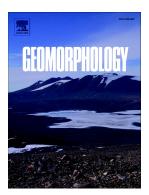
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# The use of historical cartography and ALS technology to map the geomorphological changes of volcanic areas: a case study from Gran Cono of Somma-Vesuvius volcano

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#### ABSTRACT

The eruptive activity of a volcano modifiers is surface topography through morphological changes generated by the deposition of errittee volcanic material and resulting gravity-driven processes, which can form accumulation or material in addition to the most common erosional phenomena. Mapping and quantifying such morphological changes allow to derive new data useful to better describe and understand the eruptive history of the volcano itself. Nowadays, one of the mostly used method to identify such morphological changes consists of comparing Digital Elevation Models (DEM) of the volcanic area before and after an eruptive event. If the eruptive event is referred to periods prior to 1980's, the only method to reproduce DEMs consists of elaborating the historical cartography that is often available only in paper format. In this work we aim to prove the reliability of this approach, presenting a study on the morphological changes (from 1876 to 1944) of the summit caldera of the Somma – Vesuvio volcano (Italy). For the first time, we compare DEMs derived from historical maps (1876, 1906 and 1929) and a DEM dated 2012 obtained by remote

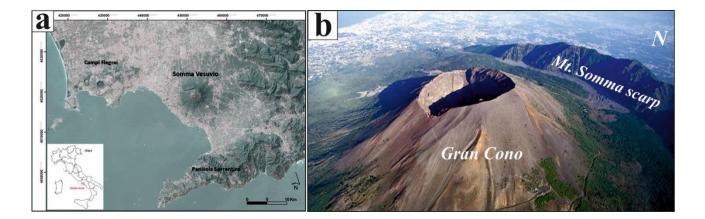
sensing. The four models of the caldera, digitally reproduced at the same spatial resolution, are morphologically investigated through specific maps derived from the DEMs and a set of height profiles. In addition, further morphometric analyses and accurate quantifications in volume and surface are presented and discussed for a portion of the Somma-Vesuvio summit caldera, represented by the Gran Cono edifice. Considering the different typology of the source data used in this study, it is also provided a discussion on the respective accuracies that, especially for the historical maps, represent a crucial point for obtaining DEMs able to reproduce topographies more realistic as possible. For this reason, despite data source were processed following rigours criteria, the calculations of volume, surface and distance related to the morphological changes of the volcano are associated to an accurate quantification of the error. Following this, the main results obtained in this study are: i) the identification of several pa.<sup>+</sup> volcanic deposits and the estimation of the related thicknesses, both in good agreement win published literature; ii) the quantification of the morphological changes of the Gran Co. of om 1876 to 1944 resulting in a volume and surface growth of  $133 \times 10^6$  m<sup>3</sup> (±5%) and ~ 0.14 km<sup>2</sup>, respectively; iii) the identification of a possible migration path of the centroid of the Cron Cono crater along the SW-NE preferential direction during the investigated period.

**Keywords**: Digital Elevation Model, Historical Cartography, Airborne Laser Scanning technology, Somma - Vesuvio volcano, Morphological Changes Quantifications

#### 1. INTRODUCTION

In the last 40 years, the topographies of an active volcano have been digitally reproduced with spatial resolutions and accuracies ranging from several meters to some tens of centimeters by using traditional methodologies based on digitalization of contour lines and elevation points from raster cartography (Favalli et al., 1999, Tarquini et al., 2007) and more innovative techniques from remote sensing (Wu, 1979; Ferretti et al, 1999; Garvin et al., 2018). In the last two decades, specific remote sensing techniques such as airborne laser scanning (ALS), aerial stereo photogrammetry combined with SfM (Structure from Motion) elaboration, and satellite ster op, otogrammetry, result to be the most diffused to create Digital Elevation Model (DEM) in vo'can c areas (Pesci et al., 2007; James and Robson, 2012; Jones et al., 2015; Bisson et al.,  $201\epsilon$ ; M iller et al., 2017; Beyer et al., 2018; De Beni et al., 2019). Such models, properly analyzed and compared, allowed i) to study the morphology of the same volcano in different time: (Neri et al., 2008), ii) to produce a series of data useful to better document and understand the vuptive history of the volcano itself (Marsella et al., 2012; Dai and Howat 2017; Whelley et al., 2017, Ganci et al., 2018) and iii) to simulate volcanic phenomena and their impact on the centrory producing maps useful to mitigate the hazard and risk volcanic (Felpeto et al., 2001; Huggel et al., 2008; Cappello et al., 2011; Salvatici et al. 2016; Annis et al, 2020). The morphology of a volcano before the 1980's cannot be obviously reproduced with the previously mentioned remote sensing techniques and hereupon the available historical cartography, with all its limits (Balletti, 2000), becomes the only reference data. The main volcanoes in Italy, Mt. Etna in Sicily and Somma-Vesuvio (SV) and Campi Flegrei in Campania, are characterized by a large amount of historical drawings and illustrations that often show an approximate representation of the territory during a particular eruptive event. Most of the oldest representations related to the three volcanoes date back to the 16<sup>th</sup> and 17<sup>th</sup> centuries (Naddeo, 2004; Branca and Abate, 2017). As regards SV, the first topographic map reproducing the entire area of the volcano with a reference of geographic coordinates dates back to the second half of 19<sup>th</sup> century, and was produced by the "Allievi of the Istituto Topografico Militare" during the 1875-1876

(Palmentieri, 2016), whereas for smaller areas some cartographies were already produced during the second half of 18<sup>th</sup> century (Ronza, 2017). During the 20<sup>th</sup> century, the cartography of SV area, as well as for the entire national territory, started to be produced by IGM (Istituto Geografico Militare), that became the main reference for the cartography in Italy. Up to now, IGM has reproduced the entire Italian territory at different scales (1:25,000, 1:50,000, 1:100,000, 1:250,000) through aerial photogrammetry. At the beginning, the cartography has been produced in the Roma 40 reference system, and subsequently, starting from the 1980's, firstly in the ED 50 and finally in the WGS84 (www.igmi.org) reference systems. Nowadays, the most detailed reproductions of a territory, or its portion, are obtained by elaborating ALS dat. or by coupling the SfM technique to the UAV (Unmanned Aerial Vehicle) photogrammetry. Such systems allow to reproduce at high spatial resolution and accuracy (few tens of cm) all 'he features (natural and anthropic) of the territory. In this work, which extends the prelimin, rv work of Bisson et al. (2020) by presenting new data and further analyses, we present a stu ly that compares the topographies of a volcanic area through time, and derived from very different sources (historical cartography and ALS acquisition). The aim of this work is firstly to demonstrate that the historical topographic maps, if processed and elaborated with rigours criteria, allow to derive reliable information for quantifying morphological changes of volcanic areas before the 1980's. In particular, this paper documents and analyses the morphological changes be, veen 1876 and 1944 of SV (Fig.1), one of the most hazardous volcanoes in the world, not only for the very high population density (up to 50,000 inhabitants/km<sup>2</sup> - ISTAT, 2011) of its surroundings, but also for the extreme variability of its eruptive style, capable of producing large-magnitude explosive eruptions (Cioni et al., 2008).



**Figure 1** - The study area: a) A Landsat 7-ETM image of the Camy anian plain including the Somma - Vesuvio (SV) volcano. The coordinates are expressed in the UTM WGS5-: 33N reference coordinate system; b) aerial image of SV volcano with the city of Naples in the backgroup 1.

As regards the 3D digital reproduction of past topographies at  $S^{\vee}$ , in literature we found only two works. The first work (Favalli and Pareschi, 2004) reconstructs at different time frames (1900, 1906, 1920 and 1988) the DEM of a portion of the SV cosresponding to the current "Gran Cono" edifice with the aim to test the "DEST" algorith. a variant of Delaunay triangulation able to improve the TIN construction and consequently, the morphology reproduction. The second work (Ventura and Vilardo, 2008) reconstructs u.º 3D model of the 1944 lava flow from ALS data with a spatial resolution of 0.33 m. This very high resolution allowed a very detailed morphological representation of the flow itself us ful to infer some emplacement mechanisms as well. Differently from these two works, we propose a multi-temporal study to document the topography changes at SV over the past century. Ve reconstruct the topographies by using historical maps and elevation data acquired by remote sensing techniques. In detail, the DEMs reproduce the entire caldera zone that will be analysed through specific maps and a set of height profiles, while the estimations in area and volume of the morphological changes are referred to a specific zone, common to all DEMs, identified with the "Gran Cono" edifice. The caldera zone and Gran Cono edifice investigated in this work are defined according to Tadini et al. (2017). After a brief description of the geological and volcanological setting of the study area, we will present the source data and the derived DEMs associated to a quantification of the plane-altimetry errors. Then, the results obtained from the morphological analyses will be described and discussed pointing out the main critical issues and the

possible implications with the geo-volcanological aspects. Finally, we will conclude summarizing the main remarks.

#### 2. GEO-VOLCANOLOGICAL SETTING

Somma-Vesuvio (SV) is a relatively young composite volcano of moderate altitude (~1200 m a.s.l.) located in the Campanian Plain (Central Italy) between the Campi Flegrei volcano and the Penisola Sorrentina (Fig. 1a). Its oldest products postdate the Campanian Unimbrite deposit of the nearby Campi Flegrei volcano (39 ka BP; Cioni et al., 2008), and are rolated to the first eruptive phase of the volcano. During this latter (~39-22 ka BP; Sbrana et al, 2020), effusive activity resulted in the accumulation of thick piles of lava flows and scoria deposite, which built up the ~1900 m high Mt. Somma (Cioni et al., 1999). Such edifice has been partially dismantled by four large-magnitude eruptions, starting from the Pomici di Base 7 n. ian eruption of ~22 ka BP (Bertagnini et al., 1998; Santacroce et al., 2008) up to the world-famou. AD 79 Plinian eruption (Sigurdsson et al., 1985). This intense explosive activity punct at at the last 20 ka, it produced thick tephra fallout and pyroclastic density currents deposite and developed the characteristic summit caldera of the SV volcano (Cioni et al., 1999). At present, the Mount Somma edifice remnants are visible along a morphological scarp (facing to the south the SV summit caldera - Fig. 1b) located in the north side of the edifice and along doeply incised valley on the northern part of the volcano (Santacroce and Sbrana, 2003; Sbrana et al. 2020). The Gran Cono (or Vesuvio sensu stricto - Fig. 1b) edifice started to grow within the summit caldera partially after the AD 79 eruption, and more substantially after the AD 472 sub-Plinian eruption (Cioni et al., 1999). Such growth occurred in consequence of effusive to mild explosive activity during period of open conduit conditions, which have been at least 4 after the AD 79 eruption (Cioni et al., 2008). The latter period with open conduit conditions started with the AD 1631 sub-Plinian eruption (Rosi et al., 1993) and lasted from 1631 to 1944. Since the last VEI 3 eruption of AD 1944 (Cole and Scarpati, 2010), SV is considered dormant in a

closed conduit condition (Scandone et al., 2008). In general, during an open conduit period at SV, eruptions are mostly characterized by non-stationary strombolian activity where magma is in a shallow magma chamber (Scandone et al., 2008). When a new batch of magma arrives into the shallow conduit filling it, magma could either overflow in gentle lava effusion or, in case of higher inputs, generate larger magnitude, violent strombolian eruptions (e.g. the 1906 eruption; Bertagnini et al., 1991; Scandone et al., 2008). When these latter eruptions occur, the shallower magmatic system tends to be evacuated, and the successive new magma batches result in a reappraisal of the non-stationary strombolian activity (Scandone et al., 2008). During the period 1631-1944, the numerous historical accounts and the abundance of studied deports allowed to reconstruct precisely the volcanic history of this last period of activity (Scandone et al., 2008): Table 1 summarizes all the eruptive events occurred in the years 1876-1944, correspending to the period investigated in this work. For each eruptive event, the table reports the date explanation of the Volcanic Explosive Index (VEI).

	Table 1- Summary o	of SV volcanic	e cupi ons in	the period 1876-1944
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Beginning (dd/mm/yyyy)	Enc (dd/m <sup>·</sup> u <sup>/</sup> yyy)	Length (days)	Type <sup>a</sup>	Deposit volume (m <sup>3</sup> )	VEI
16/12/1881	$\frac{(44)}{31/c^{1/1}}$	( <b>au</b> <sub>3</sub> s) 776	S+LF	n.a.	1-2
02/05/1885	01/07/1386	425	S+LF	n.a.	1
16/04/1887	19/04/1887	3	S+LF	n.a.	1
01/05/1889	30/09/1889	152	S+LF	n.a.	1
07/06/1891	03/06/1894	1092	LF	3.6×10 <sup>7</sup> (lava)	2
03/07/1895	07/09/1899	1527	LF	$1 \times 10^8$ (lava)	2
27/08/1903	30/09/1904	400	S+LF	$5 \times 10^5$ (lava)	2
03/02/1906	03/04/1906	59	S+LF	$1.8 \times 10^{6}$ (lava)	1
04/04/1906	22/04/1906	18	VS+LF	$2 \times 10^7$ (lava) 7.1×10 <sup>7</sup> (tephra)	3–4
27/11/1926	28/11/1926	1	S+LF	n.a.	1
01/08/1927	02/08/1927	1	S+LF	n.a.	1
11/08/1928	12/08/1928	1	S+LF	n.a.	1
03/06/1929	08/06/1929	5	S+LF	$1.2 \times 10^7$ (lava)	2–3

11/07/1930	30/07/1930	19	S+LF	n.a.	1
02/10/1930	09/11/1930	38	S+LF	n.a.	1
01/06/1933	19/11/1934	536	S+LF	n.a.	1-2
12/02/1935	31/03/1935	47	S+LF	n.a.	1
08/07/1935	21/08/1935	44	S+LF	n.a.	1
28/03/1936	24/09/1936	180	S+LF	n.a.	1
04/06/1937	07/07/1937	33	S+LF	n.a.	0
08/08/1939	09/08/1939	1	S+LF	n.a.	0
26/06/1940	31/07/1940	35	S+LF	n.a.	0
22/10/1941	15/12/1942	419	S+LF	n.a.	1–2
06/01/1944	23/02/1944	48	S+LF	n.a.	0
				$2 \times 10^7$ (lava)	
18/03/1944	04/04/1944	17	VS+LF	$6.6 \times 10^{7}$	3
				(tephra)	

<sup>a</sup>Type: S – Strombolian activity; LF – Lava Flow; VS – Violent Stron boli n eruption. The type comes from Scandone et al. (2008) with integrations from Arrighi et al. (2001) find Cioni et al. (2008).

# 3. CONSTRUCTION OF DEMS FROM HISTOR CAL MAPS AND ALS DATA

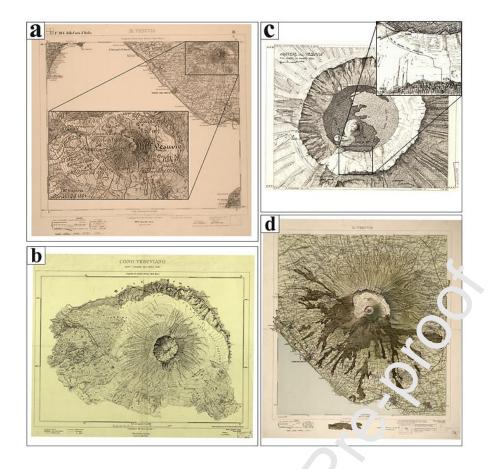
The data source and the methodologies used to reconstruct the topographies assumed by the SV summit caldera during the last century are presented in the two following sub-sections: i) Topography from historical maps; ii) Topography from Airborne Laser Scanning (ALS). In both cases, the ESRI GIS platform was plated as main working environment to process source data and reconstruct the topography modele, whereas specific software such as Tn-ShArc (Terranova platform), CartLab 2.0 and Traspunto (www.igmi.org) are used to uniform the cartographic reference system of all data, fundamental requirement for analysing and comparing geo-spatial data in GIS environment. In this study, all data are managed in the WGS84 UTM 33N cartographic reference system.

#### 3.1 Topography from historical Maps

The 3D SV topographies of 19<sup>th</sup> and 20<sup>th</sup> century were reproduced starting from IGM historical maps that contain geographic references in latitude and longitude, generally placed at the boundary of the maps themselves. These references are fundamental to assign to the map, in raster format, the

spatial coordinates and subsequently extract in geo-referenced digital vector formats (points, polylines, polygons) the 3D topographic information necessary to create the DEM. The historical maps used to reproduce the topography of SV summit caldera are referred to the 1876, 1906 and 1929 years (Fig. 2a,b,d); moreover, a supplementary map dated 1920 and centred on the crater area (Fig. 2c) was used specifically to better constrain the crater rim outline evolution.

The four maps, selected among those available, show the topography of the volcano before and immediately after the main eruptive events occurred from 1876 to 1944 covering a total period of about 50 years. The 1876 map corresponds to sheet n° 184 of the Ca ta d'Italia produced by IGM at scale 1:50,000 with contours lines every 10 m. This map allows to reconstruct the topography of the SV summit caldera before two eruptive events occurred in 1891-94 and 1895-99, respectively (see Table 1). These events created two volcanic structures, "Colle Margherita" and "Colle Umberto", named "dome-shaped accumulations of the lava "lows" by Arrighi et al. (2001) and "exhogenous tholoids" by Santacroce and Sbrana (2003) respectively. The 1906 map, produced at scale 1:10,000, updates the topography of the volcano and use violent eruption of April 1906 (Table 1), considered the strongest event of the 20<sup>th</sup> cer ury because it heavily modified the morphology of the Gran Cono (Bertagnini et al., 1991). With the aim of obtaining a detailed reconstruction of the SV summit caldera after this trong eruptive event, the 1906 map was produced by reporting the contours lines every 5 m (in tead of 10 m). The 1920 map, instead, provides a detailed outline of the Gran Cono crater rim during a period of intense intra-crateric activity (Scandone et al., 2008) before the June 1929 eruption. Finally, the 1929 map, produced at scale 1:25,000, delineates the topography before the 1944 eruption (Cole and Scarpati, 2010) reporting not only the topography of the volcano and its crater, but also portions of surrounding areas after the June 1929 eruption.



**Figure 2** - IGM historical maps. a) 1876; b) 1906; ) 1920; d) 1929. In the inset of 2c it is possible to read the highest points of the 1920 crater rim. The maps are provided by IGM in agreement with the authorization  $n^{\circ}$  0013597 dated 24/09/2019.

All the four maps were acquired by scanning the original paper at a resolution of 300 dpi and the outputs were stored in 8-bit raster format (grey tones). The high resolution allowed to better reproduce the cartographic dements (contour lines and height points) indispensable for our purpose, whereas grey tones highlighted some morphological details (i.e. very steep slopes, gullies, edges and others) that, inserted in the interpolation procedure, improved the resulting DEM. The scanning procedure was made by using a plane scanner to limit the deformation errors as much as possible. Subsequently, the obtained raster maps were geo-referenced by using the ground control points method (De Leeuw et al., 1988). In particular, for each raster, 8 points were selected in order to satisfy the following requirements: i) obtaining a spatial distribution of the points in order to cover in homogenous mode the entire image; ii) selecting the points rigorously either in the geographic grid reported as the map boundary or, when not possible, in some details or features well

recognizable both in the historical maps and in the nowadays cartographies geo-coded in the WGS 84 UTM 33N reference cartographic system. The original latitude and longitude coordinates, expressed in degrees according to Roma 40 System and readable on the boundary of the historical maps, were transformed in 8 points geo-referenced in the WGS 84 UTM 33N metric system through the software Traspunto (www.igmi.org), that assures an error of positioning < 1 m (Travaglini, 2004). Subsequently the 8 points were used in the geo-referencing procedure based on the affine transformation and performed on the TN-ShArc platform (https://www.agros.it/softwareterranova). In detail, after the transformation, the procedure calculates the RMSE (Root Mean Square Error) and if this latter results < 2 pixel, the final generative map is automatically memorized according to the raster format selected by the user. In this study, the output maps were stored in TIFF format. Finally, by digitizing the elevation data in line and point vector formats (shapefile), for each historical map two 3D drta are were created and used to reconstruct the topography of the caldera zone related to the 1876, 1906 and 1929 years. By interpolating the datasets with the Delaunay algorithm (Sn. wchuk, 2002), three elevation vector models stored in ESRI TIN format were produced and the vonverted in ESRI GRID format with a spatial resolution of 5 m, a value coherent with the original scale of the source maps. Figure 3a-c shows the three morphologies of the past through the Shaded Relief map derived from the respective DEMs.

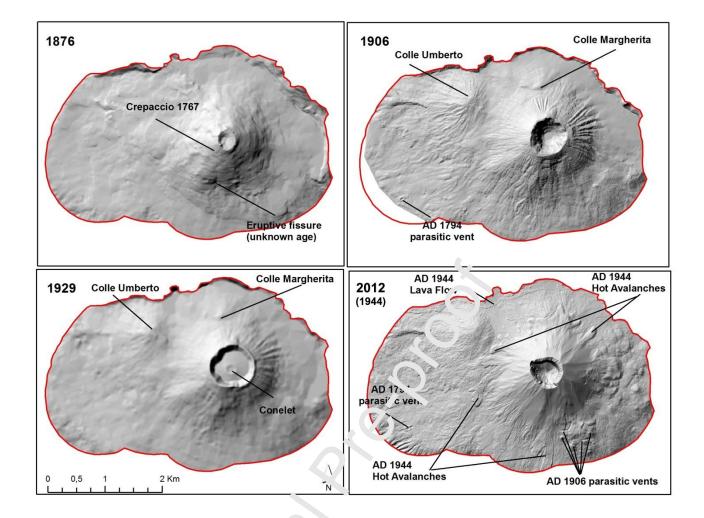


Figure 3: Shaded reliefs of SV summit caldera for the a) 1876, b) 1906 and c) 1929 years; d) shaded relief of the 1944 year obtained using ALS data (2012)

#### 3.2 Topography from ALS

For the present-day topography, ALS data acquired during 2012 year by the Province of Naples have been processed to obtain a very high resolution (1 m) elevation model. The ALS or LiDAR (Light Detection and Ranging) is a remote sensing technique able to acquire millions of a surface X, Y, Z points ("points cloud") by scanning the surface itself with an instrumental system (Infrared Sensor, IMU and GPS) mounted on an aircraft or helicopter. The source data covering the study area were elaborated and organized following the procedure described in Pizzimenti et al. (2016). The result is an accurate DEM that reproduces the actual morphology of SV caldera zone (Fig. 3d) with very high spatial resolution (cell size 1 m) and planimetry and altimetry accuracy equal to  $\pm 30$ 

cm and  $\pm 2$  cm, respectively (Pizzimenti et al., 2016). Taking into account that the volcano is quiescent from 1944 and that from this date the remarkable mass wasting/gravity-driven phenomena were recorded only in areas located outside the SV summit caldera (Bisson et al., 2017 and references therein), the DEM from ALS was considered representative of the SV caldera zone topography after the eruption of 1944. With this assumption, the morphological changes of minor entity generated by secondary erosion phenomena (e.g. wind, intensive precipitation, water runoff) are not here considered. Since these latter can be considered negligible with respect to morphological changes caused by the eruptive activity, we think that it is reasonable to include their contribution in the range of error estimated for each volume *a* to the this study.

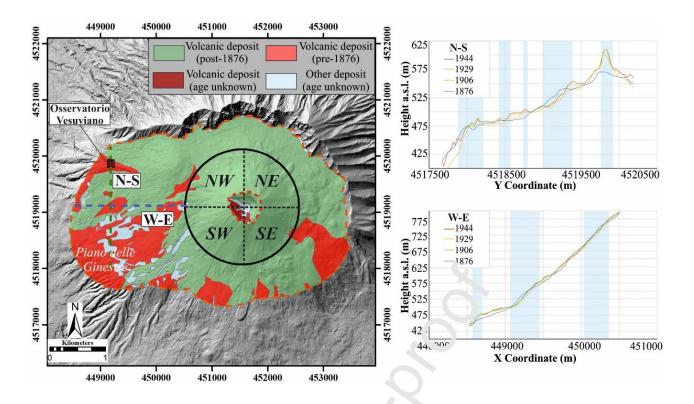
#### 3.3 Spatial accuracy of the 3D historical topographies

The accuracy of the data represents a crucial point for obtaining DEMs capable of reproducing topographies as realistic as possible and consequently to derive reliable quantitative data. For this reason, it is fundamental to provide an eximation of such accuracy by quantifying the errors of the source data and derived products arberefore, in order to calculate the error in x, y and z of the historical topographies, the 2012 ALS DEM has been used as the reference model because it reproduces the actual top only with very high resolution and accuracy (see previous section). The spatial error in x and y of the historical maps has been calculated by using the Vesuvius Observatory as a reference, because it is an historical building well recognizable in each map. In particular, we have identified the centroid of the edifice in each historical map and compared the respective coordinates (x and y) with those of the centroid of the same edifice reproduced in the 2012 ALS Model (x: 449185 m; y: 4519829 m). The differences in x and y (x/y) of the centroid were calculated by subtracting the respective historical coordinates from the ALS coordinates, and is equal to +48 m/+24 m (1929), -21 m/+30 m (1906) and +40 m/+50 m (1876).

To estimate the error in height of the historical topographies, we have constructed two profiles defined "calibration profiles" that analyse the elevation of each DEM in the same points. The profiles, one in N-S direction and the other in the W-E direction (Fig. 4), have been drawn to cross areas not affected by post 1876 morphological changes, according to literature data referring to the investigated time. In addition, to assure a correct analysis of the profiles, the geo-referenced historical maps were preventively co-registered to the ALS DEM by using the Vesuvius Observatory as tie point and the errors in x and y of each historical map have been also recalculated by using the RMSE reported in Table 2.

Historical Map	Number of° points	RMSE ju A	<b>AMSE in Y</b>
1876	10	8	± 23
1906	10	11.5	± 13
1929	10	± 19	± 16

The calibrations profiles (V/-E and N-S) were constructed by plotting the heights every 25 m, a value defined according to the tolerance error of each historical map. For the height error calculation, only the points falling in areas not affected by the post-1876 deposits have been considered because they are representative of a topography that should not have changed in the last century. Therefore, despite the calibration profiles W-E and N-S (Fig. 4) plot a total of 80 and 102 points, respectively, only 84 points (32 from W-E profile and 52 from N-S profile) are used for estimating the height error of each historical topography through the calculation of the RMSE. The values of RMSE result  $\pm 11$  m,  $\pm 4.5$  m and  $\pm 11.5$  m for the 1876, 1906 and 1929 elevation topography, respectively.



**Figure 4** – The traces (left) and relative calibration racfiles (right). The image shows the simplified geological units from Santacroce and Sbrana (2003) superimposed to the shaded relief of ALS DEM. The extent of SV summit caldera is represented by cange dashed line, whereas the 1944 SV crater rim is outlined with a red dashed line. The "NW", NF, "SW" and "SE" abbreviations identify the four Gran Cono quadrants defined according to the cardinal points. The coordinates are in the UTM WGS84 33N system. In the graphs, the portions of the proview for the calibration are highlighted in light blue.

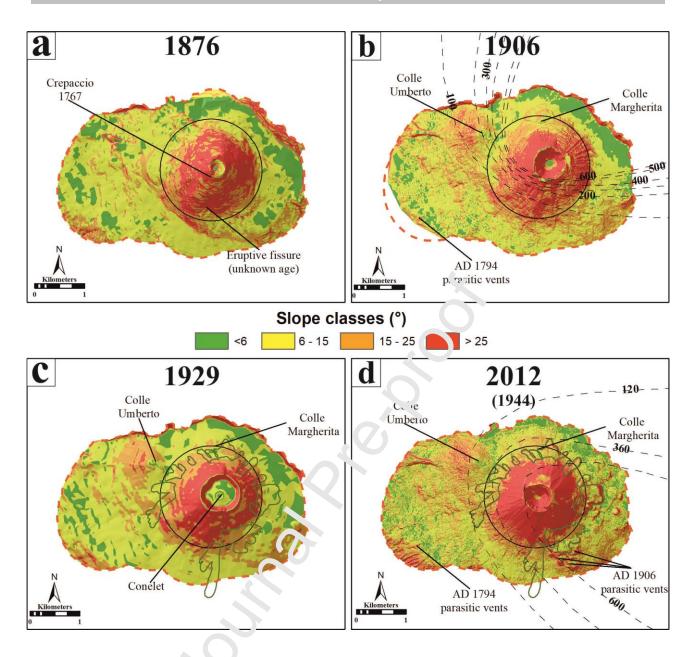
# 4. MORPHOLOGICAL CHANGES OF GRAN CONO FROM 1876 TO 1944: ANALYSES AND DISCUSSION

Since the four DEMs (Fig. 3) do not have the same extent, we have determined a common area to analyse the morphological changes during the last century. This latter corresponds to the Gran Cono area, outlined by a circle with a radius of 1 km (Fig. 4, 6) as defined in Tadini et al. (2017). The analyses applied to the elevation models and the obtained results are organized in three different subsections: i) identification of the most important volcanic morphologies (i.e. lava and pyroclastic deposits, avalanches, parasitic effusive vents related to past specific eruptions) through specific maps derived from the DEMs and a set of elevation profiles; ii) description of the most significant

changes occurred in the crater area; iii) estimations in surface and volume changes of the Gran Cono edifice from 1906-1944.

#### 4.1 Identification of volcanic morphologies

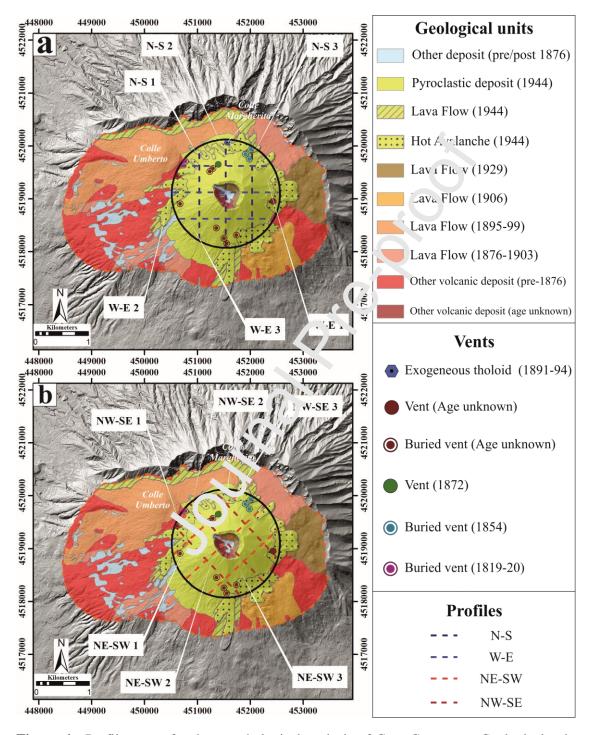
Despite the different accuracies of the source data, the shaded reliefs of the four DEMs (Fig. 3) show very clearly a series of morphological changes occurred in the SV caldera zone during the 1876-1944 period. The four shaded reliefs highlight: i) substantial enlargements and reductions of the Gran Cono crater rim through time; ii) the presence of some volcanic structures such as "Colle Margherita" and "Colle Umberto" since 1906 (Fig. 3b); iii) u.e formation of a "conelet" (Scandone et al., 1993) inside the crater of 1929; iv) a series of pa aside vents dated 1794 or of unknown age; v) some eruptive fractures such as "Crepaccio 1757" ell recognizable in the 1876 topography; vi) lava flows and "Hot Avalanche" deposits 1, 19'ed to AD 1944 eruption and well reproduced in the ALS topography (Fig. 3d). Also the slop, maps shown in Fig. 5 allow to appreciate the above mentioned morphological changes highlighting, in particular, the strong modifications of the Gran Cono crater. With the AD 1906 er. tion, the extent of the crater increases, and its internal portion is characterized by very steep and deep walls. Then, with the AD 1929 eruption, the crater rim further enlarges, and the morphology becomes gentler, since the internal walls shows moderate slopes (< 15°) and the bottom results prevalently flat. Finally, with the AD 1944 eruption, the morphology changes again, the extent of the crater diminishes assuming the nowadays dimension and the internal portion presents steeper walls that remind those of the 1906 crater.



**Figure 5** – Slope maps for the 1876 (a), 1906 (b), 1929 (c) and 1944 (d) DEMs. In b) the dashed black lines are loading isolines (in  $kg/m^{21}$ ) of the tephra fallout deposit from the AD 1906 eruption (Bisson et al., 2007). In c) and d) the solid green lines are the traces of the "Hot Avalanches" deposits of the AD 1944 eruption (modified from Santacroce and Sbrana 2003). In d) the dashed black lines represent the loads (in  $kg/m^{2}$ ) of the tephra fallout deposits from the AD 1944 eruption (Bisson et al., 2018). The solid black circle defines the Gran Cono area.

To better analyse and quantify the morphological changes, a set of elevation profiles were derived along four main directions (N-S, W-E, NW-SE, NE-SW). In detail, for each direction, we have defined 3 traces (Fig. 6a, b) distant 500 m apart from each other and we have plotted the height value every 25 m from the 4 DEMs (1876, 1906, 1929 and 1944). In addition to the traces, Fig. 6a and b display the distribution of parasitic effusive vents (Tadini et al., 2017) and a simplified

version of the geological map of Santacroce and Sbrana (2003). The elevation profiles obtained along the N-S and W-E direction are shown in Fig. 7, while those along the NW-SE and SW-NE directions are displayed in Fig. 8.



**Figure 6** - Profile traces for the morphological analysis of Gran Cono area. Geological units are modified after Santacroce and Sbrana (2003). Vent positions, SV summit caldera outline (dark orange dashed line) and Gran Cono area (solid black circle with a radius of 1 km) are reported after Tadini et al. (2017). In background, the shaded relief derived from 2012 ALS DEM. The coordinates are expressed in the UTM WGS84 33N cartographic system.

With the aim of facilitating the description of the morphological changes derived from the elevation profiles (Figs. 7 and 8), we prefer to introduce the four geographic quadrants (NW, NE, SW and SE) drawn in Fig. 4 and to distinguish two zones in the Gran Cono (the "crater" zone and "flanks" zone), whose morphological changes will be discussed separately. The "crater" zone is represented by the surface bounded by the 1944 rim (red dashed line in Fig. 4) used as reference, while the "flanks" zone corresponds to the surface between the 1944 crater rim and the Gran Cono bottom (black circle in Fig. 6).

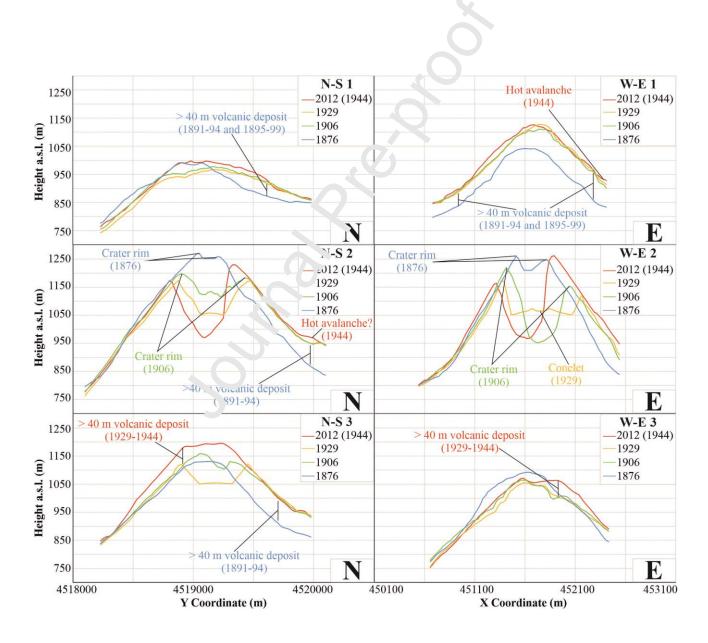


Figure 7 - N-S and W-E profiles referred to Figure 5

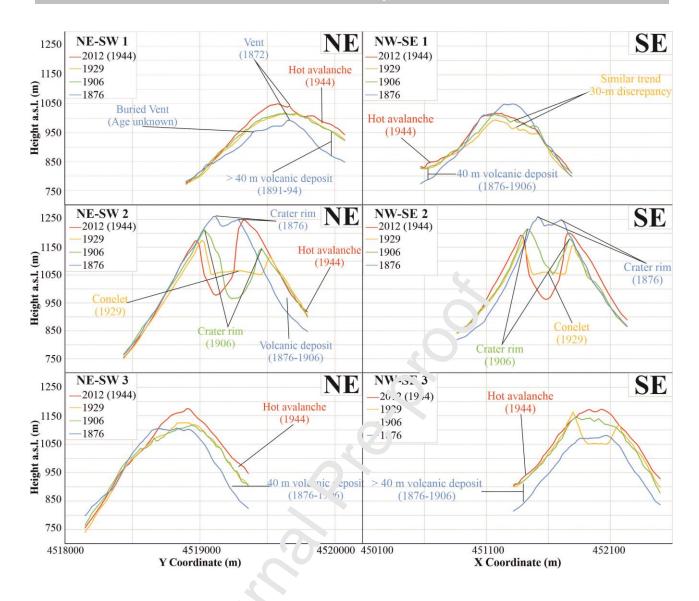


Figure 8- NE-SW and NW-SE promiss referred to Figure 5.

#### 4.1.1 The "crater" zone

To identify the morphological changes in the rim and bottom of the "crater" zone, we analyse the four reference profiles (W-E<sub>2</sub>, N-S<sub>2</sub> in Fig.6; NE-SW<sub>2</sub>, NW-SE<sub>2</sub> in Fig.8). The profiles show that the crater rim of 1876, with an elevation ranging from 1250 m to 1270 m a.s.l., is clearly the highest one. All profiles indicate a gradual and progressive decrease in elevation of the crater rim from 1876 to 1929. With the AD 1944 eruption, the N-S<sub>2</sub> and NW-SE<sub>2</sub> profiles show that the crater rim continues to decrease only in the western quadrants, whereas in the eastern quadrants it increases reaching the elevation of 1266 m a.s.l, very close to 1270 m a.s.l of the 1876 crater (W-E<sub>2</sub> and NE-

 $SW_2$  profiles). These observations suggest that the eruptions occurred from 1876 to 1929 progressively lowered the crater rim. Subsequently, the AD 1944 eruption continued to lower the crater rim only in the western quadrants, generating at the same time its remarkable growth in the eastern quadrant.

For the crater bottoms, between 1876 and 1944, the profiles show a general and progressive lowering in elevation resulting in 270 m ±11.5 m over a period of 70 years (from 1240 m to 970 m visible in the N-S<sub>2</sub> profile). In detail, with respect to 1876, the 1906 crater bottom decreases in elevation with different values according to the analysed profile. In 1 ct, this decrease ranges from a minimum of  $75 \pm 16$  m (see N-S<sub>2</sub> section) to a maximum of  $250 \pm 270 \pm 16$  m (W-E<sub>2</sub> and NE-SW<sub>2</sub>) sections). After the 1929 eruption, the crater bottom furth r decreases of 70  $\pm$  27 m in the N-S<sub>2</sub> section, whereas it increases of 70 - 100  $\pm$  27 m in the other sections (see NE-SW<sub>2</sub> and W-E<sub>2</sub>). With the AD 1944 eruption, all sections show that the trater bottom again decreases of 70 m reaching an elevation of 970 m a.s.l., a value that could have been already reached during one of the two previous eruptions (see 1906 and 1929 profiles). For completeness, we also report the value of the lowest point of the crater bottom derived from the ALS DEM and resulting equal to 954 m. a.s.l. This point is located around 25 means to the south of W-E<sub>2</sub> trace (see Fig. 6a) and its elevation results very close to the v due of  $953 \pm 5m$  measured by Imbò (1949) in June 1949. All these observations suggest that: 1 starting from 1876, the subsequent eruption of 1906 has strongly lowered the bottom of the crater indicating a strong eruptive activity; ii) the 1929 eruption has filled the bottom with products generated by a low magnitude explosive eruption that could be represented by an intra-caldera activity capable of forming a small conelet inside the caldera itself. These two hypotheses agree with Bertagnini et al. (1991) and Scandone et al. (2008), respectively.

4.1.2 The "flank " zone

Most of the profiles (Fig. 7 and 8) show that the older topographies are lower than the subsequent ones, indicating that the flanks of the volcano, from 1876 to 1944, progressively grow due to the accumulation of the erupted products along the flanks themselves. Only four profiles (N-S<sub>1</sub>, W-E<sub>3</sub> in Fig.7; NW-SE<sub>1</sub>, NE-SW<sub>3</sub> in Fig.8) show areas where the oldest topography (1876) is the highest one. Such areas have been individuated in some portions of the SSE and SSW flanks, particularly between the crater rim and the hot avalanche deposits (Fig. 6). A cording to the errors in elevation previously calculated (see section 3.2), in such areas the real elevation of the 1876 topography results 15 - 20 m lower than the one shown by the profiles. Particularly, during the AD 1906 eruption, such areas were affected by detachment phenomena and erosion processes that removed material and lowered the pre-existing topography surface as inferred from the profiles. Excluding such specific areas, the 1906 profiles show that the flanks of the 1876 volcano have been covered by the products erupted by the subsequent v. canic activity (from 16/12/1881 to 22/04/1906, see Table 1). Such products reach significant thicknesses (> 40 m) in the NW, NE and SE quadrants, whereas in the SW quadrant their th ckn sses are negligible in the SSW portion (as the values result equal or slightly greater than the elevation error), and around 10 - 15 m in the WSW portion. In detail, the 40 m of difference shown in N-S and W-E profiles (Fig. 7) could be reasonably correlated with the product erupted during the 1891-94 and 1895-99 events, whose deposits formed, respectively, the "Colle Margherita" and "Colle Umberto" morphologies (see Figs. 3b and 6 and also Scandone et al. 2008). In particular, the "Colle Margherita" edifice, now partially buried by the subsequent products, might be identified in the North part of the N-S<sub>2</sub> profile (Fig. 7).

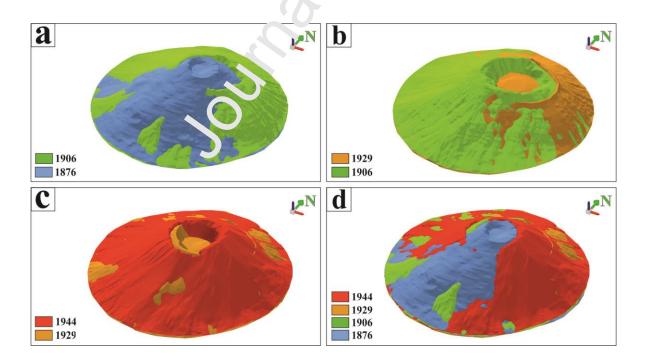
The comparisons between the 1906 and 1929 profiles highlight very similar topographies. This suggests that the AD 1929 eruption did not produce significant deposits on the flanks and that the 25-30 m discrepancy, highlighted in the central portion of the NW-SE<sub>1</sub> profile (Fig. 8) and reduced to around 10 m considering the elevation errors (Table 2), is probably not real. This hypothesis

seems to be supported also by the fact that the discrepancy remains constant along the entire central portion of the profile. In our opinion, such discrepancy of 10 meters could be explained considering that the central portion of NW-SE<sub>1</sub> profile crosses a particular area of the SW quadrant where the historical maps draw the zone affected by the eruptive fissure named "Crepaccio 1767" (visible in Fig. 3a) with morphologies poorly resolved and therefore difficult to draw during the digitalization. For this reason, it is possible that some contour lines of the older map (1906), characterized by a more detailed scale (1:10,000), have been directly copied to the younger one (1929) acquiring the altimetry error of the younger map (Fig. 2, Table 2) that is produced at less detailed scale (1:25,000). These issues are often common when working with historical maps, and should be taken into account to derive the most accurate descriptions from profile analysis. Also the comparison between the 1944 topography and the antecedent 1929 on, shows significant differences in height (> 40 m) in the N-S<sub>3</sub> and W-E<sub>3</sub> profiles (Fig.  $\bigcirc$  In addition, most of the profiles identify morphologies clearly correlated to peculi.<sup>•</sup> d posits generated from the AD 1944 eruption and named "Hot avalanche" (Santacroce and Lbrana 2003; Cole and Scarpati 2010 - see also Fig. 3d). These deposits are well recognizable in the N-S<sub>2</sub>, N-S<sub>3</sub>, NW-SE<sub>1</sub> and NW-SE<sub>3</sub> and are in agreement with the maps of Figs. 3d and 5, where the deposits cover several portions of the volcano flanks mainly involving the NW, NE and SE quadrants. As described by Hazlett et al. (1991), such peculiar type of debris avalances, triggered by a seismic crisis, took place near the end of the main eruptive phase (although long before the end of the eruption) and removed  $\sim$ 5-10×10<sup>6</sup> m<sup>3</sup> of material from elevations > 900 m a.s.l. in the Gran Cono area. The "Hot avalanche" deposits in the profiles reach thicknesses of 20 - 25 m, consistent with those reported in Hazlett et al. (1991). The last observations inferred from the profiles are related to some particular morphologies named "parasitic effusive vents". Such morphologies, well recognizable also on the shaded relief derived topographies (Fig. 3), are visible in the NE-SW<sub>1</sub> profile referred to 1876 topography (Fig. 8). In fact, such profile is characterized by a sequence of concave/convex shapes that could indicate the presence of some parasitic effusive vents located in the NW quadrant (see Fig. 6). In particular, the

convex shape to the NE might be correlated with the 1872 vent, whereas the other shapes can be correlated with the two buried vents visible along the NE-SW<sub>1</sub> trace of Figure 5. In addition, the 1872 vent can be also identified in the 1944 NE-SW<sub>1</sub> profile by a peculiar angular shape located in the same position assumed by the vent itself in the 1876 profile. Such vent is identified also in the 1944 shaded relief map (Fig. 3d).

In order to improve the discussion about the "Hot avalanche" deposits and to provide some quantitative measurements on the volumetric changes along the Gran Cono flanks, we have analysed the slope maps (Figure 5) overlaying the boundaries of the "Hot avalanche" deposits and the loading isolines  $(kg/m^2)$  of the tephra fallout deposits  $c^{f}$  the AD 1944 eruption. Despite the different degree of resolution of the maps, Figure 5 allows further observations beyond those already discussed in the section 4.1. In detail, the slope maps help to better constrain the limits of the AD 1944 "Hot avalanche" deposits. In particular, Figure 8c and 8d, highlight that for such deposits, in agreement with what observed by Hazlett et al. (1991), i) their source areas are located in zones characterized by slopes  $> 25^{\circ}$  and that ii) their accumulation areas are located in correspondence of areas with slores ranging between 6° and 15°, recognized in literature as thresholds values for the deposition phase (Bisson et al. 2007 and reference therein). In addition, Figures 8c and 8d highlight that such deposits are not present in the SW flank suggesting that their occurrence could depend of the sector of the volcano. The "sector-dependency" hypothesis has been already formulated by Hazlett et al. (1991), who explained that the AD 1944 hot avalanches developed along slopes characterized by at least one of the following elements: i) being preloaded with AD 1944 lava flows (for the N, S and W deposits, see also Fig. 6) or ii) having had high accumulation of tephra deposits from both the AD 1906 and AD 1944 eruptions (for the E deposits, see loading lines of both AD 1906 and AD 1944 eruptions in Fig. 5). None of these two elements characterize the SW flank of Gran Cono.

The visualization of the above-mentioned main topography changes can be better appreciated through the 3D perspective views of the Gran Cono performed at different time frames and shown in Figure 9. The 3D views allow to present the main steps of the morphological evolution of the Gran Cono. Firstly, we observe a crater area (rim and bottom) prevalently dismantled at least until the 1929 eruption (Fig. 9a,b). With the AD 1944 eruption, the trend changes only for the eastern crater portion that, with a remarkable re-growth, reaches elevations close to those of the 1876 crater rim (Fig. 9c). Moreover, if we consider the entire investigated period (1876 - 1944) as a whole (Fig. 9d), it is possible to observe how the South flank of the Gran Cono. after the AD 1944 eruption, does not reach the elevations of the 1876 oldest topography. Yor's respect to the flanks, it is evident a progressive growth that involves the NW, NE, and SF scators. These observations are in good agreement with the fact that the crater areas of the volcances are generally characterized by cycles of erosion and growth in elevation, whereas the crares are generally affected by an increase in elevation due the deposition and accumula, on of products emitted during an eruption, such as, for example, lava flows and fall out deposits.



**Figure 9** – 3D perspective views of the morphological changes for different time frames: a) 1876-1906; b) 1906-1929; c) 1929-1944; d) the whole investigated period (1876 - 1944).

#### 4.2 Crater rims comparison

The morphological changes of the crater area through time were investigated not only through the elevation profiles (Figs. 6 and 7) but also by analysing some specific geometric parameters of the crater area itself (see Table 3). These elements were obtained after the digitalization of the four crater rims and the representation of their geometric centroids (Fig. 10). Figure 10 shows significant changes in the extent of the crater rim that mainly involve the E and N-E areas of the Gran Cono summit portion. In addition, Figure 10 highlights an alignment of the centroids along the NE-SW direction suggesting possible displacements of the eruptive central along such direction. These displacements, quantified with the parameters reported in Table 3, are calculated considering the errors in x and y previously defined (Table 2). The results show a real migration of at least 50 m of the crater centroid towards NE from 1876 to 1929, and of at least 170 m towards SW between 1929 and 1944. Moreover, it is evident a substantial entargement of the crater rim from 1876 to 1929, in fact its extent remarkably increases from 0.0c km<sup>2</sup> to 0.46 km<sup>2</sup>. Subsequently, with the AD 1944 eruption, the crater rim significantly reduces its extent, reaching the current area of ~0.25 km<sup>2</sup>.

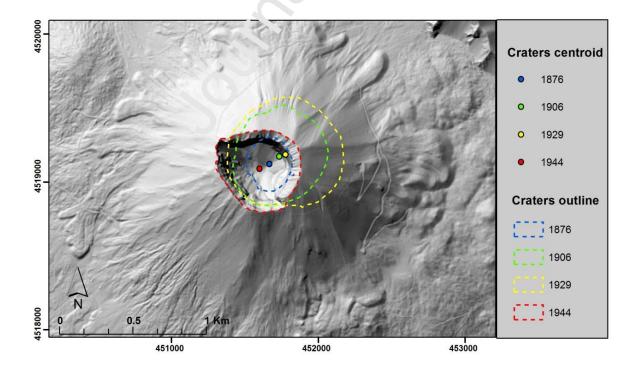


Figure 10 - Crater rim traces (1876, 1906, 1929, 1944)

Year	Crater centroid X coordinate (m)	Crater centroid Y coordinate (m)	Crater Extent (m <sup>2</sup> )	Crater centroid distance from the 1944 one (m)	Crater centroid distance from previous one (m)	Crater centroid migration direction from previous one
1876	451667±18	4519121±23	85345	75±23	-	-
1906	451732±15	4519172±13	329651	156±20	61(35 - 130)	NE
1929	451774±19	4519186±16	463629	200±27	45 (16 - 92)	NE
1944	451599	4519088	248729	-	200 (173 – 227)	SW

**Table 3**: Gran Cono crater rim parameters calculated in different time.

The values in brackets indicate the minimum and maximum distance a cording to the x and y errors of the respective maps.

By analysing more in detail the spatial position of the crater centroid during the investigated period, we can observe a progressive migration of the centroid towards NE from 1876 to 1929, and then, during the 1929-1944 period, towards the S<sup>VV</sup> (Fig. 10). This latter migration seems to be contemporaneous to the strong increase in height of crater rim in the NE quadrant (see section 4.1), whereas the first migration is concomitent with a progressive lowering of the crater rim in the NW and SW quadrants. The distances between the centroids of all crater rims have been estimated considering the errors in x and  $\sim c_{1}$  culated after the co-registration procedure (Table 2 in section 3.3). This preferential direction of centroid migration (NE-SW) is parallel to the strike of one of the two main deep fault systems below SV (Bianco et al., 1998; Tadini et al., 2017; Tramparulo et al., 2018). In detail, the 1876 crater rim covers an area of 0.08 km<sup>2</sup> and its centroid is positioned at a distance of  $75 \pm 23$  m in NE direction with respect to 1944 centroid (see Fig. 10). With the AD 1906 eruption, the crater rim enlarges defining an area of ~  $0.33 \text{ km}^2$ , lowers in height reaching the maximum elevation of ~1220 m a.s.l. (Figs. 6 and 7) and migrates to the NE of at least 35 m. With the 1929 eruption, the crater rim enlarges even more ( $\sim 0.46 \text{ km}^2$ ) and its centroid migrates again towards NE of at least ~16 m (Table 3). Although the period 1906-1929 is not characterized by major eruptions (see Table 1 and Scandone et al. 2008), the intra-crater activity does not stop. In fact, as described in Scandone et al. (1993), in this period the crater bottom is affected by two

collapses (in the years 1911 and 1913) and lava fillings, particularly in the years 1917, 1919 and 1920. These activities have probably generated the enlargement and a lowering of the crater rim that reaches an elevation not exceeding 1176 m a.s.l. (Fig. 6). This value is probably already reached by the 1920 crater rim as showed in the inset of Fig. 2c. This indicates that the intra-crater activity of the period 1920-1929 (see Scandone et al. 1993 for details) and, in particular, the AD 1929 eruption (Table 1 and Scandone et al. 2008) does not cause collapses on the crater rim, but, on the contrary, they contribute to reconstruct the Gran Cono original outline. Finally, the present day crater, generated by the AD 1944 eruption, shows a centroid migrated of at least 170 m towards SW aligned with the other centroids. The 1944 crater rim, covering an area of ~0.25 km<sup>2</sup>, results particularly reduced with respect to the previous one (~ 50%) and, in a portion of the NE edge, is very similar to the 1876 rim (Figs. 6 and 10). This recruction could suggest accumulations of material which, at least in part, may have occurre thetween 1929 and 1944. This hypothesis is supported by Cole and Scarpati (2010) and Scarpati done et al. (1993). In fact, Cole and Scarpati (2010) confirm that, prior to the AD 1944 eruption, the crater was nearly completely filled with a small conelet in the middle area and Scandone at al. (1993) points out that in the period 1929-1944 many lavas (1930, 1935, 1936, 1937, 1939, 1940, 1942) outflows from the crater toward N, NE and E. These outflows, at least in part, could explain the growth of the Gran Cono in the NE quadrant, where the crater rim reaches values higher than 1250 m a.s.l. (see W-E<sub>2</sub> and NE-SW<sub>2</sub> profile in Fig. 6 and Fig. 7, respectively). Also Cole and Scarpati (2010) support such interpretation pointing out that during all the phases of the AD 1944 eruption, most of the pyroclastic products have accumulated in the NE and SE quadrant of the Gran Cono edifice. After the emptying of most of the intra-crater material, the AD 1944 eruption carved the present shape of the Gran Cono, whose crater, in our model, reaches the highest elevation (1277 m a.s.l.) in the NE quadrant. This value is in good agreement with the measurements of Imbò (1949).

#### 4.3 Volume and surface changes estimations

Estimates of volume and surface change of the Gran Cono edifice were made by processing the altimetry models with the specific matrix operators available on the GIS platform. In detail, we calculated the volumes by using altimetry models stored in ESRI - GRID format with a spatial resolution of 5 m and the surface from the ESRI - TIN vector format. The calculations were performed starting from a plane coincident with the circle having radius of 1 km (see Fig. 4). Table 4 reports the values of the surface and volume of the Gran Conc edifice for each altimetry model highlighting a general increase both in volume and in surface from  $1\sqrt{76}$  to 1944.

Table 4 - Volume	s and 3D Surfaces	of the Gran	Cono through time.
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Year	Gran Cono volume (m <sup>3</sup> )	Volumet ac change åre m preae is (13)	Gran Cono surface (m <sup>2</sup> )
1944	824.1×10 <sup>6</sup>	+51.4×. <sup>\\\\</sup> (±5%)	3871135
1929	772.7×10 <sup>6</sup> (±4.9%)	・ <sup>^</sup> 7 ×10 <sup>6</sup> (±7%)	3751222±3800
1906	745.7×10 <sup>6</sup> (±2%)	54.2×10 <sup>6</sup> (±7%)	3835957±2420
1876	691.5×10 <sup>6</sup> (±、2%)	-	3725202±3670

The error in % of each volume (in brackets) was calculated considering the errors in height reported in the section 3.2.

This trend is coherent with the fact that the 1876 - 1944 period was overall characterized by an intense volcanic activity (Scandone et al., 2008) resulting in a total volume increase of the Gran Cono equal to ~  $133 \times 10^6$  m<sup>3</sup> (Table 4). Most of this increase in volume is achieved during the 1876-1906 and 1929 -1944 volcanic activities. During the 1876-1906 period, the volume of Gran Cono increases of ~ 54 × 10<sup>6</sup> m<sup>3</sup> (Table 4) including two main events such as the building of "Colle Margherita" (1891- 94; Table 2) and the AD 1906 eruption. In general, the products of the AD 1906 eruption correspond to deposits with thicknesses of 30 - 40 m well recognizable in several profiles

of the Figs. 6 and 7 and consist of lava flows and proximal tephra. With the aim to calculate the volume of the proximal tephra along the Gran Cono flanks, we have superimposed the loading isolines of the AD 1906 tephra fallout deposits (Fig. 5b), individuated those that involve the Gran Cono area and worked only on the portions comprised between the identified loading lines. For each portion, the mass was calculated by multiplying the area for the mean value between the two involved loading lines. Then, the sum of all masses was transformed in volume considering a deposit density value equal to  $1100 \text{ kg/m}^3$  (Barsotti et al., 2015). The resulting volume for the AD 1906 tephra fallout deposits along the Gran Cono is  $1.13 \times 10^6 \text{ m}^3$  Thir value accounts for ~2.0% of the volume increase between 1876 and 1906 (Table 4) and ~1.5% of the total tephra deposit volume of the AD 1906 (April) eruption (Table 2).

The second relevant volumetric increase is referred to the 1929-1944 period and corresponds to ~  $51 \times 10^6$  m<sup>3</sup> of erupted material and accumulated to the Gran Cono area (Table 4). During this time frame, the most important eruption is the AL 1944 one (Table 2), which produced lava flows, hot avalanches and tephra fall out deposits Trobably, part of the tephra was emplaced on the flanks of the edifice, causing a certain growth in the edif of the Gran Cono structure (see profiles in Figs. 7 and 8). By using the same procedule performed for the AD 1906 eruption, we have calculated the volume of tephra fallout deposits of the AD 1944 eruption along the Gran Cono (Figure 8d). The resulting value is  $1.25 \times 10^6$  n<sup>3</sup> and corresponds to ~2.4 % of the volumetric increase reported in Table 4 and ~1.9% of the total tephra of the AD 1944 eruption (Table 2).

Also the surface of the Gran Cono (Table 3) shows an increase of ~0.14 km<sup>2</sup> from 1876 to 1944, although being not progressive through time. In detail, the first increase (from 1876 to 1906) is equal to ~ 0.11 km<sup>2</sup> and can be explained considering: i) the large and thick mantling of the Gran Cono flanks generated by the accumulation of the 1906 erupted deposits and ii) the contemporaneous remarkable enlargement of the related crater rim, whose extent is four times greater than the previous one (0.32 km<sup>2</sup> vs. 0.08 km<sup>2</sup>, see Table 3). Subsequently, with the 1929

eruption, the surface decreases of ~  $0.08 \text{ km}^2$  despite the further enlargement of the crater rim. This decrease can be explained considering that the strong morphological changes occurred between 1906 and 1929 are mainly due to an intra-crater activity (Scandone et al., 2008) that has filled the conduit previously emptied by the AD 1906 eruption (Bertagnini et al., 1991) raising the crater bottom at higher elevations. This feature, coupled with the lowering of the crater rim elevation of the 1929 profiles (see Fig. 6 and 7), contributed to decrease the extent of the surface of the inner part of the crater. A further reason of such decrease in surface could be partially due to the different scales of the original maps. In fact, the contour lines every 25 m on the 1929 map creates a much less detailed morphology with respect to the 1906 one (genericed with contour lines every 5 m) and, consequently, a lower value of the calculated surface. Fin. 11y, with the AD 1944 eruption, the surface of the crater further increases reaching the great  $s_1$  value of ~3.87 km<sup>2</sup> corresponding to the nowadays surface. This value is linked to: i) the on the centre of lava flows, tephra fall out and hot avalanches related to the eruption itself; ii) the increase of the crater rim elevation; iii) the remarkable lowering of the crater botton, that, consequently, increased the surface of the inner portions of the crater; iv) the very high resolution of the 1944 topography that allows to quantify the area of many morphological detail not representable in the other topographies.

# 5 CONCLUSIONS

This work presents a study on the main morphological changes of Somma – Vesuvio (SV) caldera from 1876 to 1944 based on a multi-temporal analysis of the volcano topography. For the first time, three DEMs coming from historical cartography and a DEM derived by ALS technology are compared and analysed providing a quantification of the volcano morphological changes associated to an error estimation, fundamental element to understand the accuracy/reliability degree of the obtained results. Because our results are very consistent with the available literature data, this study testifies that historical topographies of volcanic areas, if elaborated with rigorous criteria, allow to

derive reliable quantitative and qualitative data very useful to improve the knowledge of the volcano in the past.

Four 3D topography models of Somma Vesuvio summit caldera were here reconstructed focusing then the attention on a specific zone represented by the Gran Cono edifice. The four models reproduce the SV summit caldera topography of 1876, 1906, 1929 and 2012 year. The first three models derive from the IGM historical maps, whereas the fourth model derives from the ALS data acquired on 2012 and it was assumed to be representative of the volcano topography immediately after the 1944 eruption. The four models, produced and an lyst 1 on GIS platform at spatial resolution of 5 m, were firstly elaborated to obtain two maps (shaded reliefs and slopes) highlighting how the morphological changes on the Gran Cono involve not only the "flanks" zone with the accumulation of several volcanic deposits (mostly lava flows and pyroclastic deposits), but also the "crater" zone that significantly changes in shape and dimension. The results indicate that from 1876 to 1944 the Gran Cono edifice gro 's in volume of  $133 \times 10^6$  m<sup>3</sup> (±5%) and in surface of ~0.14 km<sup>2</sup>. The volume growth is main<sup>1</sup>, <sup>1</sup>ue to the AD 1906 and AD 1944 eruptions that generate, respectively, an increase of  $54.2 \times 10^{\circ}$  h<sup>-3</sup> (±7%) and  $51.4 \times 10^{6}$  m<sup>3</sup> (±5%), whereas with the 1929 eruption the growth is around  $27 \cdot 10^6 \text{ m}^3$  ( $\pm 7\%$ ) corresponding only to the 20 % of the total growth. Also the surface growth is mainly due to the AD 1906 and AD 1944 eruptions but it is not progressive from 1876 to 1)44: in fact, after a consistent increase occurred with the AD 1906 eruption (~ 0.11 km<sup>2</sup>), the AD 1929 eruption generated a surface reduction of ~0.08 km<sup>2</sup>, followed by another increase of ~  $0.12 \text{ km}^2$  that allowed to reach the current value of ~ $3.87 \text{ km}^2$ . The volume and surface growths meet the observations derived from the heights profiles (12 for each altimetry model), showing an increase in elevation of at least 40 m in the Gran Cono flanks, especially in the NW, NE and SE quadrants, and a lowering of the crater bottom of at least 25 m from 1876 (1210  $\pm$ 11 m a.s.l.) to 1944 (970 m a.s.l.). In addition, the profiles well identify a series of volcanic morphologies related to specific eruptions, in particular: i) the "Colle Margherita" "exhogeneous tholoid" (or "dome-shaped accumulations of the lava flows"), formed during the AD 1891-94

eruption, is well identified by the profiles crossing the N quadrants. For this structure, a thickness of around 40 m is estimated in the N-NW flank of the Gran Cono; ii) the "Hot avalanche" deposits of the AD 1944 eruption, well recognizable in several profiles crossing the NW, NE and SE quadrants, show a thickness of 20-25 m, a value in good agreement with the estimates of Hazlett et al. (1991); iii) the parasitic effusive vents detected by the 1876 NE-SW<sub>1</sub> profile and probably correlated with the 1872 vent and with some buried vents (of age unknown). However, the most peculiar changes seem to be referred to the 1906 - 1929 period when the Gran Cono increases in volume of ~  $27 \times 10^6$  $m^3$  and, contemporaneously decreases in surface of 0.08 km<sup>2</sup> These changes can be actually justified considering the intra-calderic activity of such period na producing abundant material, has strongly filled the crater forming also new small cones, and, in particular, has increased the elevation of the crater bottom thus generating a decrease of the flanks surfaces inside the crater. Also, the crater rim shows significant modification. Consisting of a remarkable enlargement from 1876 to 1929 (from 0.08 km<sup>2</sup> to 0.46 km<sup>2</sup>, followed, during the AD 1944 eruption, by a strong reduction resulting in ~ 0.24  $\text{km}^2$  and preserved until now. These enlargements are contemporaneous to the progressive NE migration of the crater centroid of at least 50 m followed, with the AD 1944 eruption, by relation SW migration of at least 170 m. Such migrations, estimated taking into account the errors in x and y of source data, occur along the SW - NE direction individuated by up augnment of the centroids. Such alignment is in agreement with one of the two main deep faults system trend located below Somma -Vesuvio. In conclusion, despite the different typology and accuracy degree of the data source, this study allowed to better document the morphological changes of a portion of the Somma-Vesuvio volcano from 1876 to 1944, providing quantifications of these changes that, associated to an estimation of error, result in very good agreement with the available literature data.

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#### References

Arrighi, S., Principe, C., and Rosi, M. (2001). Violent strombolian and subplinian eruptions at Vesuvius during post-1631 activity. *Bulletin of Volcanology*, 63(2-3), 126-150, doi: https://doi.org/10.1007/s004450100130.

Annis, A., Nardi, F., Petroselli, A., Apollonio, C., Arcangeletti, E., Tauro, F., ... & Grimaldi, S. (2020). UAV-DEMs for Small-Scale Flood Hazard Mapping. *Water*, 12(6), 1717.

Balletti, C. (2000). Analytical and quantitative methods for the analysis of the geometrical content of historical cartography. *International Archives of Photogra nm. try and Remote Sensing*, 33(B5/1; PART 5), 30-37.

Barsotti, S., Neri, A., Bertagnini, A., Cioni, R., Mula, M., and Mundula, F. (2015). Dynamics and tephra dispersal of Violent Strombolian eregetices at Vesuvius: insights from field data, wind reconstruction and numerical simulation of the '906 event. *Bulletin of Volcanology*, 77, 1-19.

Bertagnini, A., Landi, P., Santacroce, R, and Sbrana, A. (1991). The 1906 eruption of Vesuvius: from magmatic to phreatomagmatic activity through the flashing of a shallow depth hydrothermal system. *Bulletin of Volcanolog*, 52(7), 517-532, doi: https://doi.org/10.1007/BF00298153.

Bertagnini, A., Landi, P., K. si, M., and Vigliargio, A. (1998). The Pomici di Base plinian eruption of Somma-Vesuvius. Journal of Volcanology and Geothermal Research, 83(3-4), 219-239, doi: https://doi.org/10.1016/S0377-0273(98)00025-0.

Bianco, F., M. Castellano, G. Milano, G. Ventura, and G. Vilardo (1998), The Somma- Vesuvius stress field induced by regional tectonics: Evidences from seismological and mesostructural data, *J. Volcanol. Geotherm. Res.*, 82(1), 199–218, doi: https://doi.org/10.1016/S0377-0273(97)00065-6

Bisson, M., Pareschi, M.T., Zanchetta, G., Sulpizio, R., Santacroce R. (2007). "Volcaniclastic debris flow occurrences in the Campania region (southern Italy) and their relation to Holocene -

Late Pleistocene pyroclastic fall deposits: implications for large-scale hazard mapping", *Bulletin of Volcanology*, 70, 157–167, doi: 10.1007/s00445-007-0127-4

Bisson, M., Spinetti, C., Neri, M., and Bonforte, A. (2016). Mt. Etna volcano high resolution topography: Airborne Lidar modelling validated by GPS data. *International Journal of Digital Earth*, 9(7), 710-732, doi: https://doi.org/10.1080/17538947.2015.1119208.

Bisson, M., Paolillo, A., Tadini, A., Sulpizio, R., & Zanchetta, G. (2018). Volcanoclastic flow hazard assessment in highly populated areas: a GIS-based approach applied to Torre del Greco municipality (Somma-Vesuvius, Italy). *Geosciences Journal*, 22(2):001-522.

Bisson M., Tadini A., Gianardi R., & Angioletti A. (2020) Morphological Evolution of Somma-Vesuvio During the Last Century: Integration Between Fistorical Maps and Airborne LiDAR Survey. In: Parente C., Troisi S., Vettore A. (eds' K.<sup>2</sup> <sup>2</sup> <sup>1</sup> Geomatics: Research, Results and Review. R3GEO 2019. *Communications in Computer and Information Science*, vol 1246. Springer, Cham. https://doi.org/10.1007/978-3-030-62800-0\_24.

Beyer, R.A, Alexandrov, O., and McMichael, S. (2018). The Ames Stereo Pipeline: NASA's Open Source Software for Deriving and Processing Terrain Data. *Earth and Space Science*, 5(9), 537-548, doi: https://doi.org/10.025/2018EA000409.

Branca, S., and Abate, T. (2017). Current knowledge of Etna's flank eruptions (Italy) occurring over the past 2500 years. From the iconographies of the XVII century to modern geological cartography. *Journal of Volcanology and Geothermal Research*, doi: https://doi.org/10.1016/j.jvolgeores.2017.11.004.

Cappello, A., Vicari, A., & Del Negro, C. (2011). Assessment and modeling of lava flow hazard on Mt. Etna volcano. *Bollettino di Geofisica Teorica e Applicata*.

Cappello, A., Ganci, G., Bilotta, G., Herault, A., Zago, V., & Del Negro, C. (2018). Satellite-driven modeling approach for monitoring lava flow hazards during the 2017 Etna eruption. Annals of Geophysics.

Cioni, R., Bertagnini, A., Santacroce, R., and Andronico, D. (2008). Explosive activity and eruption scenarios at Somma-Vesuvius (Italy): towards a new classification scheme. *Journal of Volcanology and Geothermal Research*, 178(3), 331-346, doi: https://doi.org/10.1016/j.jvolgeores.2008.04.024.

Cioni, R., Santacroce, R., and Sbrana, A. (1999). Pyroclastic depc\_<sup>its</sup> as a guide for reconstructing the multi-stage evolution of the Somma-Vesuvius Caldera. *Bulletin of Volcanology*, 61(4), 207-222, doi: https://doi.org/10.1007/s004450050272.

Cole, P. D., and Scarpati, C. (2010). The 1944 eruption of visuvius, Italy: combining contemporary accounts and field studies for a new volcanological *rec*construction. *Geological Magazine*, 147(3), 391-415, doi: https://doi.org/10.1007/s0044\_00.0272.

Dai, C., & Howat, I. M. (2017). Measure of laws with ArcticDEM: Application to the 2012–2013 eruption of Tolbachik, Kamch. tka. *Geophysical Research Letters*, 44(24), 12-133.

De Beni, E., Cantarero, M., & Mersina, A., 2019: UAVs for volcano monitoring: A new approach applied on an active lava florr on Mt. Etna (Italy), during the 27 February–02 March 2017 eruption. Journal of Volcanology and Geothermal Research, 369, 250-262.

De Leeuw, A. J., Veugen, L. M. M., & Van Stokkom, H. T. C. (1988). Geometric correction of remotely-sensed imagery usiing ground control points and orthogonal polynomials. International journal of remote sensing, 9(10-11), 1751-1759.

Favalli, M., and Pareschi, M. T. (2004). Digital elevation model construction from structured topographic data: the DEST algorithm. *Journal of Geophysical Research: Earth Surface*, *109*(F4), doi: https://doi.org/10.1029/2004JF000150.

Favalli, M., Innocenti, F., Teresa Pareschi, M., Pasquarè, G., Mazzarini, F., Branca, S., ... & Tibaldi, A. (1999). The DEM or Mt. Etna: geomorphological and structural implications. *Geodinamica Acta*, 12(5), 279-290.

Felpeto, A., Araña, V., Ortiz, R., Astiz, M., & García, A. (2001). Assessment and modelling of lava flow hazard on Lanzarote (Canary Islands). *Natural hazards*, 23(2-3), 247-257.

Ferretti, A., Prati, C., & Rocca, F. (1999). Multibaseline InSAR DEM reconstruction: The wavelet approach. *IEEE Transactions on geoscience and remote sensing*, 3 (2), 705-715.

Ganci, G., Cappello, A., Bilotta, G., Herault, A., Zago, V, & Del Negro, C. (2018). Mapping Volcanic Deposits of the 2011–2015 Etna Eruptive Events Using Satellite Remote Sensing. Frontiers in Earth Science, 6, 83.

Garvin, J. B., Slayback, D. A., Ferrini, V., Frowny, J., Giguere, C., Asrar, G. R., & Andersen, K. (2018). Monitoring and Modeling the Rapid Evolution of Earth's Newest Volcanic Island: Hunga Tonga Hunga Ha'apai (Tonga) Using *Falgh* Spatial Resolution Satellite Observations. *Geophysical Research Letters*, 45(8), 3445-3452 Gwinner, K., Coltelli, M., Flohrera, J., Jaumanna, R., Matza, K. D., Marsella, M., Roatscha, T., Coholtena, F., and Trauthana, F. (2006). The HRSC-AX Mt. Etna Project: High-resolution O the mages and 1 m DEM at Regional Scale. In *Proceedings ISPRS XXXVI*, *T05–23 (Part 1)*.

Hazlett, R. W., Buesch, D., Anderson, J. L., Elan, R., and Scandone, R. (1991). Geology, failure conditions, and implications of seismogenic avalanches of the 1944 eruption at Vesuvius, Italy. *Journal of Volcanology and Geothermal Research*, 47(3-4), 249-264, doi: https://doi.org/10.1016/0377-0273(91)90004-J.

Huggel, C., Schneider, D., Miranda, P. J., Granados, H. D., & Kääb, A. (2008). Evaluation of ASTER and SRTM DEM data for lahar modeling: a case study on lahars from Popocatépetl Volcano, Mexico. *Journal of Volcanology and Geothermal Research*, 170(1-2), 99-110.

Imbò, G. (1949). Considerazioni sulla presente attività del Vesuvio. *Bulletin Volcanologique*, 8, 123–132, doi: https://doi.org/10.1007/BF02596785

ISTAT (2011), Censimento popolazione 2011, Istituto Nazionale di Statistica, Database accessed March 2015.

James, M. R., & Robson, S. (2012). Straightforward reconstruction of 3D surfaces and topography with a camera: Accuracy and geoscience application. *Journal of Geophysical Research: Earth Surface*, 117 (F3).

Jones, L. K., Kyle, P. R., Oppenheimer, C., Frechette, J. D., 'z O'al, M. H. (2015). Terrestrial laser scanning observations of geomorphic changes and varying lava lake levels at Erebus volcano, Antarctica. *Journal of Volcanology and Geothermal Research*, 295, 43-54.

Marsella, M., Baldi, P., Coltelli, M., & Fahris, M. (2012). The morphological evolution of the Sciara del Fuoco since 1868: reconstructing the effusive activity at Stromboli volcano. Bulletin of Volcanology, 74(1), 231-248.

Müller, D., Walter, T. R., Schöp I, F. Witt, T., Steinke, B., Gudmundsson, M. T., & Dürig, T. (2017). High-resolution digital elevation modeling from TLS and UAV campaign reveals structural complexity at the 2014/2015 Uclubraun eruption site, Iceland. *Frontiers in Earth Science*, 5, 59.

Naddeo, B. A. (2004). Topographies of Difference: Cartography of the City of Naples, 1627–1775. *Imago Mundi*, 56(1), 23-47, doi: https://doi.org/10.1080/0308569032000172932.

Neri, M., Mazzarini, F., Tarquini, S., Bisson, M., Isola, I., Behncke, B., and Pareschi, M.T. (2008). Topographic Map of the Summit Craters of Mount Etna from The changing face of Mount Etna's summit area documented with Lidar technology. *Geophysical Research Letters*, vol. 35, L09305, doi:10.1029/2008GL033740 Palmentieri, S. (2016). La cartografia storica per l'analisi dell'evoluzione funzionale delle ville vesuviana. *Bollettino della Associazione Italiana di Cartografia*, 158, 120-130, doi: 10.13137/2282-572X/20336.

Pesci, A., Fabris, M., Conforti, D., Loddo, F., Baldi, P., & Anzidei, M. (2007). Integration of ground-based laser scanner and aerial digital photogrammetry for topographic modelling of Vesuvio volcano. *Journal of Volcanology and Geothermal Research*, 162(3-4), 123-138

Pizzimenti, L., Tadini, A., Gianardi, R., Spinetti, C., Bisson, M., a. <sup>4</sup> Brunori, C. A. (2016). Digital Elevation Models derived by ALS data: Sorrentina Peninsula cost areas. *Rapporti tecnici INGV*, n°361.

Principe, C., Tanguy, J. C., Arrighi, S., Paiotti, A., Le Goti, M., and Zoppi, U. (2004). Chronology of Vesuvius' activity from AD 79 to 1631 based on archeomagnetism of lavas and historical sources. Bulletin of Volcanology, 66(8), 70, 774, doi: https://doi.org/10.1007/s00445-004-0348-8.

Ronza, M. (2017). Ville e processi di urba sizzazione: la collina del Vomero nella città di Napoli. La cartografia storica per il cultural pranning. *Bollettino della Associazione Italiana di Cartografia*, 159, 54-68, doi: 10.13137/2282-572x/20559.

Rosi, M., Principe, C., a. 1 .<sup>1</sup> e<sup>-</sup> ci, R. (1993). The 1631 Vesuvius eruption. A reconstruction based on historical and stratigraphical data. *Journal of Volcanology and Geothermal Research*, 58(1), 151-182, doi: https://doi.org/10.1016/0377-0273(93)90106-2.

Santacroce, R. and Sbrana, A. (2003). The Vesuvius geological map, CARG project. *Geological Survey of Italy*. Editor: SELCA Firenze.

Santacroce, R., Cioni, R., Marianelli, P., Sbrana, A., Sulpizio, R., Zanchetta, G., Donahue, D. J. and Joron, J. L. (2008). Age and whole rock–glass compositions of proximal pyroclastics from the major explosive eruptions of Somma-Vesuvius: A review as a tool for distal

tephrostratigraphy. *Journal of Volcanology and Geothermal Research*, 177(1), 1-18, doi: https://doi.org/10.1016/j.jvolgeores.2008.06.009.

Sbrana, A., Cioni, R., Marianelli, P., Sulpizio, R., Andronico, D., and Pasquini, G. (2020). Volcanic evolution of the Somma-Vesuvius Complex (Italy). Journal of Maps, 1-11, doi: https://doi.org/10.1080/17445647.2019.1706653.

Scandone, R., Giacomelli, L., and Gasparini, P. (1993). Mount Vesuvius: 2000 years of volcanological observations. *Journal of Volcanology and Geothermal Research*, 58(1), 5-25, doi https://doi.org/10.1016/0377-0273(93)90099-D.

Scandone, R., Giacomelli, L., and Speranza, F. F. (2008) Det. stent activity and violent strombolian eruptions at Vesuvius between 1631 and 1944. *Journa' of Volcanology and Geothermal Research*, 170(3–4), 167-180, doi: https://doi.org/10.1016/j.jvo.gcores.2007.09.014.

Salvatici, T., Di Roberto, A., Di Traglia, F., Lisson, M., Morelli, S., Fidolini, F., ... & Casagli, N. (2016). From hot rocks to glowing ava'a ches: numerical modelling of gravity-induced pyroclastic density currents and hazard maps at the Etromboli Volcano (Italy). *Geomorphology*, 273, 93-106.

Shewchuk, J. R. (2002). Lolau ay refinement algorithms for triangular mesh generation. *Computational geometry*, 22(1-2), 21-74, doi: https://doi.org/10.1016/S0925-7721(01)00047-5.

Sigurdsson, H., Carey, S. N., Cornell, W., and Pescatore, T. (1985). The eruption of Vesuvius in AD 79. *National Geographic Research*, 1(3), 332-387.

Tadini, A., Bisson, M., Neri, A., Cioni, R., Bevilacqua, A., and Aspinall, W. P. (2017). Assessing future vent opening locations at the Somma- Vesuvio volcanic complex: 1. A new information geodatabase with uncertainty characterizations. *Journal of Geophysical Research: Solid Earth*, 122(6), 4336-4356, doi: https://doi.org/10.1002/2016JB013858.

Tarquini, S., Isola, I., Favalli, M., Mazzarini, F., Bisson, M., Pareschi, M. T., and Boschi, E. (2007). TINITALY/01: a new triangular irregular network of Italy. *Annals of Geophysics*, 50(3), doi: http://hdl.handle.net/2122/3673.

Tramparulo, F. D. A., Vitale, S., Isaia, R., Tadini, A., Bisson, M., & Prinzi, E. P. (2018). Relation between alternating open/closed-conduit conditions and deformation patterns: An example from the Somma-Vesuvius volcano (southern Italy). *Journal of Structural Geology*, 112, 138-153.

Travaglini, D. (2004). Trasformazioni tra sistemi di coordinate. software disponibili, limiti e potenzialità. *Forest@-Journal of Silviculture and Forest*. *Ecology*, 1, 128-134, doi: 10.3832/efor0228-0010128.

Ventura, G., and Vilardo, G. (2008). Emplacement mer handsm of gravity flows inferred from high resolution Lidar data: The 1944 Somma–Vesuvit s here flow (Italy). *Geomorphology*, 95(3-4), 223-235, doi: https://doi.org/10.1016/j.geomorph. 2007.06.005.

Whelley, P. L., Garry, W. B., Hamiltor, C. W., and Bleacher, J. E. (2017). LiDAR-derived surface roughness signatures of basaltic lova types at the Muliwai a Pele Lava Channel, Mauna Ulu, Hawaii". *Bulletin of Volcanology*, 79(11), 75, doi: https://doi.org/10.1007/s00445-017-1161-5.

Wu, S. S. (1979). Photogrammetric portrayal of Mars topography. Journal of Geophysical Research: Solid Earth, 84(B14), 7955-7959.

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

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# Highlights

- A morphometric study of Somma-Vesuvio caldera is presented for 1876 1944 period
- Past and recent topographies are reconstructed from historical maps and ALS data
- The analyses are based on height profiles, volume and distance quantifications
- The quantifications of the morphological changes are associated to error estimations
- Reliable morphological quantitative data are derived from historical cartography