Mosar, 1999; Yilmaz et al.,
41 1996; Ziegler, 1988). Several evidences of dinosaurs (e.g. Dal Sasso, 2003; Dal Sasso and Signore,
42 1998; Dal Sasso et al., 2016; Garilli et al., 2009), especially track-sites (Dalla Vecchia, 1999, 2001;
43 Dalla Vecchia and Venturini, 1995; Nicosia et al., 2000; among others see Petti et al., 2020 for an
44 exhaustive review), from the latest Jurassic to Late Cretaceous carbonate-platform domains of the
45 peri-Adriatic region (e.g. Panormide, Apennine, Apulian platforms) have questioned the classic
46 palaeogeographic models of the Mediterranean area. Indeed, the dinosaur record from Italy
47 provided evidence for a structured and diversified dinosaur assemblage needing freshwater
48 availability, food supply and potential nesting sites implying a certain amount of landmass area
49 (Nicosia et al., 2007). The age of dinosaur tracks and bones from Italy substantially agrees with
50 structural, sedimentological and stratigraphic evidences in suggesting that the peri-Adriatic
51 carbonate platforms constituted a wide land-mass area (*Adria sensu* Channel et al., 1979) connected

52 to Gondwana during the early Cretaceous throughout an intermittent geographical connection
53 (Bosellini 2002; Dal Sasso et al., 2016; Frizon de Lamotte et al., 2011; Nicosia et al., 2000, 2007;
54 Petti et al., 2006, 2008, 2020; Vitale et al., 2018; Zarcone, 2008; Zarcone and Di Stefano, 2008;
55 Zarcone et al., 2010). The migration path of Gondwanian dinosaurs toward Adria is supported,
56 among other evidences, by a theropod bone discovered in the Cretaceous shallow-water limestones
57 of the Panormide Carbonate Platform (PCP hereafter). This finding at Pizzo Muletta, near Capaci
58 (Palermo Mountains, Sicily) is ascribed to the Cenomanian (Garilli et al., 2009) and has pointed the
59 PCP as key area to the knowledge of the land bridge between Africa and Adria (Canudo et al.,
60 2009; Zarcone and Di Stefano, 2010; Zarcone et al., 2010). The present paper deals with the
61 sedimentological and biostratigraphic re-study of the Pizzo Muletta section. The obtained results
62 allow to redefine the stratigraphic position of the theropod bone and to discuss the tectono-
63 sedimentary evolution of the PCP in the frame of the geodynamic context of the Western Tethys.

64

65 **2. Geological setting**

66 Shallow-water limestones of Cretaceous age from northern Sicily belong to tectonic units of the
67 Maghrebian thrust-and-fold belt (Fig. 1A) and pertain to the PCP, a palaeogeographic unit located
68 at the westward termination of the Ionian Tethys during the Triassic (Todaro et al., 2017).
69 Thousand metres of shallow-water limestones pertaining to this paleogeographic unit crop out in the
70 Madonie and Palermo Mountains (Fig. 1B), and were deposited until Eocene times (Abate et al.,
71 1978, 1982a; Broquet, 1968; Caflish, 1966; Catalano and D'Argenio, 1982; Catalano et al., 1996,
72 Di Stefano and Ruberti, 2000). The PCP-derived series also occur offshore western Sicily, along the
73 Tyrrhenian margin (D'Argenio, 1999; Finetti, 2005; Gattacceca and Speranza, 2002; Pepe et al.,
74 2004) and in north-eastern Sicily in the subsurface of the Nebrodi Mountains (Maragone and
75 Bellafontana boreholes, Bianchi et al., 1987; location in Fig. 1B).

76 **INSERT FIG. 1**

77 During Late Triassic-Early Jurassic times, the PCP was part of a wider rimmed Bahamian-like
78 carbonate platform (Siculo-Tunisian platform *sensu* Di Stefano et al., 1996) leading to the
79 emplacement of a thick succession of peritidal limestones and dolostones (Todaro et al., 2017;
80 2018). The PCP fed throughout the Triassic–Eocene times the adjacent slope-to-basin areas namely
81 Imerese Basin (Abate et al., 1982b; Camoin, 1982; Catalano and D’Argenio, 1982; Grasso et al.,
82 1978; Ogniben, 1960) and Western Sicily Cretaceous Escarpment (WSCE hereafter; Randazzo et
83 al., 2020a, 2020b). Subaerial exposures related to tectonic uplifts (Basilone and Sulli, 2018;
84 Zarcone and Di Stefano, 2010) caused two major unconformities in the Middle Jurassic and Late
85 Cretaceous (Di Stefano and Ruberti, 2000) pointing to an anomalous subsidence history for the PCP
86 during the Mesozoic (Zarcone, 2008; Zarcone and Di Stefano, 2008) when compared to the
87 adjacent palaeogeographic domains of Sicily (e.g. Hyblean, Yellin Dror et al. 1997).

88

89 **3. Material and methods**

90 The carbonate succession exposed at Pizzo Muletta has been logged and sampled. The complex
91 morphology of the outcrops hinders to study and sample a single continuous section, the whole
92 studied succession thus consists of three distinct transects being sections M, PM and MZ (Fig. 2A),
93 whose total thickness is about 135 m. The lower part of the succession (section M) is well exposed
94 along the E-W oriented cliff of Pizzo Muletta facing north (Fig. 2B). A second segment (section
95 PM) has been studied inside an abandoned quarry adjacent to Pizzo Muletta (Fig. 2C) that extracted
96 the Cretaceous limestones to produce calcareous gravels and sands. A third upper segment (section
97 MZ) has been studied along a road known as Via Zarcate, south of Pizzo Muletta (Fig. 2A). The
98 correlation between the three sections has been based on bed attitude, distance and topographic
99 elevation between correlative points, taking also into account some extensional faults with metre-
100 scale displacement that occur between the first and second section. Facies and textural analysis
101 were based on the classifications proposed by Dunham (1962) and Embry and Klovan (1971). The
102 accurate analysis of about 50 thin sections allowed to complement the field observations. The

103 microfacies characterisation was performed together with the identification of skeletal and non-
104 skeletal components (Frigel, 2004), under an optical microscope (Leitz Laborlux 12 Pol).
105 Photomicrographs were taken using the camera Zeiss AxioCam ICc 5. New age constraints for the
106 Pizzo Muletta limestones were defined on the base of the observed foraminifera and algae
107 assemblages following the biostratigraphic schemes proposed for the Western Tethys realm
108 (Chiocchini, 2008; Chiocchini et al., 2008; Mancinelli and Chiocchini, 2006; Premoli Silva and
109 Verga, 2004; Velić, 2007; see Fig. 3).

110 Comprehensive photographic documentation, rock samples and thin sections are stored in the
111 Laboratory of Stratigraphic Geology at the Department of Earth and Marine Sciences, University of
112 Palermo, Italy (collection V. Randazzo).

113 **INSERT FIG. 2**

114 **INSERT FIG. 3**

115

116 **4. The Pizzo Muletta Section**

117 This chapter reports the previous contributions on the Cretaceous succession of Pizzo Muletta and
118 describe new sedimentological data concerning the identified carbonate facies and their stacking
119 patterns.

120

121 *4.1. Previous contributions*

122 The theropod bone was found in a wall of Grotta Lunga, a small cave located on the northern side
123 of Pizzo Muletta (38°10'11.08"N 13°13'55.01", Fig. 4). Grotta Lunga has a horizontal development
124 of about 10 m and a variable height ranging between 4 and 2 m. The Pleistocene and Holocene
125 sediments filling the cave were excavated by Fabiani (1928) which recovered a fossil assemblage of
126 Pleistocene mammalian vertebrates. The histology of the dinosaur bone from Pizzo Muletta and the
127 main sedimentological and biostratigraphic features of the Pizzo Muletta limestones were addressed
128 by Garilli et al. (2009). An about 80 m thick carbonate succession was described as consisting of

129 mudstone/floatstone alternating to wackestone to rudstone/packstone. The age attribution was based
130 on the identified rudists *Sauvagesia nicaisei* (COQUAND), *Eoradiolites* sp., *Ichtiosarcolites*
131 *bicarinatus* (GEMMELLARO), *I. monocarinatus* (SLIŠKOVIĆ) and gastropods (nerineids)
132 *Eunerinea ernesti* (PARONA) *Multiptyxis olisiponensis* (SHARPE) and *Iteria* sp., thus suggesting a
133 Cenomanian age for the whole Pizzo Muletta section. The rudist assemblages mainly occur in the
134 upper part, while the nerineid species are widespread along the Pizzo Muletta section. In particular,
135 the co-occurrence of *Eunerinea ernesti* and *Multiptyxis olisiponensi* with rudist fragments suggested
136 a backreef depositional environment (Garilli et al., 2009; Sirna, 1995). A good match with another
137 section of Cenomanian age pertaining to the Panormide domain and exposed at Mt. Pellegrino (Di
138 Stefano and Ruberti, 2000) has been assessed by the authors on the base of the macrofossil
139 assemblage and sedimentological features. With regard to the dinosaur bone, analyses on the tissue
140 carried out by Garilli et al (2009), revealed specific microanatomical features that point to a
141 theropod affinity (i.e. “the open medullary cavity and the laminar organization”; see Bybee et al.,
142 2006; Chinsamy, 1990, Horner et al., 2001; Horner and Padian, 2004) and a non-juvenile stage for
143 the Pizzo Muletta specimen (i.e. “well-developed primary osteons in the laminar fibro-lamellar
144 bone”; see Klein and Sander, 2008; Sander, 2000).

145 **INSERT FIG. 4**

146

147 *4.2 Facies analysis of the Pizzo Muletta succession*

148 Seven different facies types have been distinguished, being: stromatolitic bindstone, gastropod
149 floatstone, gastropod rudstone, black-pebble floatstone, black-pebble rudstone, rudist floatstone and
150 rudist rudstone. Their stratigraphic distribution (see Fig. 5) provide information about the
151 depositional environments of the three sections constituting the Pizzo Muletta succession (i.e.
152 sections M, PM and MZ).

153 Facies descriptions and related environmental interpretations for the three sections follow. The main
154 macro- and microfacies features are depicted in Figs. 6 and 7 respectively.

155

156 **INSERT FIG. 5**

157

158 *Section M*

159 This section crops out along the northern cliff of Pizzo Muletta and exposes about 80 m of metre-
160 scale, parallel-bedded grey limestones, dipping 30° to SE (Fig. 2B). Along the cliff several other
161 caves of marine origin and parallel-bedded discontinuity surfaces occur along the section. The latter
162 appear as regular and sometimes open fissures a few cm thick. In most cases, the original filling is
163 replaced by Pleistocene calcarenites (Fig. 6A). These discontinuity surfaces are usually overlain by
164 thin levels of laminated bindstones (Figs. 5 and 6B) that alternate to skeletal floatstone, less often
165 rudstones, rich in small gastropods (Figs. 5 and 6C). The bindstones consist of micritic laminae of
166 microbial origin alternating with thin peloidal levels with ostracods. *Fenestral fabric* is commonly
167 associated to the laminated bindstones (Fig. 6B). In the lower and uppermost parts of the section
168 some levels of rudstone with black pebbles occur (Fig. 6D). Thin section analysis of the gastropod
169 rudstone/floatstone mainly revealed a packstone/wackestone texture rich in gastropods, benthic
170 foraminifers, ostracods and, alternatively, rudist fragments (Fig. 7A and B). Grotta Lunga and the
171 theropod bone occur in the upper part of section M (Fig. 5). The alteration/encrustation of the wall
172 surfaces of the cave hampers the direct observation of the facies type in which the bone occurs.
173 However, the correlative bed outside the cave is a floatstone/rudstone about 50 cm thick with
174 reverse grading, rich in mollusc and echinoderm fragments, orbitolinids, and bacinellid textures
175 *sensu* Maurin et al. (1985). This facies is intercalated in muddier facies (i.e. floatstones) rich in
176 benthic foraminifers and dasycladalean algae (Fig. 5). Rare vug-type dissolution cavities filled up
177 by barren micrite were randomly observed (Fig. 8A).

178 Environmental interpretation.

179 The facies association in this tract of the Pizzo Muletta succession is typical of a low energy
180 peritidal environment. The sharp discontinuity surfaces were most probably created by the

181 Pleistocene erosion of marly horizons related to paleosoils, as commonly observed in coeval
182 successions showing comparable facies (e.g. Raffo Rosso, Nicchitta, 1999).

183 The facies stacking pattern shows a cyclical arrangement of both symmetric and asymmetric
184 deepening-shallowing cycles. This low-energy environment appears randomly punctuated by
185 higher-energy events as suggested by the occurrence of mollusc (Fig. 6C) or black-pebble (Fig. 6D)
186 rudstones (storm layers). In particular, the rudstone/floatstone bed containing the theropod bone can
187 be interpreted as a tempestite on the base of the observed reverse grading. This supports the idea
188 that skeletal components of the dinosaurs were most probably transported by a tidal or subaerially-
189 exposed area adjacent to the lagoon.

190

191 *Section PM*

192 This segment of the Pizzo Muletta section is about 37 m thick and has been logged inside the
193 abandoned quarry (Fig. 2C). The lower part consists of parallel beds of skeletal rudstones
194 alternating with floatstones containing dominant black pebbles and, alternatively, rudist fragments
195 (Fig. 6E-G). Average bed thickness is of about 80-90 cm. Upwards, there is a gradual increase of
196 black pebble rudstones facies (Fig. 5), to such an extent that were observed distinctive beds almost
197 completely made up of black-pebble conglomerates (Fig. 6E and F). In thin section, the black-
198 pebble rudstone/floatstone mainly consists of grainstone/packstone with radiolitid fragments often
199 affected by microboring, abundant orbitolinids (Fig. 7C and D), and calcareous algae (Fig. 7E).

200 Centimetre-scale, dissolution vugs are randomly observed. They show a thin isopachous rim of
201 calcite cements and a subsequent filling of micrite (Fig. 8B and C).

202 Environmental interpretation.

203 The facies stacking in this section shows a gradual shifting of lagoonal environments to more open
204 and high-energy conditions that accounts for a relative sea-level rise. The subsequent flooding of
205 adjacent emerged lands with ponds could also explain the observed increase of black pebbles in the
206 sediment composition.

207

208 *Section MZ*

209 The last segment of the studied section at Pizzo Muletta (ca. 18 m thick) was logged along a road
210 namely “via Zarcate” (Fig. 2A) and mostly consists of alternating rudist floatstone and rudstone
211 (Figs. 5 and 6H). Rudist fragments are the main constituent, though loose and well-preserved
212 specimens often occur (Fig. 6I). Rudist-rich beds are thicker and more frequent in the upper part of
213 section MZ. Gastropods are also abundant and widespread along the whole section, while large
214 Chondrodonta mainly occur in the lower beds. Corals are rare. Thin section analysis of the rudist
215 rudstone/floatstone mainly revealed a grainstone/packstone texture with rudist fragments and
216 benthic foraminifers (Fig. 7G). The whole section is crosscut by neptunian dykes with irregular and
217 smoothed walls, filled up by laminated mudstones and wackestones with planktonic foraminifers
218 (Figs. 7H and 8D-G).

219 Environmental interpretation.

220 The facies associations of section MZ are well tuned with that of the section PM as they record a
221 progressive shifting towards more external and high-energy zones of the platform that consist of
222 biostromal bodies dominated by large rudists associated with huge amounts of skeletal debris.
223 Rudists often appear as toppled specimens, thus suggesting a continuous sediment removal between
224 the organisms as described from different Upper Cretaceous Periadriatic platforms (Ruberti et al.,
225 2006). The neptunian dykes suggest a tectonic uplift of the platform and a karstic dissolution
226 overprint followed by a subsequent drowning into deeper marine settings.

227 **INSERT FIG. 6**

228 **INSERT FIG. 7**

229 **INSERT FIG. 8**

230

231 **5. Biostratigraphy**

232 The whole Pizzo Muletta succession shows an abundant and well-diversified fossil content
233 consisting of various systematic groups (e.g. rudists, calcareous algae, benthonic and planktonic
234 foraminifers). These assemblages allow the correlation of the three studied sections comprising the
235 Pizzo Muletta succession (Fig. 9). The observed assemblages are shortly described below; the
236 related biomarkers and the most representative taxa are shown in Figs. 10 and 11.

237 **INSERT FIG.9**

238

239 5.1 Section M assemblage

240 The floatstone to rudstone beds of Section M show a diverse assemblage of benthic foraminifers
241 (Fig. 10), which are commonly associated with gastropods and ostracods. Calcareous algae such as
242 *Cayeuxia* sp., *Salpingoporella* sp., as well as bryozoans and bacinellid textures, are occasionally
243 present. The foraminifers are represented by *Spiroloculina* sp., *Spiroloculina cretacea* REUSS (Fig.
244 10A), Haplophragmoididae indet (?*Debarina hahounerensis* Fourcade et al.) (Fig. 10B),
245 *Haplophragmoides* sp., *Nezzazata isabellae* ARNAUD VANNEAU AND SLITER (Fig. 10C and
246 D), *Bolivinopsis* sp., *Quinqueloculina* sp., *Quinqueloculina histri* NEAGU (Fig. 10E), *Glomospira*
247 sp. (Fig. 10F), *Pseudocyclamina* sp., *Pseudonummoloculina* sp., *Akcaya* sp., *Vercorsella* gr.
248 *camposaurii-laurentii* (SARTONI AND CRESCENTI), *Cuneolina* sp., *Cuneolina* aff. *sliteri*
249 ARNAUD VANNEAU AND SLITER (Fig. 10G-I), *Simplorbitolina aquitanica* (SCHROEDER),
250 ?*Paracoskinolina* sp. (Fig. 10J) and other undetermined orbitolinids (Fig. 10K). The bindstone
251 beds of Section M are sometimes rich in ostracods while completely lack benthic foraminifers or
252 other biota. Greenish marls are barren.

253

254 5.2 Section PM assemblage

255 In this section rudist fragments become dominant and mainly refer to Radiolitidae family. Section
256 PM also shows a rich and diversified benthic foraminifereral assemblage similar to what is
257 observed through the section M. These include *V. gr. camposaurii-laurentii*, *Cuneolina* sp., *C.*

258 *sliteri*, ?*Debarina hahounerensis*, *Haplophragmoides* sp., *Quinqueloculina* sp., *Q. histri*, *Nezzazata*
259 *isabellae* (Fig. 10L), *Akcaya* sp., *Akcaya minuta* (HOFKER) (Fig. 10M), *Pseudonummoloculina* sp.
260 (Fig. 10N). Orbitolinids are represented by different genera: *Simplorbitolina aquitanica* (Fig. 10O
261 and P), Orbitolinidae indet. (Fig. 10Q), *Mesorbitolina birmanica* (SAHNI) (Fig. 10R and S),
262 ?*Paracoskinolina* sp. (Fig. 10T), *Orbitolinopsis cuvillieri* MOULLADE (Fig. 10U and V). The
263 occurrence of *Mesorbitolina* specimens is restricted just to a specific interval (samples PMA3 and
264 PM12), in which the corresponding microfacies texture is a grainstone and the assemblage becomes
265 oligotypic and dominated by the representatives of the genus. Calcareous algae are present and
266 mostly occur in the upper part of the section. They include *Triploporella* sp. (Fig. 10W),
267 *Morelletpora turgida* (RADOIČIĆ) (Fig. 10X), *Cylindroporella? lyrata* MASSE AND LUPERTO-
268 SINNI (Fig. 10Y). Ostracods, gastropods and *bacinella* structures were also observed.

269 **INSERT FIG. 10**

270

271 5.3 Section MZ assemblage

272 The assemblage is rich in molluscs such as rudists, gastropods and chondrodonta, though the
273 assemblage is dominated by the rudists (Garilli et al., 2009). The latter mostly refer to Caprinidae
274 family (Fig. 6I). Benthic foraminifers such as *Spiroloculina* sp. (Fig. 11A, B)
275 *Pseudonummoloculina?* sp. (Fig. 11C) *Vercorsella* sp. (Fig. 11D), ?*Debarina hahounerensis* (Fig.
276 11E), *Nezzazata isabellae* (Fig. 11F), *Cuneolina* sp. (Fig. 11G-I), Orbitolinidae indet. (Fig. 11J-M),
277 *Coscinophragma?* sp. (Fig. 11O), as well as calcareous algae e.g. *Triploporella* sp. (Fig. 11N),
278 *Heteroporella? graeca* CONRAD, PAVLOPOULUS, PEYBERNES AND RADOIČIĆ (Fig. 11P)
279 are also abundant. The pelagites filling the neptunian dykes show a sharply different assemblage
280 that almost exclusively consists of planktonic foraminifers such as *Rotalipora* sp. (Fig. 11Q),
281 *Rotalipora cf. appenninica gandolfi* LUTERBACHER AND PREMOLI-SILVA (Fig. 11R),
282 *Rotalipora cf. cushmani* (MORROW) cf. *greenhornensis* (MORROW) (Fig. 11S), *Rotalipora cf.*
283 *reicheli* MORNOD (Fig. 11T), *Whiteinella cf. praehelvetica* (TRUJILLO) (Fig. 11U), *Whiteinella*

284 sp. (Fig. 11V), *Whiteinella* cf. *paradubia* (SIGAL) (Fig. 11W, X), *Whiteinella baltica* DOUGLAS
285 AND RANKIN (Fig. 11Y, Z), though resedimented benthic foraminifers (e.g. orbitolinids) may
286 occasionally occur.

287 **INSERT FIG. 11**

288

289 **6. Discussion**

290 The new sedimentological and biostratigraphic data from the Pizzo Muletta section allow to re-
291 define the age of a migration pulse of African dinosaurs towards Europe and to trace a detailed
292 evolution of a sector of the possible land bridge represented by the Panormide Platform.

293

294 *6.1. Time constraints for the dinosaur bone and biostratigraphy remarks*

295 The assemblages coming from Section M and Section PM show good matches with the associations
296 observed in coeval inner platform/backreef environments belonging to the Apennine Platform
297 (Chiocchini et al., 2008; Mancinelli and Chiocchini, 2006; Ruberti et al., 2013). These similarities
298 do not surprise since as this palaeogeographic domain was neighbouring and connected to the
299 Panormide Platform during the Early Cretaceous (Zarcone et al., 2010). Although the assemblages
300 of sections M and PM share several biota, some remarkable differences have been observed. The
301 rare record of rudists characterizing Section M is consistent with an inner-platform environment as
302 suggested by the observed alternation of peritidal facies. In contrast, rudist fragments and loose
303 specimens are common in Section PM, though no *in situ* bouquet have been observed as it would be
304 expected in a reef/fore-reef environment (Cestari and Sartorio, 1995). Furthermore, representatives
305 of the genus *Mesorbitolina* only occur in mud-free skeletal streams coming from the margin area
306 and emplaced by storm/high-energy events. These data suggest more open marine conditions for
307 Section PM when compared with section M and an open lagoon environment can be thus inferred in
308 agreement with the sedimentological interpretation. A gradual transition toward more open marine
309 conditions is also supported by the increase in calcareous algae and the occurrence of biota typical

310 of reefal-platform margin environment (i.e. *Triploporella* sp.) observed in the uppermost strata of
311 Section PM. The biostratigraphic ranges of some of the main characteristic biota found in Section
312 M and Section PM is rather uncertain (e.g. *Cuneolina sliteri*, *Mesorbitolina birmanica*, *Nezzazata*
313 *isabellae*, *Morelletpora turgida*; see Chiocchini et al., 2008; Di Lucia, 2009; Mancinelli and
314 Chiocchini, 2006; Schlagintweit and Wilmsen 2014; Schmitt et al., 2019; Velić, 2007). However,
315 the different interpretations remain constrained to the late Aptian–early Albian interval. In
316 particular, the occurrence of *Simplorbitolina aquitanica* would suggest a late Aptian–?earliest
317 Albian age for Section M and Section PM (Chiocchini et al., 2008; Fourcade, 1978; Moullade and
318 Peybernès, 1979; Schroeder and Poignant, 1964). As a consequence, the dinosaur bone occurring in
319 the uppermost strata of section M and considered until now Cenomanian in age (Garilli et al.,
320 2009), should be dated back to the late Aptian–(?early Albian). Furthermore, the inferred time
321 interval is consistent with the age of the titanosaurid bones recently discovered in the northern
322 sector of the Apennine platform (i.e. Latium-Abruzzi platform) (Dal Sasso et al., 2016) and strongly
323 supports the Adria bridge dispersal model.

324 Upsection the microfossil assemblage of the Section MZ shows the presence of *Spiroloculina* sp.,
325 *Cuneolina* sp. and, undeterminable orbitolinids. This assemblage suggests a late Albian–middle
326 Cenomanian age, in agreement with age and the macrofossil assemblage determined by Garilli et al.
327 (2009). The neptunian dykes crosscutting section MZ are filled up with calcilutites containing
328 planktonic foraminifers pertaining to the *Dicarinella algeriana* sub-biozone (Premoli Silva and
329 Verga, 2004) that suggests a latest Cenomanian age. This constrains the emersion and karstification
330 of the platform to the late Cenomanian.

331

332 6.2. Black pebbles conglomerates

333 Black pebbles are widespread in the Mesozoic carbonate platforms from southern Italy (Bosellini,
334 2002; Eberli et al., 1993; Spalluto, 2012) and Sicily (Todaro et al., 2017), though they normally
335 occur as thin and sporadic lag levels at the base of peritidal shallowing-upward cycles. These

336 intraclasts likewise occur in the peritidal limestones of section M, but they represent the main
337 constituent of the lagoonal facies of the section PM. Organic plant material acts as the “blackening”
338 agent (Lang and Tucci, 1997; Platt, 1992; Strasser and Davaud, 1983) according to three models
339 (Strasser, 1984) implying (i) anoxic conditions, (ii) uplift stages of the carbonate platform
340 (Leinfelder, 1987; Strasser, 1984) and (iii) the absorption of organic matter at the surface of
341 calcretes during pedogenetic micritization processes (Sāsāran, 2006; Strasser, 1984). A partially
342 different hypothesis ascribes the blackening rather to “instantaneous” forest fire heating (e.g.
343 Florida Keys, Shinn and Lidz, 1988) as supported by heating experiments demonstrating that
344 limestone pebbles can be blackened at temperatures between 400 °C and 500 °C (Flügel, 2004).
345 Besides the actual blackening agent, all these hypotheses call into question the contemporaneous
346 presence of nearby subaerial environments. Black-pebble occurrences are thus diagnostic indicators
347 for partial to complete subaerial exposure of subsequently reworked carbonate deposits (Strasser
348 and Davaud, 1983) as well as relative fluctuations of the sea level (Mircescu et al., 2013). The
349 widespread occurrence of black pebbles in section PM is thus a clear evidence of a relatively long-
350 lasting emergent land adjacent to this sector of the PCP during Aptian–Albian times.

351

352 *6.3 Cretaceous tectono-sedimentary evolution of the Panormide Carbonate Platform and the fate of* 353 *the connection between Africa and Adria*

354 Looking in detail to the Pizzo Muletta section, the whole stacking pattern of the facies associations
355 indicate a progressive shifting from peritidal (i.e. Section M), to lagoonal environments (i.e. section
356 PM) during late Aptian–early Albian (Fig. 12A) and, successively, to a high-energy, rudist-
357 dominated open shelf (i.e. section MZ) during late Albian–middle Cenomanian (Fig. 12B). The
358 Cretaceous global sea-level curve displays a long-term period of stable and high sea-level during
359 Aptian and early Albian followed by a progressive rise during late Albian (Haq, 2014). This eustatic
360 rise could justify the observed facies shifting at Pizzo Muletta toward more open environments even

361 though the global trend is punctuated by third-order fluctuations that could have caused the
362 periodical emersion of the platform.

363 Both the theropod bone found in the upper part of section M and the widespread occurrence of
364 black pebbles in the upper Aptian–?lower Albian strata point to the presence of a landmass area
365 close to Pizzo Muletta (Fig. 12A). This nearby landmass could have provided nesting sites and
366 fresh-water to the dinosaur community inhabiting the carbonate platform (Nicosia et al., 2007).

367 Although, the origin of the vugs occurring in both Section M and Section PM remains doubtful, the
368 neptunian dykes on top of the Pizzo Muletta suggest that the platform was affected by subaerial
369 dissolution phenomena before being filled by pelagic wackestones. This intense dissolution
370 accounts for an emersion phase (Fig. 12C) experienced by this sector of the PCP during
371 Cenomanian times and could have allowed a further pulse of dinosaur migration from Gondwana
372 toward Adria (Citton et al., 2015). The pelagic infill of neptunian dykes accounts for a sudden
373 drowning of this sector of the platform during latest Cenomanian times (Fig. 12D). This event
374 matches the reactivation of transtensional crustal shears during latest Cenomanian–earliest
375 Turonian as documented by the tectonically-driven submarine volcanism in the adjacent WSCE.
376 Since then, the progressive fragmentation of the platform is documented by the huge input of
377 calciclastic sediments into the escarpments (Randazzo et al., 2020a and b).

378 **INSERT FIG.12**

379

380 Although the drowning and step-back of the Cretaceous PCP and other coeval carbonate platforms
381 may be related to the increase of ocean-crust formation (up to 50%- 70% between 120 and 80 Ma;
382 Larson, 1991) and to the subsequent rise of sea level (Mullins et al., 1991), there are several lines of
383 evidence that tie the dismembering of most of the carbonate platforms pertaining to the southern
384 sector of Adria to the Late Cretaceous extensional regime induced by the opening of the Sirte Basin
385 (Capitanio et al., 2009; Frizon de Lamotte et al., 2011). Tectonically induced variation in the
386 sedimentary basins are documented both in deep (e.g. north Africa: Grasso et al., 1999; Jongsma et

387 al., 1985; Sicily: Di Stefano et al., 1996; Randazzo et al., 2020a) and shallow-water settings (e.g.
388 Apulia: D'Argenio and Mindszenty, 1995; Graziano, 2000; Masse and Borgomano, 1987. Apennine
389 Platform: Carannante et al., 2009; Capotorti et al., 1997; Tavani et al., 2013; Vitale et al., 2018).

390 As a consequence, the platforms of the southern sector of Adria area were progressively
391 dismembered in isolated smaller carbonate platforms, by the drowning of discrete sectors, as in the
392 case of the Pizzo Muletta succession. This event was in turn responsible of the break-down of the
393 continental connection linking Gondwana to Adria (Fig. 13C).

394 **INSERT FIG.13**

395

396 **7. Conclusions**

397 New sedimentological and biostratigraphic data from the Pizzo Muletta section allow to constrain
398 facies and stratigraphic position of the dinosaur bone found in the Panormide Carbonate Platform as
399 well as to depict the complex dynamic of the platform across the Aptian–Cenomanian interval. On
400 the base of the facies associations three different zones can be recognized along the succession: i)
401 the lower part (c.a. 80 m thick) consists of peritidal facies, stacked in more or less regular peritidal
402 cycles. This part contains the theropod bone at about 20 m from the top and this part of the
403 succession can be assigned to the upper Aptian–?lower Albian. Upward, the second part consists of
404 a thick interval (about 40 m) of lagoonal packstone and grainstone dominated by black pebbles and
405 orbitolinids. The microfossil assemblage allows to assign also this part of the succession to the
406 upper Aptian–?lower Albian. The presence of a large amount of black pebbles in this part of the
407 succession suggests the presence of a persistent exposed land adjacent to the nearshore area. The
408 third portion of the section (about 20 m) consists of rudist floatstone and rudstone with abundant
409 and a well-diversified assemblage of rudist, Chondrodonta, and gastropods. This tract of the studied
410 succession can be assigned to the upper Albian–lower/middle Cenomanian. In the whole, the facies
411 sequence shows a trend from peritidal environments to high energy, rudist dominated, shelf margin
412 environments. Dissolution cavities filled up by pelagic wackestone on top of this succession

413 account for an emersion and subsequent drowning of this sector of the carbonate platform during
414 upper Cenomanian. The Cenomanian uplift could be tuned to the pulse of dinosaur migration from
415 Africa. The contemporary submarine basalt emissions in the adjacent slope namely Western Sicily
416 Cretaceous Escarpment, suggest the activity of deep-rooted crustal shears dissecting the PCP and,
417 more in general, the carbonate platforms of the central Mediterranean area.
418 Besides retro-dating the presence of dinosaurs in the Panormide Platform to the late Aptian–early
419 Albian interval, the complex dynamics observed throughout the Pizzo Muletta section documents
420 the irregularity of the connections between Africa and Adria through this land bridge.

421

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713

714 **FIGURE CAPTIONS**

715

716 **Fig. 1.** (A) Simplified structural map of Sicily and the Southern Apennines. (B) Present-day
717 distribution in northern Sicily (see black rectangle in A) of 1) Cretaceous and 2) Triassic-Jurassic

718 PCP carbonates, 3) WSCE slope carbonates, 4) slope to deep-water carbonates and cherts pertaining
719 to the Imerese Basin (IB).

720

721 **Fig. 2.** (A) Google-Earth view of the Pizzo Muletta area and locations of sections M, PM and MZ.
722 (B) Panoramic view of section M, sampling points and Grotta Lunga location. (C) The abandoned
723 quarry on the southern side of Pizzo Muletta and the sampling points of section PM.

724

725 **Fig. 3.** Biozonal schemes for the Aptian–Cenomanian interval (pelagic basins: Premoli Silva and
726 Verga, 2004; periadriatic carbonate platforms: Chiocchini et al., 2008; Velić, 2007).

727

728 **Fig. 4.** (A) The entrance to Grotta Lunga and (B) the theropod bone occurring inside the cave.

729

730 **Fig. 5.** Stratigraphic logs of the Pizzo Muletta succession showing sample and facies distribution
731 and the stratigraphic position of the dinosaur bone. Notice the different scale between section M
732 and sections PM and MZ.

733

734 **Fig. 6.** Main facies types of section M (A-D), section PM (E-G) and section MZ (H, I). (A) Sharp
735 stratal surface filled up by Pleistocene calcarenites. (B) Bindstone with fenestral fabric overlying
736 black pebbles conglomerate. (C) Gastropod floatstone to rudstone. (D) Black-pebble level
737 occurring in section M. (E) Decimetre-thick beds of black-pebble rudstones and floatstones. F)
738 Close-up of black-pebble floatstone (down) and rudstone (up). G) Rudists scattered in a black-
739 pebble rudstone. H) Rudist rudstone to floatstone crosscut by a neptunian dyke filled up by pelagic
740 calcilutites. I) Well-preserved caprinids in section MZ.

741

742 **Fig. 7.** Main microfacies types of section M (A, B), section PM (C-E) and section MZ (F-H). (A)
743 Skeletal packstone with miliolids and cuneolinids, sample M11, (B) Wackestone with gastropods,

744 benthic foraminifers, and ostracods, sample MG3, (C) Skeletal packstone with radiolitid fragments
745 and orbitolinids. Microboring (green arrows) is evident in places, sample PM5, (D) Well-sorted
746 skeletal grainstone with rudist fragments and orbitolinids. The elongated grains show a rightward
747 imbrication, sample PM12, (E) Grainstone showing micritized skeletal grains of calcareous algae,
748 sample PM15, (F) Grainstone with orbitolinids and micritized radiolitid fragments. Microboring
749 (green arrows) is evident in places, sample MZ3, (G) Grainstone-packstone with rudist fragments
750 and benthic foraminifers, sample MZ4, (H) Wackestone rich in planktonic foraminifers occurring as
751 filling of neptunian dykes, sample MZ1. Scale bar is 500 μm .

752

753 **Fig. 8.** Dissolution overprint on the Pizzo Muletta succession. (A-C) Dissolution cavities (vugs)
754 pervading section M (A) and section PM (B, C). (D, E) Section MZ hostrock consisting of rudist
755 floatstones and rudstones crosscut by neptunian dykes showing internal laminations that suggest the
756 progressive fill of open cavities. (F) Original and (G) interpreted images of an irregular dissolution
757 cavity filled up by pelagic limestones.

758

759 **Fig. 9.** Composite column of the Pizzo Muletta succession paralleled to the distribution of the main
760 identified taxa.

761

762 **Fig. 10.** Section M (A-K) and section PM assemblages (L-Y). (A) *Spiroloculina* cf. *cretacea*
763 REUSS, sample M2, (B) Haplophragmoididae indet (?*Debarina hahounerensis* Fourcade et al.),
764 sample M3, (C, D) *Nezzazata isabellae* ARNAUD-VANNEAU AND SLITER, samples M3, M11,
765 (E) *Quinqueloculina histri* NEAGU, sample M11, (F) *Glomospira* sp., sample M13, (G-I)
766 *Cuneolina* aff. *sliteri* ARNAUD-VANNEAU AND PREMOLI SILVA, samples M7, M11, M16, (J)
767 ?*Paracoskinolina* sp., sample MG1, (K) Orbitolinidae indet., sample MG3. (L) *Nezzazata isabellae*,
768 sample PM10, (M) *Akcaya minuta* (HOFKER), sample PM6, (N) *Pseudonummoloculina* sp.,
769 sample PM11, (O, P) *Simplorbitolina aquitanica* (SCHROEDER), samples PM11, PM5, (Q)

770 Orbitolinidae indet., sample PM12, (R-S) *Mesorbitolina birmanica* (SAHNI), sample PM12, (T)
771 ?*Paracoskinolina* sp., sample PM12, (U, V) *Orbitolinopsis* gr. *cuvillieri* MOULLADE, sample
772 PM5, (W) *Triploporella* sp., sample PM15, (X) *Morelletpora turgida* (RADOIČIĆ), sample PM13,
773 (Y) *Cylindroporella? lyrata* MASSE AND LUPERTO SINNI, sample PM6. Scale bar is 200 µm.
774

775 **Fig. 11.** Section MZ assemblage (A-P hostrock, Q-Z neptunian dyke). (A, B) *Spiroloculina* sp.,
776 sample MZ7, (C) *Pseudonummoloculina?* sp., sample MZ7, (D) *Vercorsella* sp., sample MZ7, (E)
777 Haplophragmoididae indet (?*Debarina hahounerensis* Fourcade et al.), sample MZ7, (F) *Nezzazata*
778 *isabellae*, sample MZ4, (G-I) *Cuneolina* sp., samples MZ4, PM3, MZ3, (J-M) Orbitolinidae indet.,
779 samples MZ4, MZ2, (N) *Triploporella* sp., samples MZ7, (O) *Coscinophragma?* sp., sample MZ3,
780 (P) *Heteroporella? grecae* CONRAD, PAVLOPOULUS, PEYBERNES AND RADOIČIĆ, sample
781 MZ3. (Q) *Rotalipora* sp., sample MZ5, (R) *Rotalipora* cf. *gandolfi* LUTERBACHER AND
782 PREMOLI-SILVA, sample PM1, (S) *Rotalipora* cf. *cushmani* (MORROW) cf. *greenhornensis*
783 (MORROW), sample MZ1, (T) *Rotalipora* cf. *reicheli* MORNOD, sample MZ1, (U) *Whiteinella* cf.
784 *praehelvetica* (TRUJILLO), sample PM2, (V) *Whiteinella* sp., sample PM2, (W, X) *Whiteinella* cf.
785 *paradubia* (SIGAL), samples PM3F, MZ1, (Y, Z) *Whiteinella baltica* DOUGLAS AND RANKIN
786 sample MZ1. Scale bar is 200 µm.

787
788 **Fig. 12.** Pizzo Muletta tectono-sedimentary evolution. (A) Environmental interpretation of section
789 M and PM during Aptian–Albian. (B) Environmental interpretation of section MZ during early–
790 mid Cenomanian. (C) Subaerial exposure and genesis of dissolution cavities in response to the
791 tectonic uplift during late(?) Cenomanian. (D) Tectonic drowning and filling of the dissolution
792 cavities with pelagic sediments during latest Cenomanian. Not to scale.

793
794 **Fig. 13.** (A) Palaeogeographic map of the Mediterranean area during Cretaceous times showing: i)
795 the continental connection between Africa and Adria (black square), ii) the subduction zones in the

796 Ionian and Alpine Tethys, and iii) the extensional stress field in the Sirte Basin (modified after
797 Frizon de Lamotte et al., 2011). Not to scale. (B, C) Evolution of the continental connection
798 between Africa and Adria during Early (B) and Late Cretaceous (C). Not to scale. Red dashed lines
799 are deep-rooted extensional faults. Not to scale. APCP=Apennine Carbonate Platform; AP=Apulian
800 Platform; CCP=Constantine Carbonate Platform; IB=Imerese Basin; LB=Lagonegro Basin;
801 PCP=Panormide Carbonate Platform; TB=Trapanese Basin; UMB=Umbria-Marche Basin;
802 WSCE=Western Sicily Cretaceous Escarpment.
803