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18 **Paleomagnetic dating of **prehistoric** lava flows from the urban**  
 19 **district of Catania (Etna volcano, Italy)**

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31 **ABSTRACT**

32 **Determining the ages of past eruptions of active volcanoes whose slopes were historically**  
 33 **inhabited is vitally important** for investigating the relationships between eruptive phenomena and  
 34 human settlements. During its almost three-millennia-long history, Catania—the biggest city  
 35 lying at the toe of Etna volcano—was directly impacted only once by the huge lava flow  
 36 emplaced during the **A.D. 1669** **[[Author: GSA places A.D. before the year. This was**  
 37 **changed throughout. Periods after A.D., B.C., and B.P. were added, and commas were**  
 38 **removed from 4-digit numbers in keeping with GSA style.]]** Etna flank eruption. However,  
 39 other lava flows reached the present-day Catania urban district in **prehistoric** ages **before the**  
 40 **founding of the city in** Greek times (729/728 B.C., i.e., 2679/2678 yr B.P.). In this work, the  
 41 Holocene lava flows of Barriera del Bosco, Larmisi, and San Giovanni Galermo, **which are**  
 42 exposed in the Catania urban district, were paleomagnetically investigated at 12 sites (120  
 43 oriented cores). Paleomagnetic dating **was** obtained by comparing flow-mean paleomagnetic  
 44 directions to updated geomagnetic reference models for the Holocene. The Barriera del Bosco  
 45 flow turns out to represent the oldest eruptive event and is paleomagnetically dated to the

46 11,234–10,941 yr B.P. and 8395–8236 yr B.P. age intervals. The mean paleomagnetic directions  
47 from the San Giovanni Galermo and Larmisi flows overlap when statistical uncertainties are  
48 considered [\[\[OK?\]\]](#). This datum, along with geologic, geochemical, and petrologic evidence,  
49 implies that the two lava flows can be considered as parts of a single lava field that erupted in a  
50 narrow time window between 5494 yr B.P. and 5387 yr B.P. The emplacement of such a huge  
51 lava flow field may have buried several Neolithic settlements, which would thus explain the  
52 scarce occurrence of archaeological sites of that age found below the town of Catania.

## 53 INTRODUCTION

54 Mount Etna is the largest continental basaltic volcano in Europe and one of the most  
55 active in the world. It covers an area of ~1200 km<sup>2</sup> and has a maximum diameter of 45 km and a  
56 height of 3328 m above sea level (a.s.l.) (Fig. 1). During its eruptive history of the last 220 ka,  
57 Etna was characterized by magmas of predominantly basaltic composition and by volcanic  
58 events of variable intensity and magnitude that ranged from fissural to central activity, effusive  
59 to explosive phenomena, and strombolian to plinian eruptions (Branca et al., 2011a, 2011b; De  
60 Beni et al., 2011). The detailed knowledge of Etna's eruptive activity in the past 2700 years  
61 provided by historical accounts shows that the long-term behavior of the volcano, usually  
62 characterized by frequent and moderate summit eruptions, is accompanied by the occurrence of  
63 large flank eruptions such as that of A.D. 1669 (Branca and Del Carlo, 2005; Proietti et al., 2011;  
64 Branca and Abate, 2019). The SE Etna flank (Fig. 1), which is currently the most densely  
65 populated, is the volcano sector characterized by the highest number of flank eruptions in  
66 historical times. The town of Catania, which is located at the toe of the SE Etna flank and now  
67 hosts more than 330,000 inhabitants, was only directly impacted by the A.D. 1669 lava flow and  
68 was threatened in the twelfth century A.D. by the Mt. Arsi di Santa Maria lava flow (Branca et  
69 al., 2011b, 2016). As a consequence, Etna volcano hazard map (Del Negro et al., 2013) assigns a  
70 low-medium [\[\[low to medium?\]\]](#) probability of lava flow inundation to the present-day Catania  
71 urban area. The low frequency of lava flow invasions and favorable climatological, hydrological,  
72 and pedological conditions have allowed the development of human settlements at such  
73 peripheral areas of the volcano since the Neolithic epoch (Branca et al., 2017).

74 Less is known about the chronology of lava flows that reached the Catania area in  
75 prehistoric times. Although several studies were carried out on the Holocene Etna lavas exposed  
76 within the Catania urban district (Monaco et al., 2000; Tanguy et al., 2003, 2007, 2012; Speranza  
77 et al., 2006; Branca et al., 2011b, 2016), dating of volcanics from the town and its surroundings  
78 is limited. The aim of our work is to reconstruct the age of the prehistoric lavas on which the city  
79 of Catania is built by means of paleomagnetism using a dating/correlation tool for volcanic  
80 products of the last 14,000 years that has been used increasingly during the last 30 years to study  
81 volcanos in Italy and worldwide (Jurado-Chichay et al., 1996; Gonzalez et al., 1997; Zanella,  
82 1998 [\[\[Zanella, 1998 is not in the reference list.\]\]](#); Tanguy et al., 2003, 2007; Speranza et al.,  
83 2006, 2008; Pavón-Carrasco and Villasante-Marcos, 2010; Di Chiara et al., 2014; Greve and  
84 Turner, 2017; Pinton et al., 2018; Branca et al., 2019; Risica et al., 2019, 2020, among many  
85 others). To corroborate our dating results, paleomagnetic results were compared with geological,  
86 geochemical, and archaeological evidence to properly define the Holocene eruptive history of the  
87 Catania urban area in relation to the human settlements that have developed since prehistoric  
88 times.

## 89 GEOLOGICAL SETTING OF THE CATANIA AREA

90 The geological map of Etna volcano by Branca et al. (2011b), the third in the history of  
91 Etna's geological cartography (Branca and Abate, 2019), along with the volcanic evolution

92 summarized by Branca et al. (2011a), documents an updated and accurate reconstruction of the  
 93 **stratigraphy of the volcano**. The last 15 ka were characterized by the activity of the “Mongibello  
 94 **volcano**,” the most recent lithostomatic unit of Etna, **which** represents the currently active center  
 95 of volcanic activity (Il Piano Synthem, **stratovolcano** phase of Branca et al., 2011a). The  
 96 Mongibello **volcano** generated scoria cones, pyroclastic fallout deposits, and lava flows  
 97 extending onto over 85% of the Etna surface. Two principal tephra layers, largely spread over  
 98 the eastern and southeastern sectors of the volcanic edifice, were used to stratigraphically  
 99 constrain the ages of **the** Mongibello lava flows. **The older flow** is correlated to a  $3930 \pm 60$   $^{14}\text{C}$   
 100 yr B.P. sub-plinian eruption (FS **[[Provide words for FS rather than using acronym.]]**  
 101 pyroclastic fall deposit in Coltelli et al., 2000), whereas the younger is associated **with** the 122  
 102 B.C. (2072 yr B.P.) basaltic plinian eruption (FG **[[Provide words for FG rather than using**  
 103 **acronym.]]** tephra layer in Coltelli et al., 1998, 2000).

104 The stratigraphic sequence of the lower SE flank of Etna starts with marine marly clays  
 105 evolving to littoral sands and continental polygenetic conglomerates; **this succession represents**  
 106 the Pleistocene regressive phase (Fig. 1). The oldest Etna volcanics exposed at Catania are lava  
 107 flows that belong to the Timpe **phase** (Branca et al., 2011a, 2011b) **and** lie on the **early–middle**  
 108 Pleistocene marly clays **and crop** out discontinuously along the northern **suburbs of the town**.  
 109 Radiometric (Ar/Ar) dating of such lavas yielded **a result of** ca. 130 ka (De Beni et al., 2011).  
 110 However, the volcanics exposed over the lower SE Etna flank are predominantly represented by  
 111 younger lava flows emplaced between 15 ka B.P. and 3.9 ka B.P. (lower member of **the**  
 112 Pietracannone Formation of Branca et al., 2011b) and between 3.9 ka B.P. and the 122 B.C.  
 113 (2072 yr B.P.) tephra layer (upper member of **the** Pietracannone Formation of Branca et al.,  
 114 2011b). During the last 2 ka, several historical lava flows emplaced in the area between the  
 115 towns of Nicolosi, Trecastagni, San Giovanni La Punta, and Mascalucia (Fig. 1). In particular, in  
 116 the late Roman epoch, the Monpeloso, S.G. La Punta, and Piazza Sant’Alfio lava flows (**mp, sq,**  
 117 **and io** in Fig. 1) erupted at around A.D. 300 and A.D. 450, respectively. Considering the A.D.  
 118  $300 \pm 100$  paleomagnetic age by Tanguy et al. (2012), the Monpeloso lava flow could be  
 119 associated with the A.D. 252 eruption quoted in the historical sources. Among all **of the lava**  
 120 **flows of the Middle Ages**, only the Mt. Arsi di Santa Maria flow (**sm** in Fig. 1) reached the  
 121 Ionian coast at **the** Ognina locality (~2.3 km NE of the Medieval town of Catania), **and this**  
 122 **occurred** during the **twelfth** century A.D. During **later** centuries, the A.D. 1408 lava flow caused  
 123 considerable damage to cultivated lands and to the village of Pedara, and the A.D. 1537 lava  
 124 flow damaged the Nicolosi and Monpilieri villages. The best known eruptive event of the  
 125 historical period occurred in A.D. 1669 **and affected** the lower SE flank of Etna. The 17-km-long  
 126 A.D. 1669 lava flow covered  $\sim 40$  km<sup>2</sup> of a highly **urbanized** and agriculturally productive area,  
 127 destroying several towns and the SW part of Catania itself (Branca et al., 2013, 2015b).

## 128 **STRATIGRAPHIC AND CHRONOLOGICAL CONSTRAINTS OF THE** 129 **PREHISTORIC CATANIA LAVA FLOWS**

130 Considered as a natural laboratory for volcanologists and geophysicists, **Mt.** Etna has  
 131 been one of the most studied volcanoes in the world since the **eighteenth** century (Branca and  
 132 Abate, 2019). Thanks to the presence of inhabited zones on its flanks for some 2700 yr B.P., the  
 133 eruptive events of the volcano were occasionally reported by historical documents that date back  
 134 to the Greek colonization (Tanguy, 1981; Guidoboni et al., 2014). Some of these documents  
 135 described the effects of the eruptions on the **city of Catania**, **which was begun** as a Greek colony  
 136 in 729/728 B.C. (2679/2678 yr B.P.) with the name of Katánē (Privitera, 2010). **Thus, the**  
 137 availability of historical sources **documenting** numerous eruptive events and subsequent

138 interpretations by different historians led to many errors and false ages (Guidoboni et al., 2014;  
139 Tanguy et al., 2012; Branca and Abate, 2019). It has been demonstrated **that by integrating**  
140 geological, historical, radiometric, and paleomagnetic **analyses it** is possible to adequately  
141 resolve doubts about the **ages** of debated volcanic products (Tanguy, 1969, 1980; Tanguy et al.,  
142 1985, 2003, 2007, 2012; Condomines and Tanguy, 1995; Condomines et al., 2005; Speranza et  
143 al., 2006; Branca and Vigliotti, 2015; Branca et al., 2015a, 2016).

144 Geological investigations performed by Sartorius von Waltershausen (1843–1861),  
145 author of the first geological map of Mt. Etna, helped **to first** link historical accounts to several  
146 lava flows produced by flank eruptions (Sartorius von Waltershausen, 1843). In particular, **von**  
147 **Waltershausen** deemed the information on the eruptive phenomena described in the Greek-  
148 Roman and **early** Middle Age sources too general and difficult to interpret; consequently, the  
149 oldest lava flow the author was able to robustly date was related to the year A.D. 1285 (Branca  
150 and Abate, 2019). **Later** authors devoted themselves to the study and interpretation of Etna lava  
151 flows of **the** Catania area. In the earliest geological map of the city, **which was** published at a  
152 scale of 1:21,276, Sciuto Patti (1872) **erroneously attributed** some lava flows that reached the  
153 town to a Roman age (122 B.C., i.e., 2072 yr B.P. and A.D. 253). Such incorrect age attributions  
154 conditioned **later** age interpretations that were made in the geological maps of the **twentieth**  
155 century even **though** just a few years after the publication of **Patti's** map archaeological  
156 investigations evidenced the presence of **prehistoric** settlements in the lava **flows that Patti had**  
157 **attributed** to the Roman age (for detail see Branca et al., 2016; Branca and Abate, 2019). As a  
158 consequence, in the second geological map of Etna at **the** 1:50,000 scale by Romano et al.  
159 (1979), several Greek-Roman ages (693 B.C., i.e., 2643 yr B.P.; 425 B.C., i.e., 2375 yr B.P.; 122  
160 B.C., i.e., 2072 yr B.P.; and A.D. 252–253) were erroneously assigned to some lava fields  
161 mapped in the urban area of Catania **due to** the interpretation of both the historical catalogues of  
162 Etna's eruptions and the previous geological maps (Branca and Abate, 2019). Similarly, in the  
163 recent geological map of Catania published at **the** 1:10,000 scale by Monaco et al. (1999, 2000),  
164 the authors **adopted Patti's incorrect A.D. 252 age attribution for an exposed lava flow.**

165 In the most recent 1:50,000 scale geological map of Mt. Etna, Branca et al. (2011b),  
166 **using** an interdisciplinary approach including stratigraphy, revised historic analysis, and  
167 radiometric dating of the lavas, **showed** that 85% of lavas attributed to the Greek-Roman times  
168 until the **sixteenth** century A.D. are in **fact** either prehistoric or several centuries older. In  
169 particular, according to Branca et al. (2011b), the urban area of Catania is formed by two main  
170 lava flow fields of prehistoric age named Barriera del Bosco and Larmisi (**bb** and **la** in Fig. 1,  
171 respectively, Pietracannone Formation lower member, stratigraphic age between 15 ka B.P. and  
172 3.9 ka B.P.) that are characterized **by several archaeological artifacts from the Neolithic to the**  
173 **Greek and Roman ages lying above** (Branca et al., 2016). A restricted portion of the coast at **the**  
174 Ognina locality is formed by a lava flow, named Ognina (**og** in Fig. 1), **which was** emplaced in  
175 the same stratigraphic age interval (15–3.9 ka BP) **as** the Barriera del Bosco and Larmisi flows.  
176 An additional prehistoric lava flow, named San Giovanni Galermo (**le** in Fig. 1), outcrops in the  
177 **northwestern** sector of Catania. **Early** Bronze age archaeological artifacts **were found** within **its**  
178 lava tubes, **and the flow** was dated by  $^{226}\text{Ra}$ - $^{230}\text{Th}$  technique at 2700 B.C. (+960/–750), i.e.,  
179 5610–3900 yr B.P. (Sample 090 of Tanguy et al., 2007).

180 In this work we focused on the Barriera del Bosco, Larmisi, and San Giovanni Galermo  
181 prehistoric lava flows (available age constraints are summarized in Table 1). The lower age  
182 boundary of the Barriera del Bosco flow is poorly known; however, the lava flow is considered  
183 **to be** part of the Pietracannone **Formation**, whose lower bound **is dated at the AMS-constrained**

184 ages of  $15,420 \pm 60$  yr B.P. and  $15,050 \pm 70$  yr B.P. (D1a and D2a pumice layers, respectively,  
 185 in Coltelli et al., 2000). The Middle Neolithic artifacts (6950–5950 yr B.P.; Privitera and La  
 186 Rosa, 2007; Branciforti, 2010; Frasca, 2015; Nicoletti, 2015) found below the Benedettini  
 187 Monastery, which lies above the Barriera del Bosco flow, provide the upper bound flow age.

188 The Larmisi flow frontal portion forms a 2-km-long and 10–15-km-high cliff (“Larmisi  
 189 cliff”) that lies over a depositional marine platform located at 5 m below sea level (SG **[[Provide  
 190 full term for SG rather than abbreviation.]]** layer in Monaco et al., 2000). Considering the  
 191 relative sea level change curve of Lambeck et al. (2004), the sea was ~6 m below the present  
 192 level in prehistoric times (8500–6500 yr B.P.). Modifying the Lambeck curve for a regional  
 193 Holocene uplift rate of 1.2–1.4 mm/yr for eastern Sicily, the marine platform is inferred to have a  
 194 7500–7000 yr B.P. age (C. Monaco, 2020, personal commun.), which implies that the Larmisi  
 195 flow is younger than 7500–7000 yr B.P. Several Late Copper Age artifacts (4550–4250 yr B.P.;  
 196 Privitera and La Rosa, 2007) found inside some lava tubes distributed along this lava flow can be  
 197 considered to represent the upper bound age of the Larmisi flow.

198 The San Giovanni Galermo flow had been previously dated at 5610–3900 yr B.P. by  
 199  $^{226}\text{Ra}$ - $^{230}\text{Th}$  dating (site 090 of Tanguy et al., 2007). In addition, the presence of pottery dating  
 200 back to the Early Bronze Age (3950–3350 yr B.P.; Privitera and La Rosa, 2007) confirms the  
 201 radiometric upper bound age.

202 To sum up, available geological, archaeological, and radiometric evidence constrains the  
 203 Barriera del Bosco, Larmisi, and San Giovanni Galermo flows to the 13,950–6950 yr B.P.,  
 204 7500–4550 yr B.P., and 5610–3950 yr B.P. age ranges, respectively.

## 205 PRINCIPLES OF PALEOMAGNETIC DATING

206 Paleomagnetic dating is based on the fundamental assumption that remanent  
 207 magnetization, acquired by volcanic rocks at the time of their emplacement and cooling, is  
 208 parallel to the Earth’s local magnetic field direction (e.g., Butler, 1992). Thus, comparing  
 209 paleomagnetic directions recorded by volcanics and independently available reference curves  
 210 showing the paleo-secular variation (PSV) swings of the geomagnetic field direction (some tens  
 211 of degrees in declination and inclination at Sicily coordinates, Bucur, 1994; Gallet et al., 2002;  
 212 Tanguy et al., 2003) during the last millennia can provide one or more emplacement age ranges  
 213 within a given input time window. PSV reference curves were obtained by archeomagnetism,  
 214 paleomagnetism of well-dated volcanics, and paleomagnetism of cores drilled within lacustrine  
 215 successions deposited at a high sedimentation rate. Directional PSV data have regional validity  
 216 and are abundant in Europe, where a wealth of reference data was gathered in past decades (e.g.,  
 217 Speranza et al., 2008). Thus, the accuracy of paleomagnetic dating is strictly dependent upon the  
 218 availability of neighbor reference PSV data and reaches maximum reliability and potentiality for  
 219 European volcanoes.

220 In the last few years, paleomagnetic dating was routinely done using the SHA.DIF.14k  
 221 global PSV reference model (Pavón-Carrasco et al., 2014) that extends back to the last 14 ka and  
 222 discards reference sedimentary data, which is a possible source of bias. As the last 14 ka is  
 223 virtually the same time range in which all three of the lava flows studied fall (Barriera del Bosco,  
 224 Larmisi, and San Giovanni Galermo), the SHA.DIF.14k model was chosen for this work.

225 In the last three decades, paleomagnetic dating has been applied increasingly to date  
 226 volcanic products worldwide in Iceland (Thompson and Turner, 1985; Pinton et al., 2018), the  
 227 Canary Islands (Soler et al., 1984; Pavón-Carrasco and Villasante-Marcos, 2010; Risica et al.,  
 228 2020), the Azores (Di Chiara et al., 2012, 2014), New Zealand (Cox, 1969; McClelland et al.,  
 229 2004; Greve and Turner, 2017), Hawaii (Doell and Cox, 1963; Holcomb et al., 1986; Jurado-

230 Chichay et al., 1996), and Mexico (Gonzalez et al., 1997; Böhnel et al., 2016; Mahgoub et al.,  
 231 2017). Eruptive products from famous Italian volcanoes were paleomagnetically investigated:  
 232 Vesuvius (Hoye, 1981; Incoronato et al., 2002; Tanguy et al., 2003; Principe et al., 2004),  
 233 Stromboli (Speranza et al., 2004, 2008; Arrighi et al., 2004[[Arrighi et al., 2004 is not in the  
 234 reference list. 2006 here?]]; Risica et al., 2019), Vulcano (Zanella and Lanza, 1994; Arrighi et  
 235 al., 2006), and Pantelleria (Speranza et al., 2010, 2012). On Etna, several flank eruptions of the  
 236 past 2400 years were paleomagnetically dated (Chevallier, 1925; Rolph and Shaw, 1986;  
 237 Incoronato et al., 2002; Speranza et al., 2006; Tanguy et al., 1985, 2003, 2007, 2012; Branca et  
 238 al., 2011a, 2015a, 2019).

### 239 SAMPLING AND METHODS

240 In June 2019, we paleomagnetically sampled four sites from each of the three prehistoric  
 241 Etna lava flows (Barriera del Bosco, San Giovanni Galermo, and Larmisi) exposed in the Catania  
 242 urban district for a total of 12 sites (Fig. 1 and Table 1). We carefully selected the sampling  
 243 outcrops to investigate several single flow units from the same lava field and focused on massive  
 244 cores of lava flow units, where post-emplacement tilt, i.e., the main bias source of paleomagnetic  
 245 dating (e.g., Pinton et al., 2018), is unlikely to occur.

246 Sites ETN23, 24, 25, and 26 were located at the Barriera del Bosco flow. Collection at  
 247 site ETN23 was conducted in a single lava flow unit along Viale della Regione just north of the  
 248 cemetery (Fig. 2A); at the same site, sandy sediments (likely the San Giorgio sands of Branca et  
 249 al., 2011a) were burnt by the underlying hot lava flow through the degassing process and form  
 250 the so-called “lithophysae” (Fig. 2B). Site ETN24 was sampled near San Severio Villa (Cibali  
 251 neighborhood) at a small outcrop along the street floor; site ETN25 was sampled near Guglielmo  
 252 Oberdan Street in a lava flow unit outcrop; site ETN26 was sampled along Santa Maria La  
 253 Grande Street, where a massive core of basaltic lava characterized by joints is exposed.

254 Sites ETN27, 28, 29, and 33 are located in the San Giovanni Galermo flow. Site ETN27  
 255 was sampled at the crossroads of Santa Sofia Street and Fratelli Vivaldi Street from a single lava  
 256 flow unit; collection at sites ETN28 and ETN29 was conducted along Sebastiano Catania Street  
 257 and Via della Misericordia Street, respectively; site ETN33 was sampled at Egadi Street (San  
 258 Nullo locality) from the massive basaltic lava core outcrops of three overlapping flow units (Fig.  
 259 2D) characterized by a ropy surface on their roofs (Fig. 2E).

260 Finally, four sites were sampled in the Larmisi lava flow: site ETN30 comprises several  
 261 overlapping lava flow units along Don Luigi Sturzo Street (Fig. 2G–2H); site ETN31 was  
 262 sampled along Via dei Miti Street, where three lobes with alternating massive cores and welded  
 263 autobreccias crop out; site ETN32 in Gioeni Park lies in a wide lava field consisting of several  
 264 flow units; and collection at site ETN34 sampled some lava lobes exposed near Justice Palace  
 265 along Ramondetta Street.

266 Samples were collected by drilling 2.5-cm-diameter cores with a gas-powered portable  
 267 drill cooled by water. At each site we systematically drilled 10 cores for a total number of 120  
 268 cores. All cores collected were spaced along the outcrop over a length of few meters to tens of  
 269 meters and oriented in situ using both a magnetic compass and a sun compass. The comparison  
 270 between sun and magnetic compass readings translates into local magnetic field declinations  
 271 varying from 24° to –35° (4.8° on average), which compares to the regional geomagnetic field  
 272 declination at Etna ( $D=3^\circ$ ) extrapolated for June 2019 from the International Geomagnetic  
 273 Reference Field model  
 274 (<https://www.ngdc.noaa.gov/geomag/calculators/magcalc.shtml#igrfwmm>; accessed April 2021).

275 Cores were cut into standard cylindrical specimens, and the **natural remanent**  
276 **magnetization** (NRM) of one specimen per core was measured in the shielded room of the  
277 Paleomagnetic Laboratory of Istituto Nazionale di Geofisica e Vulcanologia in Rome using a 2G  
278 Enterprises **direct current superconducting quantum interference device** cryogenic  
279 magnetometer. All specimens were demagnetized by alternating field (AF) cleaning with 10  
280 steps until a maximum AF peak of 120 mT **was reached**.

281 AF demagnetization data were plotted on orthogonal demagnetization diagrams  
282 (Zijderveld, 1967) and equal **area** projections, and magnetization components were isolated  
283 **through** principal component analysis (Kirschvink, 1980). Site-mean and lava flow-mean  
284 paleomagnetic directions (calculated **by** averaging **all individual characteristic remanent**  
285 **magnetization [ChRM]** direction components from the same flow) were computed using Fisher's  
286 (1953) statistics.

287 Paleomagnetic dating was **performed using** the Matlab tool developed by Pavón-Carrasco  
288 et al. (2011). The SHA.DIF.14k global model used in this work (Pavón-Carrasco et al., 2014)  
289 spans the 12,000 B.C.–A.D. 1900 (13,950–50 yr B.P.) period and relies on archaeomagnetic and  
290 well-dated **volcanic** paleomagnetic data from the GEOMAGIA50v2 data set (Donadini et al.,  
291 2006; Korhonen et al., 2008).

292 For two specimens from each lava flow (**a** total of six specimens), we also measured the  
293 variation of the low-field magnetic susceptibility during a heating and cooling cycle performed  
294 in air **at a** room temperature **of** up to 700 °C using an AGICO MK1-FA Kappabridge coupled  
295 with a CS-3 furnace. The Curie **temperature** ( $T_c$ ) of the magnetic minerals present in the samples  
296 was determined from the thermomagnetic curves (Fig. S1<sup>1</sup>) as the temperature, or range of  
297 temperatures, at which paramagnetic behavior starts to dominate, following the approach  
298 outlined by Petrovský and Kapička (2006).

299 For one specimen per site, hysteresis loops were also measured (Fig. S2; see footnote 1)  
300 using a Princeton Measurement Corporation MicroMag alternating gradient magnetometer  
301 (Model 2900) with a maximum applied field of 1 T. The acquired hysteresis parameters are  
302 saturation magnetization ( $M_s$ ), saturation remanent magnetization ( $M_{rs}$ ), and coercive force ( $B_c$ ).  
303 Coercivity remanence parameter  $B_{cr}$  was measured by **acquiring** an isothermal remanent  
304 magnetization (IRM) and subsequent back-field DC remagnetization (both in a succession of  
305 fields up to 1 T).

306 Rock blocks **at** some paleomagnetic sites were also gathered for petrographic analyses:  
307 **11** samples from Barriera del Bosco flow (**sites** ETN23, ETN24, ETN25, and ETN26), one from  
308 San Giovanni Galermo flow (**site** ETN27), and three from Larmisi flow (**sites** ETN30 and  
309 ETN31). Thirty- $\mu$ m-thick rock sections were obtained from each block **at** the Thin Section  
310 Laboratory of the Roma Tre University.

## 311 **RESULTS**

### 312 **Petrographic Analyses**

313 For **all three lava flows studied**, petrographic analyses show porphyritic and phaneritic  
314 texture with large phenocrysts (2–7 mm) inside a hypocrySTALLINE groundmass (Figs. **2C–2F and**  
315 **2I** **[[OK?]]** and Fig. S3; see footnote 1). The predominant phenocrysts are plagioclases (Pl) with  
316 minor clinopyroxenes (Cpx), olivines (Ol), and opaque oxides (Ox), which **comprise** the  
317 common mineralogical assemblage of Etna products (Corsaro and Pompilio, 2004). The Barriera  
318 del Bosco lava flow (Fig. 2C and Fig. S3) presents a **porphyricity index** (P.I. = total **phenocryst**  
319 abundance expressed in volume %) of around 30% with elongated plagioclases (labradorite, 3–6  
320 mm) constituting 80%–90% and euhedral clinopyroxenes/olivines representing 10%–20% (2–4

321 mm). The Pl/Cpx+Ol ratio ranges from 9 to 4. The groundmass, which forms 70% of the rock  
 322 volume, is characterized by an intersertal texture in which plagioclase microlites form a network  
 323 whose intergranular spaces are filled with mafic minerals and/or volcanic glass. Groundmass  
 324 microlites have the same paragenesis of phenocrysts.

325 In the San Giovanni Galermo lava flow (Fig. 2F), mafic minerals are more abundant,  
 326 with a Pl/Cpx+Ol+Ox ratio of 1.5–1.9. The P.I. is 40%–45%, which is greater than that of the  
 327 Barriera del Bosco flow, with plagioclase being the most abundant phenocrysts (60%–65%,  
 328 andesine, 3–5 mm) followed by euhedral clinopyroxenes (20%–22%, augite to aegirine-augite,  
 329 3–4 mm), and sub-rounded olivines (15%–18%, 1–3 mm). Pyroxenes are often arranged as  
 330 glomeroporphyritic aggregates with phenocrysts bracketed in groups. The groundmass (55%–  
 331 60% of the total volume) presents an intersertal structure with paragenesis of microlites (<0.5  
 332 mm) similar to that of phenocrysts (60% plagioclase; 30% clinopyroxene; 10% olivine).

333 The Larmisi lava flow (Fig. 2I) has Pl/Cpx+Ol+Ox ratios of between 1 and 1.9 and a P.I.  
 334 of 40–45%, which fully overlap those of San Giovanni Galermo. Elongated plagioclases are  
 335 always the predominant phenocrysts (60%–65%, andesine, 2–6 mm), followed by euhedral  
 336 clinopyroxenes (30%–20%, augite to aegirine-augite, 1–5 mm, sometimes in glomeroporphyritic  
 337 aggregates) and sub-rounded olivines (10%–20%, 0.5–3 mm). The groundmass forms 55–60% of  
 338 the rock and shows an intersertal texture with paragenesis of microlites (<0.5 mm) similar to that  
 339 of phenocrysts (50%–60% plagioclase; 30% clinopyroxene; 10%–20% olivine).

#### 340 **Magnetic Properties**

341 Almost all thermomagnetic curves show an irreversible variation trend during heating-  
 342 cooling cycles, which indicates magnetic mineralogy changes during heating (Fig. S1). Samples  
 343 ETN2301 and ETN3007 show similar heating-cooling cycles and  $T_c$  of 580 °C, which is  
 344 characteristic of magnetite. In the remaining samples, multiple  $T_c$ s are apparent in the 150–500  
 345 °C temperature range likely due to the occurrence of minerals belonging to the titanomagnetite  
 346 series. Moreover, noticeable magnetization above the magnetite  $T_c$  in samples ETN2301,  
 347 ETN3007, ETN3203, and ETN3310 may be associated with the occurrence of hematite,  
 348 contributing to magnetic susceptibility to a lesser degree. To sum up, thermomagnetic analyses  
 349 show that magnetite and Ti-magnetites are the predominant magnetic minerals of the lava flows  
 350 investigated.

351 All samples analyzed show low coercive force values ( $B_c$ ) ranging between 4.6 mT and  
 352 26 mT. The great variation in magnetization and coercivity parameters suggests a wide range of  
 353 magnetic behaviors. Some specimens (ETN2304, ETN2504, ETN3007, and ETN3107) show  
 354 hysteresis cycles with a noticeably wide shape related to single-domain (SD) grains (Fig. S2A),  
 355 while other samples (ETN2604 and ETN2810) have narrower hysteresis loops that are typical of  
 356 multi-domain (MD) grains (Fig. S2B). Remaining specimens display intermediate magnetic  
 357 behaviors between these two end members (Figs. S2C–S2D). Hysteresis parameters were plotted  
 358 in a Day plot (Day et al., 1977; Dunlop, 2002a, 2002b; Fig. S2E) of saturation remanence to  
 359 saturation magnetization ( $M_{rs} / M_s$ ) against the ratio of remanent coercive force to coercive force  
 360 ( $B_{cr} / B_c$ ). All of the samples follow the theoretical mixing curves for SD and MD magnetite  
 361 ( $Fe_3O_4$ ), with the exception of a single specimen (ETN3405) lying on a TM60 titanomagnetite  
 362 (i.e.,  $Fe_{3-x}Ti_xO_4$ , with  $x = 0.6$ ) mixing curve.

#### 363 **Paleomagnetic Directions**

364 In all AF-cleaned specimens, a well-defined ChRM was isolated in the 20–120 mT field  
 365 interval (Fig. 3). More than 90% of magnetic remanence was removed at 100 mT, thus pointing  
 366 to low-coercivity minerals as main magnetic carriers. However, for ~10% of the samples (mostly



367 from sites ETN25, ETN30, and ETN31), only 75%–80% of NRM is removed at the maximum  
 368 available 120 mT alternating field (see ETN3102 specimen in Fig. 3), **which suggests** the  
 369 occurrence of both low- and high-coercivity minerals. Given the not negligible amount of  
 370 hematite documented by some thermomagnetic curves (Fig. S1), the high-coercivity fraction is  
 371 possibly represented by both hematite and SD and/or deuterically oxidized titanomagnetite (e.g.,  
 372 Dunlop and Özdemir, 2001).

373 Site-mean declinations (except **at** site ETN26) vary from  $-9.1^\circ$  (site ETN31) to  $15.0^\circ$   
 374 (ETN25), and inclinations range from  $40.4^\circ$  (ETN23) to  $60.8^\circ$  (ETN34; Fig. 4A and Table 2).  
 375 Site ETN26 **at** the Barriera del Bosco flow shows a scattered direction that **lies**  $26^\circ$ – $27^\circ$  apart  
 376 from the directions of the remaining three sites from the same flow. This site was sampled in an  
 377 isolated, 8-m-long outcrop left **along** the **urban** Via Santa Maria La Grande **Road** and thus was  
 378 most likely tilted after emplacement. Site ETN26 was thus considered an outlier and discarded  
 379 from further consideration. The  $\alpha_{95}$  values relative to the mean paleomagnetic directions  
 380 evaluated for each site range between  $2.7^\circ$  and  $7.1^\circ$  ( $5.0^\circ$  on average, Fig. 4A and Table 2),  
 381 whereas they vary from  $2.6^\circ$  to  $3.2^\circ$  **depending on the** **[[OK?]]** lava flow (average  $2.9^\circ$ , Fig. 4B  
 382 and Table 2).

## 383 DISCUSSION

### 384 Age Determinations

385 Our paleomagnetic data reveal that **the** Larmisi and San Giovanni Galermo lava flows  
 386 share overlapping paleomagnetic directions (Figs. 4A–4B); thus, the question arises **of** whether  
 387 the two lava flows **in fact belong** to the same lava field. Indeed, while two significantly different  
 388 paleomagnetic directions from two volcanic units **are definitive** proof **of** distinct emplacement  
 389 age (with a 100–200 year time resolution, e.g., Speranza et al., 2012), similar paleomagnetic  
 390 directions may either indicate coeval emplacement **or be** related to the characteristics of the  
 391 geomagnetic field that may reoccupy the same directions **a few centuries or millennia later**  
 392 (Butler, 1992).

393 Some statistical tests are used to verify whether two data sets share a common mean  
 394 direction. The most widely used is the F-test (Watson, 1956), recently adopted by Larrea et al.  
 395 (2019), which compares a statistic parameter,  $F$ , with tabulated values for the chosen significance  
 396 level. Another test, considered **to be** more statistically reliable (Tauxe et al., 2018) and used here,  
 397 is the Watson test or  $V_w$  test (Watson, 1983). In the simplest terms, if two data sets share a  
 398 common mean direction, then the statistic parameter  $V_w$  (that increases with **growing** difference  
 399 between the mean directions of the two data sets) will be lower than a critical value ( $V_{crit}$ )  
 400 determined through the Monte Carlo simulation (see Supplemental Material S4 for the detailed  
 401 theory and calculation process; see footnote 1). We applied the Watson test, using the PmagPy  
 402 function of Tauxe et al. (2016), to **the** Larmisi and San Giovanni Galermo directions to examine  
 403 if these two data sets share a common mean direction. The obtained  $V_w$  value (5.5) is smaller  
 404 than the  $V_{crit}$  (6.2) (see Supplemental Material S4 for the results); consequently, the two data sets  
 405 share a common mean direction.

406 We also note that petrographic analyses **indeed suggest similar** characteristics of the two  
 407 flows, although this is not a conclusive criterion given the similarity of most Etna lava flows  
 408 (Corsaro and Pompilio, 2004). Both lavas have a **porphyricity index** of 40%–45%, **plagioclase**  
 409 **phenocryst** content ranging from 60% (Larmisi flow) to 65% (San Galermo Galermo flow), high  
 410 amounts of mafic minerals (Pl/Cpx+Ol+Ox ratio of  $\sim 1$ –1.9), and clinopyroxenes in places  
 411 arranged as glomeroporphyritic aggregates. Moreover, it must be recalled that there is no field  
 412 evidence of stratigraphic boundaries between the two lava flows.

413 Finally, **geochemical data from the literature** for the two flows show nearly identical  
 414 chemical compositions (Corsaro and Cristofolini, 1993; Table S1; see footnote 1), **which** further  
 415 **confirms** that the two lava flows can reliably be **considered products** of the same eruption.

416 We conclude that the San Giovanni Galermo and Larmisi flows belong to the same  
 417 volcanic unit emplaced during a single eruptive event. **Thus**, we calculated a single mean  
 418 paleomagnetic direction for the San Giovanni Galermo-Larmisi lava flows (Fig. 4C) by  
 419 averaging the **76 reliable ChRMs** obtained from the two flows.

420 Eruption ages were obtained by comparing mean paleomagnetic directions from the  
 421 Barriera del Bosco and San Giovanni **Galermo-Larmisi** flows to field **direction** values expected  
 422 at Etna considering the SHA.DIF.14k global model (Fig. 5; Pavón-Carrasco et al., 2014) by  
 423 using the “archaeo\_dating” Matlab tool.

424 Three age intervals were obtained for the Barriera del Bosco flow: 11,234–10,941 yr  
 425 B.P., 8395–8236 yr B.P., and 7309–6950 yr B.P. (Table 2 and Fig. 5). Conversely, two possible  
 426 age **windows between** 5610 yr B.P. and 4550 yr B.P. (i.e., the input time window overlapping  
 427 both individual time windows from the two flows) were obtained for the San Giovanni Galermo-  
 428 Larmisi lava flow (Table 2 and Fig. 5).

429 In Figure 6, the paleomagnetic ages are compared with input time windows arising from  
 430 **archaeological/radiometric ages reported in the literature** from the Barriera del Bosco and San  
 431 Giovanni Galermo-Larmisi lava flows. Although all **of** the paleomagnetically inferred ages are  
 432 equally probable, the presence of Middle Neolithic finds (6950–5950 yr B.P.) on the Barriera del  
 433 Bosco lava flow allows **us** to rule out its youngest age interval (7309–6950 yr B.P.), which is  
 434 temporally too close to enable the communities of that period to settle there. The most likely  
 435 paleomagnetic ages of the Barriera del Bosco flow are then 11,234–10,941 yr B.P. and 8395–  
 436 8236 yr B.P. (bold contour boxes in Fig. 6). **Both age** determinations are in agreement with the  
 437 oldest human presences recognized in the Catania area **and suggest** that the Barriera del Bosco  
 438 lava flow was emplaced during the early stage of **the** Holocene epoch and was afterwards widely  
 439 covered by lacustrine and alluvial deposits (Fig. 1).

440 Concerning the San Giovanni Galermo-Larmisi lava flow, the  $^{226}\text{Ra}/^{230}\text{Th}$  age (5610–  
 441 3900 yr B.P.) of Tanguy et al. (2007) **overlaps** with both paleomagnetic age intervals.  
 442 Furthermore, following the same criterion previously adopted for the Barriera del Bosco flow,  
 443 the youngest age span (5125–4550 yr B.P.) of the San Giovanni Galermo-Larmisi lava flow is  
 444 very close to the Late Copper Age artifacts (4550–4250 yr B.P.) discovered on the lava flow  
 445 itself and **consequently should be discarded**. Accordingly, the only paleomagnetic age for **the**  
 446 San Giovanni Galermo-Larmisi flow that is consistent with both geologic and archaeological  
 447 evidence is the narrow 5494–5387 yr B.P. age window (Fig. 6).

#### 448 **Archeological Implications**

449 The new paleomagnetic datings, cross-correlated with both archaeological and geological  
 450 evidence, allow **constraints on** the age of **the** prehistoric lava flow **that was emplaced where** the  
 451 urban district of **Catania later** developed. Undistinguishable paleomagnetic directions, along with  
 452 petrographic and geochemical evidence, show that the San Giovanni Galermo and Larmisi lava  
 453 flows, **which had been considered to be** related to distinct eruptions (Branca et al., 2011b), form  
 454 in fact a unique wide lava field generated by a flank eruption **in Etna’s** SE lower slope between  
 455 5494 yr B.P. and 5387 yr B.P. (Fig. 7). **This** wide lava field **almost entirely covered** the Barriera  
 456 del Bosco flow, whose paleomagnetic age ranges from 11,234–8236 yr B.P.

457 In addition, the new datings of the lavas lying below the metropolitan district of **the town**  
 458 **of Catania** explain the scarce distribution of Neolithic archaeological sites found in this area of

459 the volcano. **Archaeological sites here are** limited to the Mt. Vergine hill, where the Benedettini  
460 Monastery is **located** (Fig. 7). **At** this site, the human colonization of the Barriera del Bosco lava  
461 flow started **in** the Middle Neolithic (Branciforti, 2010; Nicoletti, 2015). The emplacement of the  
462 wide San Giovanni Galermo-Larmisi lava flow radically modified the morphological setting of  
463 the Late Neolithic landscape **and probably masked** further evidence of the human presence in this  
464 sector of the Etnean coast during the Neolithic period. In fact, the oldest traces of human  
465 colonization on the San Giovanni Galermo-Larmisi lava flow date back to the Upper Copper and  
466 Early Bronze periods (Fig. 7). This new morphological and geological setting represented the  
467 substratum of the following human settlements until the foundation of the Greek colony of  
468 Katánē in 729/728 B.C. (2679/2678 yr B.P.). **Since** the **founding** of the town 2700 years ago, the  
469 interaction between **volcanic events and urban** development **has been** very limited **[[OK?]]**  
470 (Branca et al., 2016). In fact, during the Greek-Roman domination (from **the eighth** century B.C.  
471 to the **fifth** century A.D.), only the 122 B.C. (2072 yr B.P.) Plinian eruption produced significant  
472 pyroclastic fallout that severely damaged the town (Coltelli et al., 1998; Guidoboni et al., 2014).  
473 After a **millennium passed without** any eruptive event, in the **twelfth** century A.D. the Mt. Arsi di  
474 Santa Maria lava flow reached the Ionian coast ~3.5 km **north** of the Medieval town (Fig. 7).  
475 Finally, in A.D. 1669, a lower flank eruption generated a 17-km-long lava flow that impacted  
476 **Catania** (Branca et al., 2015b, 2016).

#### 477 **CONCLUSION**

478 The urban area of Catania, with its 330,000 citizens, **is** the main inhabited zone around  
479 the Etna volcano. Our new paleomagnetic data, coupled with geological and petrographic  
480 evidence from the Holocene lava flows emplaced within the present metropolitan district of  
481 Catania, define their ages and **highlight** the eruption timing in relation to the human presence in  
482 **the prehistoric** epoch. Moreover, the new paleomagnetic datings **have significantly improved**  
483 **knowledge of** the geochronology of Holocene lava flows exposed within the Catania urban  
484 district.

485 The oldest archaeological evidence of human presence in this area **dates** back to **the**  
486 Middle Neolithic ages (6950–5950 yr B.P.) **and** lies on the lava field known as Barriera del  
487 Bosco, which yields a paleomagnetic age range **of** between 11,234 and 8236 yr B.P. Above this  
488 lava flow we have correlated to the same eruptive event two lava flows **that are** interpreted so far  
489 as distinct, **which are** known as **the** San Giovanni Galermo and Larmisi flows. When considered  
490 together, the two flows are paleomagnetically dated to a narrow age **range between 5494 yr B.P.**  
491 **and 5387 yr B.P.** The emplacement of such **a** wide lava field almost covered the Barriera del  
492 Bosco flow and highly impacted the Catania area; **the flow reached** the Ionian Sea. Even though  
493 the San Giovanni Galermo-Larmisi lava flow caused significant changes in the morphological  
494 setting of this sector **of Etna's** SE coast during the late Neolithic age, **continuous** development of  
495 the human presence is **evidenced** by the Early Bronze age (3950–3350 yr B.P.) archaeological  
496 artifacts discovered in the lava tubes of this flow. **A few** centuries later, the Chalcidian (Greek)  
497 settlers founded the town of Katánē on the Barriera del Bosco lava flow in 729/728 B.C./2679 yr  
498 B.P.

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- 821 Figure 1. Geological map **shows** the lower southeastern flank of **Mt.** Etna volcano (modified  
 822 from Branca et al., 2011b, 2016). The **locations** of the paleomagnetic sampling sites **are** also  
 823 shown: **Barriera del Bosco (circles)**, **Larmisi (triangles)**, and **San Giovanni Galermo (asterisks)**.
- 824 Figure 2. Representative photographs **show** the outcrops, lithologic details, and thin sections  
 825 (cross-polarized light) **of** each lava flow. (A–C) Barriera del Bosco flow images show: (A) a  
 826 flow unit with its massive core and upper autoclastic breccias; (B) particular of sedimentary

827 deposit (contoured by a white line) burned by underlying unit through “lithophysae” degassing  
828 process; (C) thin section with a detail of the porphyritic texture (note the abundance of  
829 plagioclase phenocrysts). (D–F) San Giovanni Galermo flow photographs show: (D) view of a  
830 flow outcrop; (E) particular of lava flow roof characterized by ropy surface; (F) thin section with  
831 a porphyritic texture and plagioclase (Pl), clinopyroxenes (Cpx), and olivine (Ol) phenocrysts.  
832 (G–I) Larmisi flow photographs represent: (G) a 3-m-high outcrop characterized by several  
833 stacked flow units with alternating massive cores and autobreccias; (H) detail showing the high  
834 vesicularity of the lava flow (a 2.5-cm-diameter paleomagnetic drill hole is visible); (I) thin  
835 section showing a porphyritic texture with plagioclase, clinopyroxene, and olivine phenocrysts.  
836 Figure 3. Orthogonal vector diagrams show representative alternating field demagnetization data  
837 and in situ coordinates. Black and white dots indicate projections of paleomagnetic directions  
838 onto the horizontal and vertical planes, respectively. **NRM—natural remanent magnetization.**  
839 Figure 4. (A) Equal-area projections (lower hemisphere) of site-mean paleomagnetic directions  
840 from the three lava flows are shown. Ellipses around the paleomagnetic directions are the  
841 projections of the relative  $\alpha_{95}$  cones. (B–C) Projection of flow-mean paleomagnetic directions  
842 obtained by averaging (B) all characteristic remanent magnetizations (ChRMs) from each lava  
843 flow and then (C) considering the San Giovanni Galermo and Larmisi flows together. Ellipses  
844 around the paleomagnetic directions are the projections of the relative  $\alpha_{95}$  cones. All  
845 paleomagnetic directions are listed in Table 2.  
846 Figure 5. Paleomagnetic dating results of the Barriera del Bosco and San Giovanni Galermo-  
847 Larmisi flows are shown according to the Pavón-Carrasco et al. (2011) method and software and  
848 the Pavón-Carrasco et al. (2014) paleo-secular variation (PSV) reference model. In each panel,  
849 PSV curves for the declination (left-hand panel) and the inclination (right-hand panel) are shown  
850 as thick light gray lines (thin light gray lines indicate associated errors at the 95% confidence  
851 level) with the probability density curves (in gray-shade below each PSV). Paleomagnetic  
852 declination and inclination values are shown in the PSV graphs as dark gray straight lines; the  
853 light gray straight lines above and below are the 95% associated errors. In the probability density  
854 graphs the 95% confidence level is shown as a light gray straight line. Ages are in yr B.P.  
855 (present is A.D. 1950).  
856 Figure 6. Overall framework of the paleomagnetically inferred ages for the Barriera del Bosco,  
857 Larmisi, San Giovanni Galermo, and San Giovanni Galermo-Larmisi flows (more likely age  
858 intervals are shown with bold countour boxes; see text for details). Dashed boxes indicate the  
859 lava flows’ input time windows. Arrows represent the archaeologically constrained ages. The  
860 two dark gray straight lines are the lower and upper age constraints shared by the San Giovanni  
861 Galermo and Larmisi flows, respectively, which were used to date the new unified San Giovanni  
862 Galermo-Larmisi lava flow. Ages are in yr B.P. (present is A.D. 1950).  
863 Figure 7. Updated geological map of the Catania urban district (modified from Branca et al.,  
864 2011b) considers the new paleomagnetic data and includes the locations of archaeological sites  
865 recognized within the lava flows (modified from Branca et al., 2016).