

36

37 **Keywords:** *Speleoseismology; Central Apennines; Seismic Hazard; Finite Element Modelling*

38

39 **1. Introduction**

40

41 Cave speleothems can be used not only for paleoenvironmental and paleoclimatic reconstructions
42 (Gascoyne et al., 1980; Li et al., 1989; Talma & Vogel, 1992; Lauritzen, 1995; Bar-Matthews et al., 1999;
43 McDermott, 2004), but also for speleoseismological research, i.e. the investigation of traces of ancient
44 earthquakes in caves (Kagan et al., 2005; Becker et al., 2006). Since the 80's researchers (e.g. Forti &
45 Postpischl, 1980; Forti et al., 1981; Forti & Postpischl, 1984; Postpischl et al., 1991, Forti, 2001, Rajendran et
46 al., 2016) established that in the proved absence of other deformation sources (landslide, flooding, ice flow,
47 animal or anthropic passage), the deviation of the growth axis from the vertical and the speleothem
48 collapse are evidence for seismic shaking of cave walls. Improvements in analytical techniques and
49 theoretical modelling also showed that in earthquake-prone regions, unbroken speleothems that have
50 survived all the earthquakes occurred during their life span may provide constraints on seismic hazard
51 estimates, allowing to extend the window of observation of a given historical catalogue (Di Domenica &
52 Pizzi, 2017; Gribovski et al., 2013, 2018; Ferranti et al., 2019).

53 In terms of seismic hazard, the main issue concerns the evaluation of the ground motion potentially
54 responsible (or not) of the speleothems collapse. Cadorin et al. (2001) carried out laboratory
55 measurements on stalagmite samples to study their failure in bending and estimate the ground
56 acceleration necessary to break speleothems. Similarly, Lacave et al. (2000, 2004) investigated the
57 mechanical behaviour of speleothems via static bending tests and estimated the probability that at least
58 one moderate earthquake has occurred in the past. Recently, Mendecki and Szczygieł (2019) provided
59 quantitative relationships between speleothems shape and their oscillation and acceleration during
60 shaking. In these studies, the underlying idea is based on the stalactite-stalagmite oscillatory system, which
61 represents the vertical datum. However, the difficulty in quantitative modelling of the observed
62 deformation and its direct attribution to a seismogenic source remains a major issue (Lacave et al., 2004;
63 Becker et al., 2006, Ferranti et al., 2019).

64 The aim of the present work is to reconstruct the speleoseismological history of Cola Cave (Abruzzo
65 region, Central Italy, Figure 1) and to investigate the causes that led to the collapse of a tall (173 cm high)
66 speleothem found in the cave. We were motivated to establishing whether this collapse represents the
67 record of a large paleoearthquake, and how this and other speleoseismic event that we found in the cave
68 are framed in the seismotectonic context of the Central Apennines.

69 Cola Cave is located in the internal sector of the Central Apennines (Figure 1), an area presently
70 affected by NE-SW trending geodetic extension at 2-3 mm/yr (Devoti et al., 2011; D'Agostino, 2014; Cheloni

71 et al., 2017). Extension is accommodated by systems of NW-SE striking normal faults (Figure 1; Boncio et
72 al., 2004; Pace et al., 2006; Valentini et al., 2017). The present activity of most Central Apennines faults is
73 constrained by several evidence, and a number of them are considered seismogenic sources potentially
74 able to generate M 6.5–7 earthquakes (Pace et al., 2006, 2010; Verdecchia et al., 2018; DISS Working
75 Group, 2015) documented in the Central Apennines (Figure 1; Boncio et al., 2004; Galli et al., 2008;
76 Chiaraluce et al., 2011).

77 Although a good age record exists for Late Pleistocene-Holocene earthquake slips from fault scarp
78 analysis (e.g. Galadini & Galli, 2000; Roberts & Michetti, 2004; Schlagenhauf et al., 2011), and trench
79 paleoseismology (e.g. Cinti et al., 2011; Galli et al., 2008) on the major Central Apennines faults, there are
80 still contrasting views about the capability of several faults to generate earthquakes.

81

82 **2. Seismotectonic context**

83

84 Cola cave is surrounded by major active faults and the region has been shaken by large
85 earthquakes, the largest of them occurred over the last century (Fucino 1915 Mw 7; L'Aquila 2006 Mw 6.3;
86 Central Italy 2016 Mw 6.5). The active faults are mainly south west-dipping normal faults and are organised
87 in fault systems. Some of them are considered seismogenic (Boncio et al., 2004; Valentini et al., 2017) (red
88 traces in Figure 1).

89 The most important fault system is the “Fucino” (source n. 2 in Figure 1), whose activity controlled
90 the growth of the largest intermontane basin in the Central Apennines during the last ~2-3 Ma. The ~ 38
91 km-long surface expression of the Fucino fault system (Galadini and Galli, 1999) is represented, from north
92 to south, by three main fault segments, namely the Magnola fault (MF, Figure 1), the Marsicana Highway-
93 Parasano fault (MHP, Figure 1) and the San Benedetto dei Marsi-Gioia dei Marsi fault (SBGM, Figure 1) (e.g.
94 Serva et al., 1986; Giraudi, 1988; Michetti et al., 1996; Messina, 1996). The 1915 earthquake activated most
95 of the Fucino source producing coseismic ground displacements up to about 1 m high, as described by
96 Oddone (1915) and confirmed by paleoseismological studies (Serva et al., 1986; Michetti et al., 1996;
97 Galadini et al., 1997; Galadini & Galli, 1999). The paleoseismic trenches revealed at least seven additional
98 surface rupture events during the last ~20 ka with offsets comparable to the 1915 earthquake slip (Galli et
99 al., 2008)., Galli et al. (2012) found trench evidence of recent activity of the MF segment, with one or more
100 faulting events that have affected the Neoglacial colluvia (post 4.5 ka) before Roman age (~2 ka) deposits.
101 The time span of this event roughly matches the 3.4 ± 0.3 ka age of the penultimate exhumation event
102 found on the fault scarp using the cosmogenic ^{36}Cl method (Schlagenhauf et al., 2010).

103 The “Velino” seismogenic source (source n. 4 in Figure 1) is located north of the Fucino basin and
104 is represented by a bedrock scarp along the southern slope of Mt. Velino. The fault is considered currently
105 active (Bosi, 1975; Boncio et al., 2004) as evidenced by geomorphological and paleoseismological

106 observations (Giraudi, 1992, Pace et al., 2008). Schlagenhauf et al. (2011) carried out cosmogenic ^{36}Cl
107 dating of carbonate samples from the exhumed fault plane, reconstructing the slip history over the last 7.5
108 ka. They recognized five exhumation events with the last occurred around 1.3 ± 0.3 ka. A single earthquake
109 is assigned to the structure in the historical catalogue (CPT15; Rovida et al., 2016, 2020), the event of
110 24/02/1904 (I_0 VIII-IX MCS; M_w 5.7).

111 The “*Salto Valley*” fault (source n. 3 in Figure 1) is a nearly 28 km long structure bordering the
112 north-east side of the Salto Valley. Some authors believe that the fault is currently active (Roberts &
113 Michetti, 2004; Papanikolaou et al., 2005) on the basis of morphological indications, while according to
114 others fault activity ended at the beginning of the middle Pleistocene (Galadini & Messina, 2001).

115 The “*Campo Felice-Ovindoli*” fault system (source n. 1 in Figure 1) is characterised by several
116 segments that form a single seismogenic source. The recent activity of the fault system is supported by
117 paleoseismological studies realised on the different segments (e.g. Pantosti et al., 1996; Salvi et al., 2003).
118 Three events that displaced Holocene deposits with a recurrence interval of about 3000 years were
119 identified on the southern segment. The last event occurred in a time interval between 700-1130 years BP
120 (Pantosti et al., 1996). Moreover, the event of 9 September 1349 (M_w 6.3, Rovida et al., 2016) has been
121 attributed to the “*Campo Felice-Ovindoli*” seismogenic source activity (Pace et al., 2006).

122 Finally, the Cola cave is localised exactly along the ~60 km long “*Liri*” fault system. While the
123 southern segment (Sora fault, about 20 km long) shows possible signs of seismic activity (Val Roveto
124 earthquake, 1922, M_w 5.2), the activity of the central-northern sector (about 40 km long, source n. 5 in
125 Figure 1) is still a matter of debate. Galadini & Messina (2004) consider the fault inactive, because of the
126 lack of clear displacement of late Quaternary deposits and suggest that slip occurred during the late
127 Pliocene-Middle Pleistocene. Roberts & Michetti (2004) studied the along strike variation of vertical
128 dislocation and escarpment height of the fault. Assuming that these two parameters are congruous with
129 the values measured on other active faults in central Apennines, they concluded that exposure of the fault
130 surface is due to tectonic displacement and not to erosional exhumation.

131 The choice of Cola cave, for its strategic position in this complex seismotectonic context, may help
132 to better constrain the paleoseismological record of active faults of the region, and to explore the activity
133 of doubtful faults, such as the “*Liri*” fault, whose contribution to seismic hazard and seismic risk models
134 could nevertheless be notable (Scotti et al., 2020).

135

136 **3. Cave morphology**

137

138 Cola cave opens at an elevation of ~1400 m a.s.l. on the south-west side of Mt. Arunzo, a
139 culmination of a ~ 15 km long ridge belonging to the M.ti Carseolani mountain group. The ~200 m long cave
140 is developed within thick Cretaceous limestone beds that cap a succession of Mesozoic platform

141 carbonates.

142 A narrow entrance carved along less competent limestone and marly beds (point *A*, Figure 2) leads
143 to a large sub-rectangular (~60x30 m), ~5-10 m high room occupied by fallen ceiling blocks which reach
144 dimensions up to a few m³ (point *B*, Figure 2). A steep (20° average) northwest-southeast trending slope
145 brings to the more internal part of room *B* where large speleothems and columns are grown upon ancient
146 rockfalls. Here, the cave turns to follow a northeast-southwest trend. To the northeast, the cave again
147 bifurcates (point *C*, Figure 2). The main branch continues towards northeast across rooms rich of
148 concretions (point *D*, Figure 2), whereas a narrower branch detaches to the east and terminates in a
149 circular room with a small ephemeral lake (point *E*, Figure 2). To the southwest of the main room, the cave
150 progressively descends getting narrower, and reaches a series of small rooms decorated by abundant
151 speleothems and pools (point *F*, Figure 2). The cave continues deepening to the southwest through tighter
152 passages and shafts, but we did not explore this section further.

153

154 **4. Material and Methods**

155

156 The research methodology integrates:

- 157 - field work in the cave, which includes structural analysis aimed to exclude that the observed
158 speleothem deformation is the result of local non-tectonic processes, measurement of the
159 geometry of collapsed speleothems, and collection of samples from fallen and regrown speleothems
160 to constraint events age;
- 161 - laboratory work, consisting in (²³⁰Th/U) and AMS-¹⁴C dating of pre- and post-event layers;
- 162 - seismic hazard analysis, to define a seismogenic source model and calculate the expected ground
163 motion at the site in terms of uniform hazard spectra;
- 164 - numerical modelling of a selected fallen stalagmite, to study the speleothem behaviour during the
165 passage of the expected seismic waves, and to estimate the possible seismogenic source of the
166 recorded speleoseismic event.

167

168

169 *4.1 Speleothem sampling and cave structural analysis*

170

171 We measured the geometry of 29 collapsed speleothems using a tape and sampled by hammer a
172 selection of them for radiometric dating and geo-mechanical tests. For each measured speleothem, we
173 recorded the length and the average diameter of the fallen section. For stalagmites, we also tried to
174 recognize the standing stump from which the speleothem fall. Commonly, the stump narrows upwards to
175 form a neck of the pre-failure stalagmite. Indeed, when a large earthquake occurs in the vicinity of the

176 cave, the speleothems oscillate due to the ground motion until its failure tensile stress may be overcome
177 and the speleothem breaks in the weakest part, the neck (Figure 3a, and 3b). Because the stalagmite neck
178 is characterized by a change in orientation of C axes of calcite crystals from horizontal at the base to vertical
179 above (P. Forti, personal communication, 1999), it represents the weakest section and the potential locus
180 of breakage during seismic shaking (Figure 3b; Lacave et al., 2004).

181 Based on cave topography and on the geomorphological setting (relative ease of access), we cannot
182 exclude in principle that the observed speleothem deformation is the result of non-tectonic processes such
183 as water flooding, flow of ice during past glacial stages, and animal or anthropic passage (Becker et al.,
184 2012; Jaillet et al., 2006; Kagan et al., 2005, 2017). All the above processes are expected to produce a
185 random orientation of fallen speleothems. Conversely, during seismic shaking, it is predicted that
186 speleothems fall to the ground along a direction broadly aligned with the mean direction of the seismic
187 waves (Figure 3b). To disentangle between a tectonic versus a non-tectonic origin for the observed
188 collapses, we measured the fall azimuth of 29 stalagmites. The azimuth distribution clusters along the
189 NNW-SSE direction, with minor petals to the WSW and NNE (Figure 2). The tight azimuthal distribution
190 argues for a tectonic origin, inasmuch it is roughly parallel with the mean strike of the seismogenic sources
191 of the area (Figure 1).

192 Finally, we sampled both the tip of fallen speleothem and in few cases the base of speleothems
193 grown on the top of the collapsed ones, to tightly constrain the age of the collapses (Figure 3c). In some
194 cases, if the speleothem is not vulnerable enough, instead of collapsing, it could change its growth axis. In
195 this case the evidences of earthquakes are in the inner structure of it (Figure 3d). Indeed, the youngest
196 layer of a fallen speleothem and the base of an intact stalagmite grown upon it represent the pre-event and
197 post-event layers, respectively (Forti & Postpischl, 1984; Becker et al., 2012; Kagan, 2012; Kagan et al.,
198 2005). When the growth of an intact speleothem was not observed above a collapsed speleothem, we only
199 took its youngest, pre-event layer.

200

201 *4.2 Radiometric dating*

202

203 Field samples were sawed in half to identify the layers for dating. Subsamples were extracted from
204 the chosen laminae of stalagmites and stalactites using a dental drill with diamond-encrusted blades.

205 Both the U-series decay ($^{230}\text{Th}/\text{U}$) and AMS- ^{14}C dating techniques were used to date nine samples
206 (Figures 2 and 4). The ages of five samples were accurately determined by U/Th at the Laboratoire des
207 Sciences du Climat et de l'Environnement (LSCE) at Gif-sur-Yvette, France using a ThermoScientific
208 Neptune^{Plus} multi-collector inductively coupled plasma-mass spectrometer following the protocol
209 developed at LSCE (Columbu et al., 2015; Pons-Branchu et al., 2014). Details of the analytical procedure for
210 U/Th can be found in Ferranti et al. (2019). An initial $^{230}\text{Th}/^{232}\text{Th}$ activity ratio of 1.50 ± 0.75 (Hellstrom, 2006)

211 was used to correct for the non-radiogenic detrital ^{230}Th fraction (Table 1). Four samples were AMS- ^{14}C
212 dated at the Center of Applied Physics, Dating and Diagnostics (CEDAD) Laboratory of the University of
213 Salento, Italy (three samples) and at Beta Analytic, Inc. (one sample). Following Ferranti et al. (2019), the
214 radiocarbon ages were first corrected for the “dead carbon portion” (DCP) and then converted to calendar
215 years (cal. yr BP, BP = AD 1950) using the IntCal13 calibration curve (Reimer et al., 2013) and the Calib 7.10
216 program (Stuiver et al., 2020). Based on the DCP calculated in Ferranti et al. (2019), we used values of 5%
217 and 10% (Table 2), which translates to a difference between DCP-corrected and DCP-uncorrected ^{14}C ages
218 of approximately 10-20%. As a preferred age, we consider for each sample the average between the
219 calibrated ages and related uncertainty range from the 5% and 10% DCP-corrected ages and uncertainty
220 (Table 2).

221

222 *4.3 Seismic hazard analysis*

223

224 We carried out a seismic hazard analysis to model the impact of past large earthquakes generated
225 by faults located in the surroundings of Cola cave. We defined the uniform hazard spectrum (*UHS*) for each
226 seismogenic source at the investigation site. We used the calculated spectra in a deterministic approach to
227 study the behaviour of COLA-3 speleothem (from which COLA-3 sample was taken for dating), found fallen
228 from its stump in Room *D* (Figure 2), through a numerical finite element modelling (*FEM*). The approach
229 permits to estimate which seismogenic source(s) were potentially responsible of the speleothem collapse.

230 We selected COLA-3 (Figure 4b-c) because (i) the azimuth of the fall ($\text{N}355^\circ\text{E}\pm 3$) is consistent with
231 the general NNW-SSE pattern of measured falls (Figure 2) arguing for a tectonic origin of the collapse; and
232 (ii) we were able to date both the last pre-collapse (COLA-3) layer of the stalagmite and the initial post-
233 collapse layer of two small stalagmites grown above its stump (COLA-1 and COLA-2, Figure 4b-c). We
234 reconstructed the pre-collapse geometry of COLA-3 (Figure 5), which, considering the stump, was 1.73
235 meters tall, has an average section of 0.34 meters, wider at the base (up to ~ 0.46 meters) and narrows
236 towards the tip (up to ~ 0.07 meters). In particular, the ~ 0.2 m tall stump shrinks at the top where it is ~ 30
237 cm wide, and probably this feature was a weak point along COLA-3 axis.

238 We carried out the seismic hazard analysis in three steps: 1) seismic noise measurements to
239 quantify possible amplification phenomena; 2) definition of a seismogenic source model in the area
240 surrounding the cave; and 3) definition of the uniform hazard spectra to be used for the numerical
241 modelling.

242 First, three single stations seismic noise measurements were acquired in different rooms of Cola
243 cave (COL1 and STR1 stations, Figure 2) and just outside the cave (OUT1 station, Figure 2). The stations
244 were equipped with a Lennartz 5 s velocimeter and a 24-bit Reftek 130 digitizer. Each station acquired a
245 continuous time window of noise of at least 1 hour long. The seismic data were used to evaluate the

246 resonance frequency of measurement sites and to quantify possible amplification phenomena. For this
247 purpose, the seismic noise was used to compute the spectral ratio (H/V) of the horizontal component
248 respect to the vertical (Bonnetfoy-Claudet et al., 2006). H/V curves were computed at each site using the
249 free software Geopsy (Wathelet et al., 2020; www.geopsy.org). The analysis was performed starting from
250 the continuously recorded data divided using a 40 s long moving time-window. An anti-trigger software was
251 used to remove the windows with short transient which could be produced by anthropic activities, local
252 seismic events or rockfalls. For each accepted time window, we removed the mean, the linear trend and
253 applied a 5% cosine taper. Then, we computed the Fourier Amplitude Spectra (FAS) and smoothed the
254 amplitudes following the method proposed by Konno & Omachi (1998), using a coefficient of 40 for the
255 bandwidth. The two horizontal spectra were combined with the geometrical mean and finally divided by
256 the vertical spectrum to get the H/V ratios.

257 We used a seismogenic source model that summarizes and updates the knowledge of active
258 tectonics in the inner part of the Central Apennines (Galadini & Messina, 2004; Galadini & Galli, 2000;
259 Boncio et al., 2004; Pace et al., 2006; Valentini et al., 2017). We used a distance threshold of 50 km,
260 corresponding approximately to an expected ground shaking of ~ 0.25 g (PGA) from the known seismogenic
261 sources, beyond which we consider the effect of seismogenic sources negligible for the seismic hazard in
262 the area. We defined five seismogenic sources (Figure 1), namely “*Campo Felice-Ovindoli*” (1), “*Fucino*” (2),
263 “*Salto Valley*” (3), “*Velino*” (4) and “*Liri*” (5), representative of as many southwest-dipping active normal
264 faults. The length of the sources ranges between ~ 11 (source n.4) and ~ 38 km (source n.2), and the
265 seismogenic thickness is 11-15 km. The seismogenic source geometry has been parameterized using
266 published data (Pace et al., 2014; Peruzza et al., 2011; Faure-Walker, 2014; Valentini et al., 2019; Table 3).
267 For the “*Liri*” source, we modelled the northern sector, for a total length of 27 km suggested by
268 geomorphological considerations (Faure-Walker, 2014) and by aspect ratio relationships (Peruzza & Pace,
269 2002). We did not incorporate the southern sector of the fault because the cave is too far from it and
270 would lay in its footwall, less prone to seismic shaking.

271 We used kinematic, geometric, and slip rate information for each source as inputs for the *FISH*
272 code (Pace et al., 2016) to calculate the global budget of the seismic moment rate allowed by the structure.
273 This calculation is based on predefined size-magnitude relationships in terms of the maximum expected
274 magnitude (M_{max}) and associated mean recurrence time (T_{mean}). The M_{max} values have been calculated using
275 different empirical relationships, benchmarked against observed magnitudes when available, following the
276 approach suggested by Pace et al. (2016). Results define a M_{max} ranging from M_w 6.1 (source n.4) to M_w 6.8
277 (source n.2) (Table 3).

278 Once the maximum expected magnitudes (M_{max}) for each source was computed, the OpenQuake
279 Engine software (Pagani et al., 2014) was used to calculate the Uniform Hazard Spectra (*UHS*) for the
280 following finite element modelling. The spectra were calculated at the cave site using two different ground

281 motion prediction equations (*GMPE*; Bindi et al., 2011; Cauzzi et al., 2015). These two *GMPEs* have been
282 chosen because they use two different metrics: the Joyner-Boore distance (Bindi et al., 2011) and the
283 shortest rupture distance (Cauzzi et al., 2015). To explore the impact of uncertainty of the M_{max} of each
284 source in our estimation, the *UHS* for the $M_{max} \pm$ its standard deviation ($1\text{ }sD$, Table 3) have been calculated
285 (Figure 6). Furthermore, to consider the random uncertainty linked to the standard deviation of each
286 *GMPE*, we performed a number n of simulations of the scenario decided by the user, following a Monte
287 Carlo-like approach. For this work, we performed 10,000 different simulations for each *GMPE*, for a total of
288 20,000 scenario simulations, and finally we calculated the average between the two relationships giving a
289 weight of 0.5 to each relationship. Subsequently, we calculated the median and the confidence interval at
290 16% and 84% for the averaged 10,000 scenarios (Figure 6). For the finite element modelling, the two
291 extremes have been selected for each source, i.e. the maximum and the minimum spectrum (Figure 6). In
292 doing so, we consider the uncertainty in the magnitude, the uncertainty in the individual *GMPE* and the
293 uncertainty related to the use of different *GMPEs*. Following this approach, we calculated the *UHS* for each
294 seismogenic source (Figure 7); the differences between them depend mainly on the different M_{max} and the
295 different distance from Cola cave.

296

297 4.4 Numerical modelling

298

299 Because a speleothem can be considered as a cantilever beam, with height H and diameter D (e.g.
300 Cadorin et al., 2001; Mendecki and Szczygieł, 2019; Figure 5), we performed a numerical modelling to study
301 its behaviour during the passage of the seismic waves, based on its natural frequency of vibration. We used
302 a *FEM* modelling approach taking advantage of the *SAP2000* software
303 (www.csiamerica.com/products/sap2000), which uses geometric, physical, and mechanical properties to
304 compute the response of speleothems to the ground motion. This method is a numerical technique
305 designed to approximate a problem described by partial differential equations, making the solution of the
306 problem a system of algebraic equations. By this method, the macroscopic system is subdivided into small
307 elements (Figure 5), and then the contribution of each element is cumulated to study the overall behaviour.
308 To perform the *FEM* modelling with the *SAP2000* software, it is necessary to know the geometric (H , D),
309 physical and mechanical (ρ , Ed , σ_T) parameters of the speleothems.

310

311 The geometric parameters were collected during the cave survey and are described in paragraph
312 5.1. The collected samples were analysed at the geotechnical laboratory of *DICEA* (*Department of Civil,*
313 *Architectural and Environmental Engineering - Federico II University of Naples*) to evaluate physical and
314 mechanical properties. Specifically, density values ρ of irregularly shaped samples were obtained by using
315 the water displacement method, while mechanical parameters (compressive strength σ , dynamic young
modulus Ed) were derived from uniaxial compressive load-deformation tests (Table 4) (see also Colella et

316 al., 2017). Furthermore, the tensile strength (σ_T) has been obtained empirically according to literature data
317 that indicate values between 5% and 10% for the unconfined compressive strength (Hoek & Brown, 1997;
318 Nazir et al., 2013; Trivedi, 2013; Perras & Diederichs, 2014). Specifically, we evaluated the threshold values
319 as the minimum tensile strength that must be reached to break the speleothems (dotted lines in Figures 10
320 and 11). We used the minimum measured value of unconfined compressive strength (7.35 MPa) and so the
321 thresholds range from 0.365 to 0.735 MPa. It should be noted that measured values of density ρ , dynamic
322 young modulus Ed and tensile strength σ_T agree well with previous studies conducted on speleothems
323 (stalactites, stalagmites, pillars or soda straw) from different caves and contexts (Gribovszki et al., 2008,
324 2013, 2018; Cadorin et al., 2001; Lacave et al., 2000, 2003, 2004; Szeidovitz et al., 2008; Gilli et al., 1999).

325 After the geometric and mechanical parameters are defined, the first step in this method is the
326 discretization of the whole system into finite elements, called nodes and beams (Figure 5). Once the beam
327 element with height H is created, this element is divided into n elements of smaller dimensions that are
328 connected by nodes. Then, we assigned to each element a circular section, with diameter D and a constant
329 value of ρ and Ed . In this work, the speleothems were considered without variations in mechanical
330 parameters (ρ and Ed) and diameter along their axis, but we explored the uncertainty in the ρ evaluation,
331 using the two limit values in the calculations (1.93-2.35 g/cm³). In the same way, we explored the impact of
332 the uncertainty in the dynamic young modulus Ed value, which, unlike ρ , can be considered neglectable in
333 the modelling. The last step in the *FEM* analysis is the selection of the earthquake ground acceleration; the
334 software allows the user to select different seismic input among which the response-spectrum to be used
335 in the modelling. In this work, we used the *UHS* obtained following the seismic hazard analysis described
336 before.

337 Because of the variable geometry of stalagmite COLA-3 along its axis, two types of modelling were
338 carried out, called "simple" and "complex", respectively. In the former scenario, the diameter size does not
339 vary along the axis of the speleothem, whose geometry is cylindrical. In the latter model the diameter size
340 varies along the axis in a way that best approximates the actual geometry (Figure 5). This second modelling
341 allowed to highlight the critical section of the stalagmite, such as the base and the tip, which results in a
342 very different behaviour relative to the simple modelling. The same geometric and mechanical parameters
343 were used in both simple and complex modelling, except for the speleothem diameter. In the first case the
344 diameter is equal to 0.34 cm and in the second case it varies along the axis of the speleothem (Figure 5).
345 Once all input parameters are set, the software provides, for a given direction of acceleration, the
346 maximum displacements, forces (and their moments M), and stresses for each of the vibration mode of the
347 speleothem. Using the moment (M) in the following equation:

348

349
$$\sigma = Mr/l \quad (\text{eq. 1})$$

350

351 where r is the radius of the speleothem, I is the moment of inertia ($I=(r^4*\pi)/4$) of the
352 circumferential transversal speleothem section of radius r , it is possible to calculate the failure tensile stress
353 (σ) due to the selected earthquake ground acceleration. In an analytical approach, the acceleration value is
354 obtained from eq. 1 so to reach the tensile stress value beyond which the speleothem breaks. In our
355 approach instead, we do not have to calculate the acceleration value as it is one of our input data. We
356 instead compare the changes in tensile stress data obtained through modelling (two values exploring the
357 density uncertainty, dashed and solid lines in Figures 10 and 11) with the tensile strength thresholds
358 obtained from laboratory tests (dotted lines in Figures 10 and 11). Accordingly, we can establish whether a
359 given seismic input, and therefore the seismogenic source associated with it, is able or not to cause the
360 collapse of the studied speleothem, and in the former case where the critical section of the speleothem
361 occurs.

362

363 5. Results

364

365 5.1 Age of speleoseismic events

366

367 The total number of collected samples is nine, from four different sites. The U/Th and ^{14}C dating
368 (Tables 1 and 2) provide constraints on the age of speleoseismic events (Table 5, Figure 8).

369 Four samples provide information from room *F* in the inner part of the southwestern branch of the
370 cave (Figure 2). COLA-20 and COLA-21 are two small (~10 cm length) stalactites fallen toward the SSE and
371 partly cemented on the floor (Figures 2, 4a). Nearby TRAP-2 come from the youngest (pre-event) layer of a
372 fallen stalagmite, whereas TRAP-1 is the oldest (post-event) layer of a younger stalagmite (~2 cm long)
373 grown upon the former (Figures 2, 4d). The ages of the samples TRAP-1 and TRAP-2 outline an event
374 between ~8.0-1.5 ka (Table 5). Samples COLA-20 and COLA-21 collapsed after ~8.0 and ~1.1 ka,
375 respectively. Whereas fall of COLA-20 was likely coeval to breakage of TRAP-2, collapse of COLA-21
376 occurred during a younger event.

377 In room *D* on the northeast branch of the cave, sample COLA-3 comes from the pre-event layer of
378 the stalagmite chosen for modelling and samples COLA-1 and COLA-2 were extracted from the post-event
379 layer of two small stalagmites (~10 cm length) grown above the stump left by the collapsed stalagmite
380 (Figures 2, 4b-c). Sample COLA-4 was extracted from the oldest (post-event) layer of a stalagmite grown
381 above an undated fallen speleothem nearby speleothem COLA-3. Samples COLA-3 (pre-event) and COLA-1
382 and COLA-2 (post-event) constrain a speleoseismic event whose age broadly spans between ~7-14 and ~4-5
383 ka (Table 5, Figure 8). The age bracket for this event is rather uncertain. Sample COLA-3 shows a high
384 detrital contamination, as indicated by ^{232}Th (Table 1), and its corrected age strongly depends on the
385 chosen initial $^{230}\text{Th}/^{232}\text{Th}$ activity ratio. For example, if a value of 2.96 ± 0.1 were applied, as in Ferranti et al.

386 (2019), the ^{230}Th -corrected age would be 4.3 ka. This is not the case for the other speleothem samples that
387 show a very small detrital thorium component. The post-event age bracket is more tightly constrained.
388 Although COLA-1 and COLA-2 have ages median probability that differ of ~ 1 ka, they could have started
389 growing within few hundred years from each other in the uncertainty range. Sample COLA-4 age
390 documents a speleoseismic event older than ~ 3.5 ka, which could coincide with the event that caused
391 collapse of COLA-3 within the same room (Table 5, Figures 2, 8).

392 Finally, in room *E* (east branch of Cola cave), we sampled the base of a stalagmite (CRL1 ~ 6 cm long)
393 grown upon a large limestone block fallen from the vault of the cave (Figures 2, 4e). The U/Th age of
394 sample CRL1 implies that the block fell from the ceiling prior to ~ 12.5 ka (Table 5, Figure 8).

395

396 *5.2 Evaluation of amplification phenomena from noise measurement*

397

398 Figure 9 shows the H/V ratios obtained at each station, for clarity of the figure we show the H/V
399 mean curve and standard deviations computed for each measurement site. Figure 9 also reports the results
400 of directional noise analysis and the mean FAS computed for the three different components (Z in blue, NS
401 in green and EW in red) for each station. The H/V curves show a good agreement in terms of shape within
402 the three stations. Two peaks are observed at frequency of about 1.5 Hz and 2.4 Hz in all sites. In all the
403 three cases, the observed H/V peaks are due to signal characterized by energy higher on the horizontal's
404 components compared to the vertical one, as clearly indicated by the FAS (Figure 9). Moreover, a clear
405 polarization in the north and south directions is observed for the H/V peaks around the picks. Finally, also
406 the amplitudes of H/V peaks (mean values) are comparable within the three stations with values close to 2
407 or slightly less than 2 for the peaks at 1.5 Hz and of about $1.7\div 1.8$ for the higher frequency peaks. A further
408 spectral peak around 15-20 Hz is present in only one measurement (COL1). Also, in this latter case, this
409 peak is characterized by an amplitude value (of about 1.5) significantly lower than the threshold used to
410 identify local seismic amplification effects from HV spectral ratios analysis.

411 Summing up, the H/V curves at the sites show similar characteristics both in terms of fundamental
412 frequencies and spectral shapes, the weak peaks observed (with amplitudes ≤ 2) suggest no important
413 amplification phenomena (SESAME criteria and guidelines, Bard et al., 2005). Based on these results, we
414 have not used the site contribution in theoretical spectral response simulated at the cave bedrock.

415

416 *5.3 Numerical modelling*

417

418 The numerical modelling results, obtained with both the "simple" and "complex" models, show
419 how the value of the tensile stress (σ_T) arising from seismic motion varies along the speleothem axis
420 (Figures 10 and 11). In the "simple" modelling (Figure 10) the maximum values of σ_T are at the base of the

421 speleothem, and as we move upward the values decrease until they are null at the top. For three out of the
422 five seismogenic sources for which seismic input was calculated, the minimum threshold value of tensile
423 strength was not reached. This therefore means that a possible earthquake generated by seismogenic
424 sources “Salto Valley”, “Campo Felice-Ovindoli” and “Velino” is not able to cause the collapse of the
425 speleothem when a “simple” geometry is adopted. The “simple” modelling allowed us to make a skimming
426 between the possible sources that caused collapse of COLA-3 and, based on results, we discarded the latter
427 three sources in “complex” modelling.

428 In Figure 11 we compare the variations of the tensile stress values σ_T along the speleothem axis
429 with the threshold values of tensile strength for the “complex” modelling using the input *UHS* obtained for
430 the “Fucino” and “Liri” seismogenic sources. In this scenario, the maximum value of σ_T is not at the base of
431 the speleothem, but it increases from the base to a maximum point beyond which it decreases until zero at
432 the top. This result is in excellent agreement with the geometric data of the modelled speleothem. Indeed,
433 the maximum tensions that led to the speleothem breakage occurred at the exact point where the
434 speleothem has a narrowing of its axis, which therefore represents a vulnerable section. This maximum
435 value of σ_T was produced at ~22 cm height, which perfectly matches the elevation of the stump left by the
436 breakage (Figure 5). Results of the complex modelling show that only the seismogenic source “Liri” reaches
437 the minimum threshold value of tensile strength, suggesting it is the most likely source of the earthquake
438 recorded by the collapse of speleothem COLA-3.

439

440 6. Discussion

441

442 Taken together, ages from pre- and post-event layers define three distinct speleoseismic events:
443 the pre-penultimate with an age >12.5 ka, the penultimate between ~7-14 and ~4-5 ka, and the ultimate
444 after 1.1 ka (Table 5). We relate these events to paleo-earthquakes that caused significant ground motion
445 and collapses of cave speleothems. Unfortunately, pre- and post-event layers are only available for two
446 speleothem couplets (COLA-3/COLA-1&2 and TRAP-1/TRAP-2) and refer to the penultimate event (or for
447 the ultimate event for the latter couple). For the remaining events, either pre- or post-event ages are only
448 defined (Table 5). In addition, a large time span separates pre- and post-event layers for the two
449 speleothem couplets, impacting a tight definition of the age of the penultimate and possibly the ultimate
450 events.

451 We can mitigate the above uncertainties and better refine the paleo-earthquake age estimate
452 based on the acceptable assumption that growth of post-event stalagmites occurred soon after the collapse
453 of the speleothem upon which they grew. We do not consider in this approach sampling and measurement
454 errors that are too small to match the large time gap observed.

455 For the penultimate event, although a significant time span is observed between pre- and post-

456 event layers, the simplest interpretation is that the circulation-precipitation system closed long before the
457 event and was re-opened by seismic shaking. We refer to this assumption as the post-event growth
458 hypothesis. We are aware that closure and opening of the system could arise from changes in climatic and
459 hydrologic conditions as well. However, we do not expect significant variations in the limited time span
460 studied here when the Holocene climatic optimum was attained (Rossignol-Strick, 1999; Giraudi et al.,
461 2011). The hypothesis is corroborated by the observation that, when more than one single age
462 determination is available for post-event layers as in the case of COLA-1, COLA-2 and COLA-4 from room D,
463 their value overlap or is very close. These ages may indicate a sudden opening of the circulation-
464 precipitation system in the limestone above room D that in our view is related to seismic shaking during the
465 penultimate event. The inverse hypothesis, that shaking closed the system that was fortuitously opened
466 long after, seems untenable.

467 Bearing the above in mind, we suggest that the pre-penultimate event occurred at or slightly before
468 ~12.5 ka and is documented by a stalagmite (sample CRL1) grown on a large limestone block fallen from the
469 vault of the cave in room E (Figure 2). We compare this result with the paleoseismological data (trenches
470 and ³⁶Cl cosmogenic data; Galadini & Galli, 1999; Galli et al., 2008, 2012; Schlagenhauf et al., 2010, 2011;
471 Palumbo et al., 2004) available for the “Fucino” source (Figure 8). Cosmogenic data identified a possible
472 exhumation event older than 14.7±1 ka (Schlagenhauf et al., 2011) that could be the same identified in this
473 work. An alternative view, that we regard as more likely based on the post-event growth hypothesis
474 outlined above, is that the pre-penultimate event relates to a faulting event found around 12.5 ka in
475 paleoseismic trenches on the “Fucino” fault system (Galadini & Galli, 1999), which produced a significant
476 vertical displacement (~3 m on SBGM fault and 1-2 m on the Trasacco fault, a secondary splay of the
477 system).

478 Within the post-growth hypothesis, the maximum age of the penultimate event is suggested by the
479 oldest (COLA-1) post-event layer at 5.1 ka, which, within uncertainty, could be as young as 4.8 ka (Table 5).
480 Post-event layers COLA-2 (on the same stalagmite) and COLA-4 (on a nearby speleothem), with maximum
481 ages of 4.3 and 3.8 ka, respectively, probably started growing some hundred to a thousand years after the
482 fall. Comparison of our results for the penultimate event with the “Fucino” paleoseismological observations
483 available from literature (Figure 8) does not fit perfectly the exhumation event recorded by Schlagenhauf et
484 al. (2011) at 4.8 +0.3/-0.4 ka using ³⁶Cl data on the Magnola fault plane or the trench record by Galadini &
485 Galli (1999), where the closest event occurred between ~6.0 and 5.6 ka.

486 Finally, the pre- and post-event layers TRAP-2 and TRAP-1 brackets the ultimate event at ~8.0-1.5
487 ka, and COLA-21 after 1.1 ka. Although the fall of TRAP-2 could have occurred during the penultimate
488 event, by considering the post-growth hypothesis, we regard this fall as constrained by the 1.55±0.07 ka
489 age of TRAP-1. In support of this, nearby stalactite COLA-21 (Figure 2) fell on the ground after 0.91-1.34 ka.
490 Considering uncertainties (Table 5), the two estimates may be separated by only 140 years, therefore they

491 are both assigned to the ultimate event.

492 The ultimate event can be compared to the trench rupture found between 426-782 CE on the
493 Fucino fault (Figure 8; Galli et al., 2008). These authors attributed the trench offset to the earthquakes that
494 hit Rome in 508 CE causing serious damages to the Coliseum (Galadini & Galli, 2001). The last ^{36}Cl
495 exhumation found on the Magnola fault plane (Schlagenhauf et al., 2010, 2011) has an age comparable to
496 the ultimate speleoseismic event as well. The Magnola fault represents the northern segment, closest to
497 the Cola cave, of the “Fucino” fault system. However, the damages occurred in the distant site of Rome
498 suggest that the entire ~38 km fault system ruptured contemporarily (Galli et al., 2012). In principle, we
499 cannot fully exclude that the ultimate speleoseismic event recorded in the cave occurred during the
500 devastating 1915 M_w 7 earthquake, the strongest recorded in central Italy (Rovida et al., 2016). This
501 earthquake was generated similarly by the entire “Fucino” fault system, with surface faulting recognized on
502 several segments (Oddone, 1915), and an inferred rupture directivity from the SE towards the NW (Berardi
503 et al., 1999).

504 For the penultimate event, we carried out a vulnerability analysis of COLA-3 to find a direct
505 attribution to a seismogenic source potentially responsible for the collapse. Using seismogenic source and
506 seismic hazard models, we calculated the expected ground motion (i.e. Uniform Hazard Spectra) in the cave
507 by the selected sources in turn, in a determinist approach, and compared it to the calculated failure tensile
508 stress (σ_T) of the stalagmite. The results suggest two possible seismogenic sources responsible of the
509 collapse: the “Fucino” and the “Liri” sources. The “complex” modelling, where the speleothem diameter
510 size varies along the axis, suggests “Liri” as the most likely source of the recorded shaking (see paragraph
511 5.3). It is important to highlight that we explored uncertainties related to the variability in density and
512 tensile strength, and the maximum expected magnitude and ground motion (i.e. Ground Motion Prediction
513 Equations), but several other sources of error remain unquantified. The results could be affected, for
514 example, by the rupture directivity, as happened for the 1915 earthquake, or by the actual hypocentre
515 location, hypothesized in our modelling at the bottom of the fault rupture and in the centre of the fault
516 system. Besides, we cannot exclude that changes in orientation of C axes of calcite crystals in the
517 speleothem, from horizontal at the base to vertical above, can create sections potentially prone to
518 breakage even in case of weaker ground motions. Finally, we cannot exclude the general variations of the
519 mechanical parameters, like for the example the density δ , inside the stalagmite, that we considered
520 homogeneous.

521 For these reasons, although the “Liri” fault remains the most likely source of the large paleo-
522 earthquakes recorded in the cave, and specifically of the penultimate event, we cannot exclude “Fucino” as
523 another potential source. This alternative scenario may be supported by the possible directivity of the
524 rupture towards NNW and a hypocentre closer to the Cola cave during the speleoseismic events studied
525 here. The “Liri” fault recent activity is debated and no paleoseismological data on the fault are available to

526 be compared to our event age. On the other hand, the abundance of paleoseismological data on the
527 “Fucino” source renders the comparison possible, and we were able to match our speleoseismological data
528 of two events (the ultimate and the pre-penultimate) with the paleoseismological record.

529

530 **7. Conclusions**

531

532 Our study allowed to reconstruct the speleoseismological history of the Cola cave in central Italy and to
533 identify seismogenic sources potentially responsible for the collapse of a tall stalagmite and for two
534 additional speleoseismic events.

535 Using the samples ages and the comparison with available paleoseismological data for active faults in
536 the region surrounding the cave, we identified three speleoseismic events: the pre-penultimate occurred
537 before but probably near ~ 12.5 ka, the penultimate occurred around ~ 5 ka, and the ultimate around ~ 1.5
538 ka and consistent with the 508 AD earthquake, which caused damages to the Coliseum in Rome.

539 The seismic hazard and vulnerability analysis of the tall stalagmite, collapsed during the penultimate
540 event, allowed to find a direct attribution to a seismogenic source potentially responsible for the collapse.
541 Modelling suggests the “Liri” fault as the most likely seismogenic source responsible for the ground shaking.

542 The “Fucino” fault system cannot be excluded as a potential source of collapsing in the cave. In support
543 of this, the ages of two speleoseismic event matches some of the events found in trenches and on
544 exhumed fault planes along the system.

545 We believe that the approach presented here can be useful to improve the paleoseismological data
546 record available in active areas similar to the central Apennines, mainly collected using “classical” trench
547 methods (e.g. Galli et al., 2008) or through more “innovative” investigations like fault surface exposure
548 dating by cosmogenic technique (e.g. Benedetti & Van Der Woerd, 2014; Tesson et al., 2016).

549 The numerical FEM approach, using the best approximation of the complex speleothem geometry,
550 allows to evaluate in detail the speleothem behavior during ground motion, and to precisely define the
551 point of breakage as the weakest section of the speleothem.

552 The integrated use of seismic hazard and speleothem vulnerability analysis allowed to find a direct
553 attribution of the penultimate speleoseismological event to a seismogenic source and suggested as possible
554 the activity of the “Liri” fault, that should be considered more carefully into seismic hazard and risk models.

555 Further developments of the approach could integrate fault segmentation relaxation and the possible
556 fault interactions into the seismic hazard model (e.g. Verdecchia et al., 2018; Visini et al., 2020; Valentini et
557 al., 2020) and a better treatment of the uncertainties involved in the definition of the mechanical
558 properties of broken speleothems and in the age record of speleoseismic events.

559

560 **Acknowledgements**

561 We sincerely thank S. Di Bianco, J. De Massis and P. Teodoro who helped sampling and collecting the cave
562 data; F. Visini who helped us during the seismic noise measurements acquisition and discussed with us the
563 seismic hazard results; D. Berardi who led us in the Cola cave; F. Rizzo who helped us during the initial part
564 of FEM modelling; S. Agostini for useful discussions on speleoseismology and for providing his papers
565 collection. Finally, we would like to thank the Editor T. Schildgen, the Associate Editor L. Dal Zilio and the
566 two reviewers C.P. Rajendran and J. Szczygieł for helpful comments and revisions that significantly
567 improved the manuscript. All the parametric data for sample dating, seismogenic sources and the
568 speleothems physical and mechanical data are reported in tables (Tables 1-2-3-4) and have been taken
569 from published literature, listed in the references. Data supporting the sample dating can be found also in
570 the [Pangaea](https://doi.pangaea.de/10.1594/PANGAEA.922154) repository
571 at <https://doi.pangaea.de/10.1594/PANGAEA.922154> and <https://doi.pangaea.de/10.1594/PANGAEA.922156>.

572

573 References

574

- 575 Bar-Matthews, M., Ayalon, A., Kaufman, A., & Wasserburg, G. J. (1999). The Eastern Mediterranean
576 paleoclimate as a reflection of regional events: Soreq cave, Israel. *Earth and Planetary Science Letters*,
577 *166*(1-2), 85-95. [https://doi.org/10.1016/S0012-821X\(98\)00275-1](https://doi.org/10.1016/S0012-821X(98)00275-1)
- 578 Bard, P.-Y., SESAME-Team (2005). Report D23.12, Guidelines for the Implementation of the H/V Spectral
579 Ratio Technique on Ambient Vibrations Measurements, Processing and Interpretation in European
580 Commission: Research General Directorate, Project No. EVG1-CT-2000-00026. SESAME, p. 62 available
581 online at: <http://sesame-fp5.obs.ujf-grenoble.fr>.
- 582 Becker, A., Davenport, C. A., Eichenberger, U., Gilli, E., Jeannin, P. Y., & Lacave, C. (2006). Speleoseismology:
583 A critical perspective. *Journal of Seismology*, *10*, 371-388. <https://doi.org/10.1007/s10950-006-9017-z>
- 584 Becker, A., Häuselmann, P., Eikenberg, J., & Gilli, E. (2012). Active tectonics and earthquake destructions in
585 caves of northern and central Switzerland. *International Journal of Speleology*, *41*, 35-49.
586 <https://doi.org/10.5038/1827-806X.41.1.5>
- 587 Benedetti, L. C., & Van Der Woerd, J. (2014). Cosmogenic nuclide dating of earthquakes, faults, and toppled
588 blocks. *Elements*, *10*(5), 357-361. <https://doi.org/10.2113/gselements.10.5.357>
- 589 Berardi, R., Contri, P., Galli, P., Mendez, A. & Pacor F. (1999). Modellazione degli effetti di amplificazione
590 locale nelle città di Avezzano, Ortucchio e Sora. In: Castenetto S. and Galadini F. (eds), 13 gennaio del
591 1915. Il terremoto nella Marsica, Servizio Sismico Nazionale, pp. 349-372.
- 592 Bindi, D., Pacor, F., Luzi, L., Puglia, R., Massa, M., Ameri, G., & Paolucci, R. (2011). Ground motion prediction
593 equations derived from the Italian strong motion database. *Bulletin of Earthquake Engineering*, *9*,
594 1899-1920. <https://doi.org/10.1007/s10518-011-9313-z>

595 Boncio, P., Lavecchia, G., & Pace, B. (2004). Defining a model of 3D seismogenic sources for Seismic Hazard
596 Assessment applications: The case of central Apennines (Italy). *Journal of Seismology*, 8(3), 407–425.
597 <https://doi.org/10.1023/B:JOSE.0000038449.78801.05>

598 Bonnefoy-Claudet, S., Cornou, C., Bard, P. Y., Cotton, F., Moczo, P., Kristek, J., et al. (2006). H/V ratio: a tool
599 for site effects evaluation. Results from 1-D noise simulations. *Geophysical Journal
600 International*, 167, 827–837.

601 Bosi, C. (1975). Osservazioni preliminari su faglie probabilmente attive nell'Appennino centrale. *Boll. Soc.
602 Geol. Ital.* 94 (1975) 827– 859.

603 Cadornin, J. F., Jongmans, D., Plumier, A., Camelbeeck, T., Delaby, S., & Quinif, Y. (2001). Modelling of
604 speleothems failure in the Hotton cave (Belgium). Is the failure earthquake induced? *Geologie En
605 Mijnbouw/Netherlands Journal of Geosciences*, 80(3-4), 315-321.
606 <https://doi.org/10.1017/s001677460002391x>

607 Cauzzi, C., Faccioli, E., Vanini, M., & Bianchini, A. (2015). Updated predictive equations for broadband
608 (0.01–10 s) horizontal response spectra and peak ground motions, based on a global dataset of digital
609 acceleration records. *Bulletin of Earthquake Engineering*, 13(6), 1587–1612.
610 <https://doi.org/10.1007/s10518-014-9685-y>

611 Cheloni, D., De Novellis, V., Albano, M., Antonioli, A., Anzidei, M., Atzori, S., et al. (2017). Geodetic model of
612 the 2016 Central Italy earthquake sequence inferred from InSAR and GPS data. *Geophysical Research
613 Letters*, 44(13), 6778–6787. <https://doi.org/10.1002/2017GL073580>

614 Chiaraluce, L., Valoroso, L., Piccinini, D., Di Stefano, R., & De Gori, P. (2011). The anatomy of the 2009
615 L'Aquila normal fault system (central Italy) imaged by high resolution foreshock and aftershock
616 locations. *Journal of Geophysical Research: Solid Earth*, 116(12).
617 <https://doi.org/10.1029/2011JB008352>

618 Cinti, F. R., Pantosti, D., De Martini, P. M., Pucci, S., Civico, R., Pierdominici, S., et al. (2011). Evidence for
619 surface faulting events along the Paganica fault prior to the 6 April 2009 L'Aquila earthquake (central
620 Italy). *Journal of Geophysical Research: Solid Earth*, 116(7). <https://doi.org/10.1029/2010JB007988>

621 Colella, A., Cremona, M., Di Bianco, S., Ferranti, L., Ramondini, M., & Calcaterra, D. (2017). Valutazione dei
622 principali parametri fisico- meccanici di alcuni speleotemi Italiani. In Atti III Convegno Regionale di
623 Speleologia «Campania Speleologica», 2-4 Giugno 2017, Napoli, Italy, pp. 101-112. ISBN 978-88-
624 89897-16-4.

625 Columbu, A., De Waele, J., Forti, P., Montagna, P., Picotti, V., Pons-Branchu, E., Hellstrom, J., Bajo, P., &
626 Drysdale, R. (2015). Gypsum caves as indicators of climate-driven incision and aggradation in a rapidly
627 uplifting region. *Geology*, 43, 539-542.

628 D'Agostino, N. (2014). Complete seismic release of tectonic strain and earthquake recurrence in the
629 Apennines (Italy). *Geophysical Research Letters*, 41(4), 1155–1162.
630 <https://doi.org/10.1002/2014GL059230>

631 Devoti, R., Esposito, A., Pietrantonio, G., Pisani, A. R., & Riguzzi, F. (2011). Evidence of large scale
632 deformation patterns from GPS data in the Italian subduction boundary. *Earth and Planetary Science
633 Letters*, 311(3–4), 230–241. <https://doi.org/10.1016/j.epsl.2011.09.034>

634 Di Domenica, A., & Pizzi, A. (2017). Defining a mid-Holocene earthquake through speleoseismological and
635 independent data: Implications for the outer Central Apennines (Italy) seismotectonic framework.
636 *Solid Earth*, 8, 161–176. <https://doi.org/10.5194/se-8-161-2017>

637 DISS Working Group (2015). Database of individual seismogenic sources (DISS), version 3.2.0: A compilation
638 of potential sources for earthquakes larger than M 5.5 in Italy and surrounding areas.
639 <http://diss.rm.ingv.it/diss>, Istituto Nazionale di Geofisica e Vulcanologia, doi:
640 <https://doi.org/10.6092/INGV.IT-DISS3.2.0>

641 Faure Walker, J. (2014). Mechanics of continental extension from Quaternary strain fields in the Italian
642 Apennines. *PhD Thesis*, pp 405.

643 Ferranti, L., Pace, B., Valentini, A., Montagna, P., Pons-Branchu, E., Tisnérat-Laborde, N., & Maschio, L.
644 (2019). Speleoseismological constraints on ground shaking threshold and seismogenic sources in the
645 Pollino range (Calabria, southern Italy). *Journal of Geophysical Research: Solid Earth*, 124(5), 5192-
646 5216. <https://doi.org/10.1029/2018JB017000>

647 Forti, P., & Postpischl, D. (1980). Neotectonic data from stalagmites: sampling and analysis techniques.
648 European Regional Conference on Speleology Sofia, CNR – Progetto Finalizzato Geodinamica, 351
649 Sofia, 2, 34-39.

650 Forti, P., Petrini, V., & Postpischl, D. (1981). Ricostruzione di fenomeni paleosismici da strutture carsiche.
651 *Rendiconti Società Geologica Italiana*, 1981(4),563-569.

652 Forti, P., & Postpischl, D. (1984). Seismotectonic and paleoseismic analyses using karst sediments. *Marine
653 Geology*, 55(1-2), 145-161. [https://doi.org/10.1016/0025-3227\(84\)90138-5](https://doi.org/10.1016/0025-3227(84)90138-5)

654 Forti, P. (2001). Seismotectonic and paleoseismic studies from speleothems: the state of the art. *Geologica
655 Belgica*, 4(3-4), 175-185.

656 Galadini, F., & Galli, P. (1999). The Holocene paleoearthquakes on the 1915 Avezzano earthquake faults
657 (central Italy): Implications for active tectonics in the central Apennines. *Tectonophysics*, 308(1-2), 143-
658 170. [https://doi.org/10.1016/S0040-1951\(99\)00091-8](https://doi.org/10.1016/S0040-1951(99)00091-8)

659 Galadini, F., & Galli, P. (2000). Active tectonics in the central Apennines (Italy)—input data for seismic hazard
660 assessment. *Natural Hazards*, 22, 225–270.

661 Galadini, F., & Galli, P. (2001). Archaeoseismology in Italy: Case studies and implications on long-term
662 seismicity. *Journal of Earthquake Engineering*, 5(1), 35-68.
663 <https://doi.org/10.1142/S1363246901000236>

664 Galadini, F., Galli, P., & Giraudi, C. (1997). Geological investigations of Italian earthquakes: New
665 paleoseismological data from the Fucino Plain (Central Italy). *Journal of Geodynamics*, 24(1-4), 87-103.
666 [https://doi.org/10.1016/s0264-3707\(96\)00034-8](https://doi.org/10.1016/s0264-3707(96)00034-8)

667 Galadini, F., & Messina, P. (2001). Plio-quadernary changes of the normal fault architecture in the central
668 Apennines (Italy). *Geodinamica Acta*, 14(6), 321-344.
669 <https://doi.org/10.1080/09853111.2001.10510727>

670 Galadini, F., & Messina, P. (2004). Early-Middle Pleistocene eastward migration of the Abruzzi Apennine
671 (central Italy) extensional domain. *Journal of Geodynamics*, 24(1-4), 87-103.
672 <https://doi.org/10.1016/j.jog.2003.10.002>

673 Galli, P., Galadini, F., & Pantosti, D. (2008). Twenty years of paleoseismology in Italy. *Earth-Science Reviews*,
674 88(1-2), 89-117. <https://doi.org/10.1016/j.earscirev.2008.01.001>

675 Galli, P., Messina, P., Giaccio, B., Peronace, E., & Quadrio, B. (2012). Early Pleistocene to late Holocene
676 activity of the Magnola fault (Fucino fault system, central Italy). *Bollettino Di Geofisica Teorica Ed*
677 *Applicata*, 53(4), 435-458. <https://doi.org/10.4430/bgta0054>

678 Gascoyne, M., Schwarcz, H. P., & Ford, D. C. (1980). A palaeotemperature record for the mid-Wisconsin in
679 Vancouver Island. *Nature*, 285, 474-476. <https://doi.org/10.1038/285474a0>

680 Gilli, E., Levret, A., Sollogoub, P., & Delange, P. (1999). Research on the February 18, 1996 earthquake in the
681 caves of Saint-Paul-de-Fenouillet area, (Eastern Pyrenees, France). *Geodinamica Acta*, 12(3-4), 143-
682 158. <https://doi.org/10.1080/09853111.1999.11105338>

683 Giraudi, C. (1988). Evoluzione geologica della Piana del Fucino (Abruzzo) negli ultimi 30.000 anni. *Il*
684 *Quaternario*, 1(2), 131-159.

685 Giraudi, C. (1992). Segnalazione di scarpate di faglia tardo-pleistoceniche sui Monti della Magnola
686 (Massiccio del Velino - Abruzzo). *Il Quaternario*, 5(1), 27-32.

687 Giraudi, C., Magny, M., Zanchetta, G., & Drysdale, R.N. (2011). The Holocene climatic evolution of
688 Mediterranean Italy: a review of the continental geological data. *The Holocene*, 21(1), 105-115.

689 Gribovszki, K., Paskaleva, I., Kostov, K., Varga, P., & Nikolov, G. (2008). Estimating an upper limit on
690 prehistoric peak ground acceleration using the parameters of intact speleothems in caves in
691 southwestern Bulgaria. In A. Zaicenco, I. Craifaleanu, I. Paskaleva (Eds.) Harmonization of seismic
692 hazard in Vrancea Zone with special emphasis on seismic risk reduction. NATO Science for peace and
693 security, series C: environmental security (pp. 287-308). Dordrecht: Springer. (ISBN: 978-1-4020-9241-
694 1).

695 Gribovszki, K., Bokelmann, G., Szeidovitz, G., Varga, P., Paskaleva, I., Brimich, L., & Kovacs., K. (2013).
696 Comprehensive investigation of intact, vulnerable stalagmites to estimate an upper limit on
697 prehistoric ground acceleration. In Proceedings of the Vienna congress on recent advanced in
698 earthquake engineering and structural dynamics & 13. D-A-CH Tagung, Vienna, Paper No. 445, p. 10.

699 Gribovszki, K., Esterhazy, S., & Bokelmann, G. (2018). Numerical Modeling of Stalagmite Vibrations. *Pure*
700 *and Applied Geophysics*, 175, 4501-4514. <https://doi.org/10.1007/s00024-018-1952-4>

701 Hellstrom, J. (2006). U-Th dating of speleothems with high initial ²³⁰Th using stratigraphic constraint.
702 *Quaternary Geochronology*, 1(4), 289-295. doi:10.1016/j.quageo.2007.01.004

703 Hoek, E., & Brown, E. T. (1997). Practical estimates of rock mass strength. *International Journal of Rock*
704 *Mechanics and Mining Sciences*, 34(8), 1165-1186. [https://doi.org/10.1016/S1365-1609\(97\)80069-X](https://doi.org/10.1016/S1365-1609(97)80069-X)

705 Jaillet, S., Pons-Branchu, E., Maire, R., Hamelin, B., & Brulhet, J. (2006). Enregistrement de paléo-mises en
706 charge Holocènes dans deux stalagmites du réseau du rupt-du-puits (Barrois, France). Analyses
707 morphologiques des lamines et datations U/Th en TIMs. *Geologica Belgica*, 9(3-4), 297–307.

708 Kagan, E. J., Agnon, A., Bar-Matthews, M., & Ayalon, A. (2005). Dating large infrequent earthquakes by
709 damaged cave deposits. *Geology*, 33(4), 261-264. <https://doi.org/10.1130/G21193.1>

710 Kagan, E.J., (2012). Multi-site Late Quaternary Paleoseismology in the Dead Sea Transform Region:
711 Independent Recording by Lake and Cave Sediments. Geological Survey of Israel report, GSI/17/2012.

712 Kagan, E. J., Cinti, F. R., Alfonsi, L., Civico, R., & Bar-Matthews, M. (2017). Broken speleothems reveal
713 Holocene and Late Pleistocene paleoearthquakes in Northern Calabria, Italy. *Quaternary International*,
714 451, 176-184. <https://doi.org/10.1016/j.quaint.2016.10.023>

715 Konno, K., & Ohmachi, T. (1998). Ground-motion characteristics estimated from spectral ratio between
716 horizontal and vertical components of microtremor. *Bulletin of the Seismological Society of America*,
717 88(1), 228-241.

718 Lacave, C., Levret, A., & Koller, M. (2000). Measurements of natural frequencies and damping of
719 speleothems. *Proc. of the 12th World Conference on Earthquake Engineering*, Auckland, New-
720 Zealand, Paper 2118, 2000.

721 Lacave, C., Koller, M. W., Eichenberger, U., & Jeannin, P. Y. (2003). Prevention of speleothem rupture during
722 nearby construction. *Environmental Geology*, 43, 892-900. [https://doi.org/10.1007/s00254-002-0719-](https://doi.org/10.1007/s00254-002-0719-4)
723 4

724 Lacave, C., Koller, M. G., & Egozcue, J. J. (2004). What can be concluded about seismic history from broken
725 and unbroken speleothems? *Journal of Earthquake Engineering*, 8(3), 431-455.
726 <https://doi.org/10.1080/13632460409350496>

727 Lauritzen, S. E. (1995). High-Resolution Paleotemperature Proxy Record for the Last Interglaciation Based
728 on Norwegian Speleothems. *Quaternary Research*, 43(2), 133-146.
729 <https://doi.org/10.1006/qres.1995.1015>

730 Li, W. X., Lundberg, J., Dickin, A. P., Ford, D. C., Schwarcz, H. P., McNutt, R., & Williams, D. (1989). High-
731 precision mass-spectrometric uranium-series dating of cave deposits and implications for
732 palaeoclimate studies. *Nature*, *339*, 534-536. <https://doi.org/10.1038/339534a0>

733 McDermott, F. (2004). Palaeo-climate reconstruction from stable isotope variations in speleothems: A
734 review. In *Quaternary Science Reviews*, *23*(7-8), 901-918.
735 <https://doi.org/10.1016/j.quascirev.2003.06.021>

736 Mendecki, M., & Szczygieł, J. (2019). Physical constraints on speleothem deformations caused by
737 earthquakes, seen from a new perspective: Implications for paleoseismology. *Journal of Structural*
738 *Geology*, *126*, 146–155, doi:10.1016/j.jsg.2019.06.008.

739 Messina, P. (1996). Tettonica mesopleistocenica dei terrazzi nord-orientali del Fucino (Italia centrale). *Il*
740 *Quaternario*, *9*(1), 393-398.

741 Michetti, A. M., Brunamonte, F., Serva, L., & Vittori, E. (1996). Trench investigations of the 1915 Fucino
742 earthquake fault scarps (Abruzzo, central Italy): Geological evidence of large historical events. *Journal*
743 *of Geophysical Research: Solid Earth*, *101*(B3), 5921-5936. <https://doi.org/10.1029/95jb02852>

744 Nazir, R., Momeni, E., Armaghani, D. J., & Amin, M. F. M. (2013). Correlation between unconfined
745 compressive strength and indirect tensile strength of limestone rock samples. *Electronic Journal of*
746 *Geotechnical Engineering*, *18*, 1737-1746.

747 Oddone E., (1915). Gli elementi fisici del grande terremoto marsicano-fucense del 13 gennaio 1915. *Boll.*
748 *Soc. Sism. It.*, *19*, 71-215.

749 Pace, B., Peruzza, L., Lavecchia, G., & Boncio, P. (2006). Layered seismogenic source model and probabilistic
750 seismic-hazard analyses in central Italy. *Bulletin of the Seismological Society of America*, *96*(1), 107–
751 132. <https://doi.org/10.1785/0120040231>

752 Pace, B., Boncio, P., Brozzetti, F., Lavecchia, G., & Visini, F. (2008). From regional seismic hazard to “scenario
753 earthquakes” for seismic microzoning: A new methodological tool for the Celano Project. *Soil*
754 *Dynamics and Earthquake Engineering*, *28*(10–11), 866–874.
755 <https://doi.org/10.1016/j.soildyn.2007.11.001>

756 Pace, B., Peruzza, L., & Visini, F. (2010). LASSCI2009.2: Layered earthquake rupture forecast model for
757 central Italy, submitted to the CSEP project. *Annals of Geophysics*, *53*(3), 85–97.
758 <https://doi.org/10.4401/ag-4847>

759 Pace, B., Bocchini, G. M., & Boncio, P. (2014). Do static stress changes of a moderate-magnitude earthquake
760 significantly modify the regional seismic hazard? Hints from the L’Aquila 2009 normal-faulting
761 earthquake (Mw 6.3, central Italy). *Terra Nova*, *26*(6), 430-439. <https://doi.org/10.1111/ter.12117>

762 Pace, B., Visini, F., & Peruzza, L. (2016). FiSH: MATLAB Tools to Turn Fault Data into Seismic-Hazard Models.
763 *Seismological Research Letters*, *87*(2A), 374–386. <https://doi.org/10.1785/0220150189>

764 Pagani, M., Monelli, D., Weatherill, G., Danciu, L., Crowley, H., Silva, V., et al. (2014). OpenQuake Engine: An
765 Open Hazard (and Risk) Software for the Global Earthquake Model. *Seismological Research Letters*,
766 85(3), 692–702. <https://doi.org/10.1785/0220130087>

767 Palumbo, L., Benedetti, L., Bourlès, D., Cinque, A., & Finkel, R. (2004). Slip history of the Magnola fault
768 (Apennines, Central Italy) from 36Cl surface exposure dating: Evidence for strong earthquakes over
769 the Holocene. *Earth and Planetary Science Letters*, 225(1-2), 163-176.
770 <https://doi.org/10.1016/j.epsl.2004.06.012>

771 Pantosti, D., D'Addezio, G., & Cinti, F. R. (1996). Paleoseismicity of the Ovindoli-Pezza fault, central
772 Apennines, Italy: A history including a large, previously unrecorded earthquake in the Middle Ages
773 (860-1300 A.D.). *Journal of Geophysical Research: Solid Earth*, 101(B3), 5937-5959.
774 <https://doi.org/10.1029/95jb03213>

775 Papanikolaou, I. D., Roberts, G. P., & Michetti, A. M. (2005). Fault scarps and deformation rates in Lazio-
776 Abruzzo, Central Italy: Comparison between geological fault slip-rate and GPS data. *Tectonophysics*,
777 408(1-4), 147-176. <https://doi.org/10.1016/j.tecto.2005.05.043>

778 Perras, M. A., & Diederichs, M. S. (2014). A Review of the Tensile Strength of Rock: Concepts and Testing.
779 *Geotechnical and Geological Engineering*, 32, 525-546. <https://doi.org/10.1007/s10706-014-9732-0>

780 Peruzza, L., & Pace, B. (2002). Sensitivity analysis for seismic source characteristics to probabilistic seismic
781 hazard assessment in central Apennines (Abruzzo area). *Bollettino di Geofisica Teorica ed Applicata*,
782 43(1–2), 79–100.

783 Peruzza, L., Pace, B., & Visini, F. (2011). Fault-based earthquake rupture forecast in Central Italy: Remarks
784 after the L'Aquila Mw 6.3 Event. *Bulletin of the Seismological Society of America*, 101, 404-412.
785 <https://doi.org/10.1785/0120090276>

786 Pons-Branchu, E., Douville, E., Roy-Barman, M., Dumont, E., Branchu, P., Thil, F., Frank, N., Bordier, L., &
787 Borst, W. (2014). A geochemical perspective on Parisian urban history based on U-Th dating, laminae
788 counting and yttrium and REE concentrations of recent carbonates in underground aqueducts:
789 *Quaternary Geochronology*, 24, 44–53. doi:10.1016/j.quageo .2014 .08.001.

790 Postpischl, D., Agostini, S., Forti, P., & Quinif, Y. (1991). Palaeoseismicity from karst sediments: the 'Grotta
791 del Cervo' cave case study (Central Italy). *Tectonophysics*, 193, 33-44.

792 Rajendran, C.P., Sanwal, J., Morell, K., Sandiford, M., Kotlia, R.S., Hellstrom, J. & Rajendran, K. (2016).
793 Stalagmite growth perturbations from the Kumaun Himalaya as potential earthquake recorders.
794 *Journal of Seismology*, 20, 579-594.

795 Reimer, P. J., Bard, E., Bayliss, A., Beck, J. W., Blackwell, P. G., Ramsey, C. B., et al. (2013). IntCal13 and
796 Marine13 radiocarbon age calibration curves 0–50,000 years cal BP. *Radiocarbon*, 55(4), 1869–1887.
797 https://doi.org/10.2458/azu_js_rc.55.16947. Roberts, G. P., & Michetti, A. M. (2004). Spatial and
798 temporal variations in growth rates along active normal fault systems: An example from The Lazio-

799 Abruzzo Apennines, central Italy. *Journal of Structural Geology*, 26(2), 339–376.
800 [https://doi.org/10.1016/S0191-8141\(03\)00103-2](https://doi.org/10.1016/S0191-8141(03)00103-2)

801 Roberts, G. P., & Michetti, A. M. (2004). Spatial and temporal variations in growth rates along active normal
802 fault systems: An example from the Lazio-Abruzzo Apennines, central Italy. *Journal of Structural*
803 *Geology*, 26(2), 339–376. [https://doi.org/10.1016/S0191-8141\(03\)00103-2](https://doi.org/10.1016/S0191-8141(03)00103-2)

804 Rossignol-Strick, M. (1999). The Holocene climatic optimum and pollen records of sapropel 1 in the eastern
805 Mediterranean, 9000–6000 BP. *Quaternary Science Review*, 4-5, 515–530.

806 Rovida, A., Locati, M., Camassi, R., Lolli, B., & Gasperini, P. (2016). CPTI15, the 2015 version of the
807 parametric catalogue of Italian earthquakes. Istituto Nazionale di Geofisica e Vulcanologia.
808 <https://doi.org/10.6092/INGV.IT-CPTI15>

809 Rovida, A., Locati, M., Camassi, R., Lolli, B., & Gasperini, P. (2020). The Italian earthquake catalogue CPTI15.
810 *Bulletin of Earthquake Engineering*, 18, 2953–2984. <https://doi.org/10.1007/s10518-020-00818-y>

811 Salvi, S., Cinti, F. R., Colini, L., D'Addezio, G., Doumaz, F., & Pettinelli, E. (2003). Investigation of the active
812 Celano-L'Aquila fault system, Abruzzi (central Apennines, Italy) with combined ground-penetrating
813 radar and palaeoseismic trenching. *Geophysical Journal International*, 155(3), 805–818.
814 <https://doi.org/10.1111/j.1365-246X.2003.02078.x>

815 Schlagenhauf, A., Gaudemer, Y., Benedetti, L., Manighetti, I., Palumbo, L., Schimmelpfennig, I., et al. (2010).
816 Using in situ Chlorine-36 cosmonuclide to recover past earthquake histories on limestone normal fault
817 scarps: A reappraisal of methodology and interpretations. *Geophysical Journal International*, 182(1),
818 36–72. <https://doi.org/10.1111/j.1365-246X.2010.04622.x>

819 Schlagenhauf, A., Manighetti, I., Benedetti, L., Gaudemer, Y., Finkel, R., Malavieille, J., & Pou, K. (2011).
820 Earthquake supercycles in Central Italy, inferred from ³⁶Cl exposure dating. *Earth and Planetary*
821 *Science Letters*, 307(3–4), 487–500. <https://doi.org/10.1016/j.epsl.2011.05.022>

822 Scotti, O., Visini, F., Benedetti, L., Boncio, P., Faure Walker, J., Pace, B., Peruzza, L., & Roberts, G. (2020). On
823 the importance of fault modelling for seismic risk estimate. EGU2020-19434,
824 <https://doi.org/10.5194/egusphere-egu2020-19434>

825 Serva, L., Blumetti, A. M., & Michetti, A. M. (1986). Gli effetti sul terreno del terremoto del Fucino (13
826 Gennaio 1915); tentative di interpretazione della evoluzione tettonica recente di alcune strutture.
827 *Mem. Soc. Geol. It.*, 35, 893–907.

828 Stuiver, M., Reimer, P.J., & Reimer, R.W. (2020). CALIB 7.1 [WWW program] at <http://calib.org>, accessed
829 2020-7-26.

830 Szeidovitz, G., Paskaleva, I., Gribovszki, K., Kostov, K., Surány, G., Varga, P., & Nikolov, G. (2008). Estimation
831 of an upper limit on prehistoric peak ground acceleration using the parameters of intact speleothems
832 in caves situated at the western part of Balkan Mountain Range, North-West Bulgaria. *Acta*

833 *Geodaeica et Geophysica Hungarica*, 43(2-3), 249-266. <https://doi.org/10.1556/AGeod.43.2008.2->
834 3.13

835 Talma, A. S., & Vogel, J. C. (1992). Late Quaternary paleotemperatures derived from a speleothem from
836 Cango Caves, Cape Province, South Africa. *Quaternary Research*, 37(2), 203-213.
837 [https://doi.org/10.1016/0033-5894\(92\)90082-T](https://doi.org/10.1016/0033-5894(92)90082-T)

838 Tesson, J., Pace, B., Benedetti, L., Visini, F., Delli Rocoli, M., Arnold, M., et al. (2016). Seismic slip history of
839 the Pizzalto fault (central Apennines, Italy) using in situ-produced ³⁶Cl cosmic ray exposure dating and
840 rare earth element concentrations. *Journal of Geophysical Research: Solid Earth*, 121(3), 1983-2003.
841 <https://doi.org/10.1002/2015JB012565>

842 Trivedi, A. (2013). Estimating in Situ deformation of rock masses using a hardening parameter and RQD.
843 *International Journal of Geomechanics*, 13(4). [https://doi.org/10.1061/\(ASCE\)GM.1943-5622.0000215](https://doi.org/10.1061/(ASCE)GM.1943-5622.0000215)

844 Valentini, A., Visini, F., & Pace, B. (2017). Integrating faults and past earthquakes into a probabilistic seismic
845 hazard model for peninsular Italy. *Natural Hazards and Earth System Sciences*, 17(11), 2017–2039.
846 <https://doi.org/10.5194/nhess-17-2017-2017>

847 Valentini, A., Pace, B., Boncio, P., Visini, F., Pagliaroli, A., & Pergalani, F. (2019). Definition of Seismic Input
848 From Fault-Based PSHA: Remarks After the 2016 Central Italy Earthquake Sequence. *Tectonics*, 38(2).
849 <https://doi.org/10.1029/2018TC005086>

850 Valentini, A., Duross, C. B., Field, E. H., Gold, R. D., Briggs, R. W., Visini, F., & Pace, B. (2020). Relaxing
851 segmentation on the Wasatch fault zone: Impact on seismic hazard. *Bulletin of the Seismological*
852 *Society of America*, 110(1), 83-109. <https://doi.org/10.1785/0120190088>

853 Verdecchia, A., Pace, B., Visini, F., Scotti, O., Peruzza, L., & Benedetti, L. (2018). The Role of Viscoelastic
854 Stress Transfer in Long-Term Earthquake Cascades: Insights After the Central Italy 2016–2017 Seismic
855 Sequence. *Tectonics*, 37(10), 3411-3428. <https://doi.org/10.1029/2018TC005110>

856 Visini, F., Valentini, A., Chartier, T., Scotti, O., & Pace, B. (2020). Computational tools for relaxing the fault
857 segmentation in probabilistic seismic hazard modelling in complex fault systems. *Pure Applied*
858 *Geophysics*, 177(5), 1855–1877.

859 Wathelet, M., Chatelain, J.-L., Cornou, C., Di Giulio, G., Guillier, B., Ohrnberger, M., & Savvaidis, A. (2020).
860 Geopsy: A User-Friendly Open-Source Tool Set for Ambient Vibration Processing. *Seismological*
861 *Research Letters*, 91(3), 1878-1889. doi:10.1785/0220190360.

862

863 **Figures and Tables captions**

864

865 Figure 1 - Tectonic sketch of the studied area. The star represents the Cola cave (coordinates in the box);
866 the black circles represents the epicenters of the historical earthquakes reported in the CPTI15 catalogue
867 (Rovida et al., 2016); the black lines are the major active faults outcropping in the area (in red the active
868 faults modelled in this work). The numbers represent the modelled seismogenic sources: 1: “*Campo Felice-*
869 *Ovindoli*”; 2: “*Fucino*”; 3: “*Salto Valley*”; 4: “*Velino*”; 5: “*Liri*”. *MF*: Magnola fault; *MHP*: Marsicano Highway-
870 Parasano fault; *SBGM*: San Benedetto-Gioia dei Marsi fault. The hillshade has been produced from a DEM
871 available at: <http://www.sinanet.isprambiente.it/it/sia-ispra/download-mais/dem20/view>.

872

873 Figure 2 – Map of the Cola Cave (modified from a map kindly provided by Daniele Berardi) showing the
874 falling azimuth of broken stalagmites (arrows in map gives the trend; rose diagram distribution in inset
875 provides the direction), the location and labels of dated samples and the seismic noise measurement sites
876 (stars).

877

878 Figure 3 – Conceptual model showing the process of seismic breaking of the speleothems. a) An earthquake
879 occurs in the vicinity of the cave and the speleothems oscillate due to the passage of the seismic waves. b)
880 The speleothem oscillate due to the ground motion until its failure tensile stress is overcome and the
881 speleothem breaks in the weakest part, often where the C axes of calcite crystals could change in
882 orientation from horizontal at the base to vertical above. The speleothems fall down oriented with the
883 mean direction of the seismic waves. c) After the collapse, new speleothems grow on top of the base and
884 the fallen part of the broken speleothem. Dating the tip of the fallen part (A) and the base of the new
885 speleothems (B) help to constrain the age of the collapse. d) If the speleothem is not vulnerable enough,
886 instead of collapsing, it could change its growth axis. In this case the evidences of earthquakes are in the
887 inner structure of it.

888

889 Figure 4 – Speleothem from the Cola cave: (a) site of the two fallen stalactites COLA-20 and COLA-21: in the
890 two small pictures the sample COLA-20 in details, before the cut (on the right) and after the cut (on the
891 left), in red the sampled layer; (b) site of TRAP-1 and TRAP-2 samples: TRAP-1 is the younger stalagmite
892 grown upon the fallen stalagmite TRAP-1; in the small picture the sample TRAP-1 in details after the cut, in
893 red the sampled youngest layer; (c) COLA-3 site: COLA-1 and COLA-2 are the two small stalagmites grown
894 above the stump left by the collapsed stalagmite COLA-3; in the small picture a detail of the sample COLA-2
895 after the cut, with the arrow the sampled oldest layer; (d) CRLI sample site: CRLI is a small stalagmite
896 grown upon a large limestone block fallen from the vault of the cave; in the two small pictures the sample
897 CRLI in details, before the cut (on the right) and after the cut (on the left), in red the sampled layer.

898

899 Figure 5 – Sketch of the real geometry of COLA-3 and of its approximation in the simple and complex
900 modeling. The dashed red line in the real case is the failure section, where COLA-3 broke after the
901 earthquake. The 0.34 meters average section (D) is the section used for the simple modeling and for the
902 central part of the complex modeling. The black dots in the simple and complex modeling are the nodes of
903 the finite elements, which are connected by a beam. The difference between simple and complex shape are
904 at the base and at the tip, where the complex model tries to approximate the real shape of COLA-3.

905

906 Figure 6 – An example of the uniform hazard spectra computed for a given source and the logic tree used to
907 explore the epistemic uncertainty. For each seismogenic source, we computed the $M_{max} - sD$ and the $M_{max} +$
908 sD . Then, for each moment magnitude, we run 10,000 different simulations for the two GMPEs and
909 averaged them giving a weight of 0.5 to each relationship. Finally, we computed for the 10,000 averaged
910 simulations, the median, and the confidence interval at 16% and 84%. At the end, for the modeling
911 simulations and for each seismogenic source, we picked out the two UHS extremes (dashed/bold curves).

912

913 Figure 7 – Deterministic UHS, for each seismogenic source, used in the finite element modeling. For each
914 source, we selected two UHS as described in section 4.3.

915

916 Figure 8 – Chronological scheme of the speleoseismic events recorded in Cola cave by crossing individual
917 events found in the cave. The different symbols indicate the different dating techniques and the thick bars
918 indicate the age uncertainties (Table 1). The gray bands indicate events constrained by both pre-event and
919 post-event layers (the darker grey indicate the most probable ages). Event bracketed with dotted line
920 indicates uncertainty in the pre-event or post-event age. In the upper part we show, as comparison, the
921 paleoseismological data (trench data from Galadini & Galli, 1999; Galli et al., 2008, 2012; and ^{36}Cl
922 cosmogenic data from Schlagenhauf et al., 2010, 2011; Palumbo et al., 2004) available for the “Fucino”
923 source; with the dotted circles we show the possible matching events.

924

925 Figure 9 – H/V spectral ratios for the three different sites in Cola cave. The panels on the right show the
926 average values of the spectral ratios with the relative standard deviations. The figure in the middle panels
927 show the results of H/V rotated in the horizontal plane, i.e., as a function of azimuth. In the right side, the
928 averaged Fourier Amplitude Spectra (FAS) for the three components of motion (blue curve refers to the
929 vertical component). Each panel shows the code associated to the corresponding seismic stations (top
930 right).

931

932 Figure 10 – Results of the simple modeling for the five seismogenic sources. For each source, we computed
933 the failure tensile strength (σ) along the speleothem axis for four different cases in order to explore the

934 epistemic uncertainty in the modeling parameters (UHS and the density ρ). The dotted horizontal lines are
935 the maximum and minimum threshold value of the failure tensile strength, beyond which the speleothem
936 could collapse.

937

938 Figure 11 – Results of the complex modeling for the two seismogenic sources that showed higher results in
939 the simple modeling (Figure 10). Different from the simple modeling, here, the modeled failure tensile
940 strength (σ) reaches the highest value at 0.2 meters, the same point where the speleothem is collapsed
941 (Figure 5).

942

943 Table 1 – Uranium-series measurements of the speleothem samples from the Cola cave. The
944 concentrations of ^{238}U and ^{232}Th were determined using the enriched ^{236}U and ^{229}Th isotopes, respectively.
945 $\delta^{234}\text{Um} = \{[(^{234}\text{U}/^{238}\text{U})_{\text{sample}} / (^{234}\text{U}/^{238}\text{U})_{\text{eq}}] - 1\} \times 1000$, where $(^{234}\text{U}/^{238}\text{U})_{\text{sample}}$ is the measured atomic ratio
946 and $(^{234}\text{U}/^{238}\text{U})_{\text{eq}}$ is the atomic ratio at secular equilibrium. *Corrected ages for detrital ^{230}Th fraction were
947 calculated using an initial $^{230}\text{Th}/^{232}\text{Th}$ activity ratio of 1.25 ± 0.62 .

948

949 Table 2 – AMS ^{14}C ages of the speleothem samples from Cola cave. The “dead carbon portion” DCP-
950 corrected ^{14}C ages were calculated based on two different values of dead carbon incorporated in the
951 speleothem (5%, and 10%). The ^{14}C ages were converted to calendar year (calibrated age) using the
952 IntCal13 calibration curve (Reimer et al., 2013). The mean preferred calibrated age is the average of the
953 median probability calculated from the two DCP values. The 1σ range for the preferred calibrated age
954 represents the largest age interval.

955

956 Table 3 – Geometric parameters and list of M_{max} and its standard deviation (sD) of the Seismogenic Sources.
957 Id = Identification number, L = along strike length, Dip = the inclination angle of the fault plane.

958

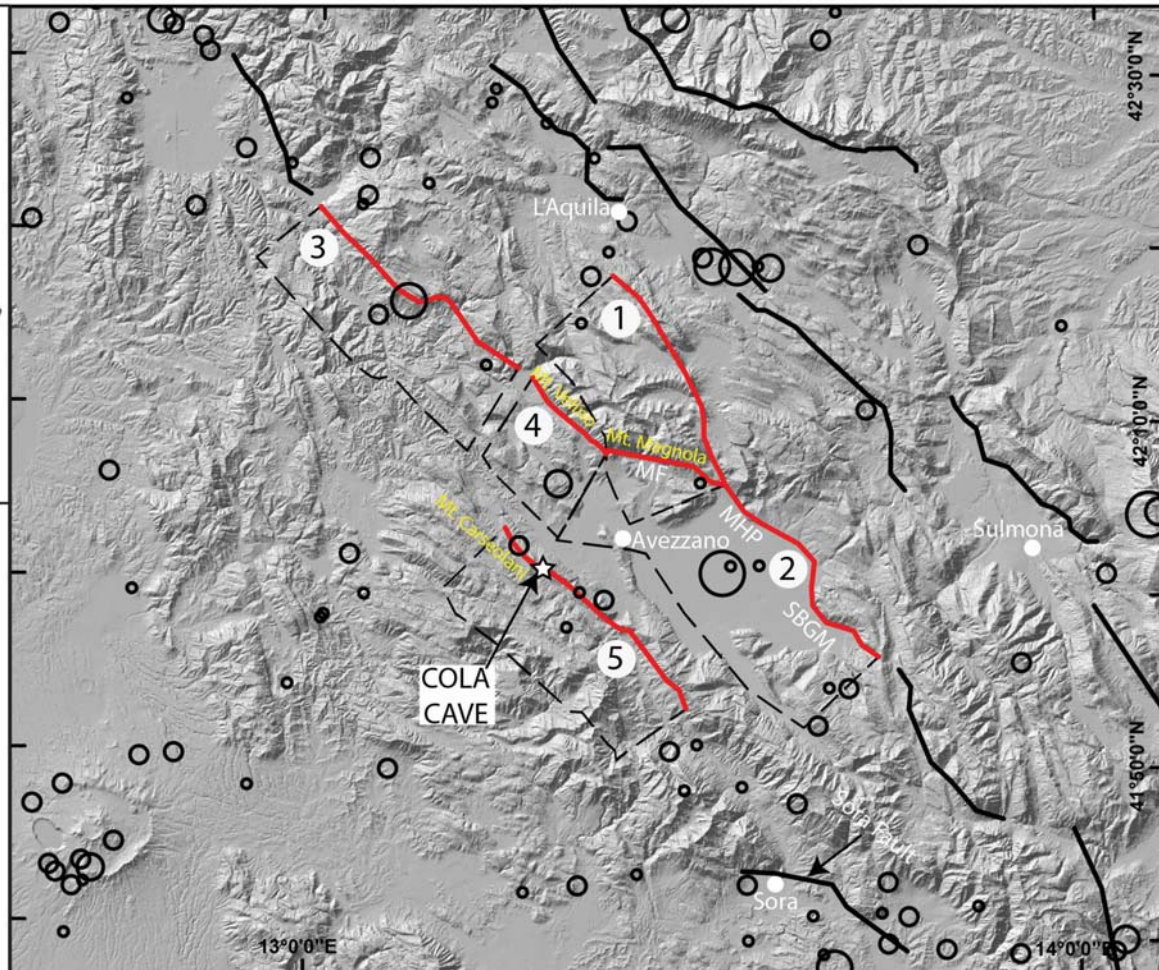
959 Table 4 – Geometrical features, physical and mechanical parameters of the investigated speleothems: H =
960 Height (in-site average value over the failure section), D = Diameter (in-site average value of the failure
961 section), δ = density, Ed = dynamic young modulus, σ_T = tensile strength thresholds.

962

963 Table 5 – Age constraints on speleoseismic events from U-Th and ^{14}C dating results.

964

Figure 1.



— Active Fault
— Active Fault (modeled)

CPTI15

- Mw < 4.5
- 4.5 < Mw < 5.5
- 5.5 < Mw < 6.0
- 6.0 < Mw < 6.5
- Mw > 6.5



Figure 2.

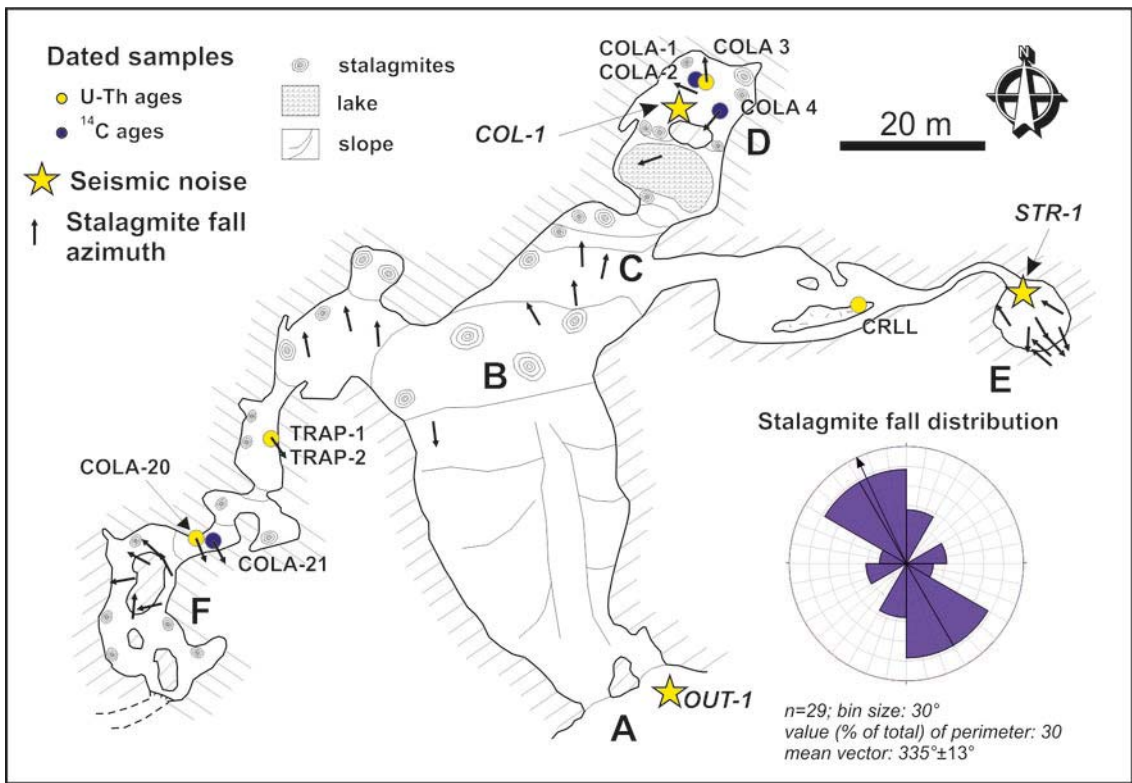


Figure 3.

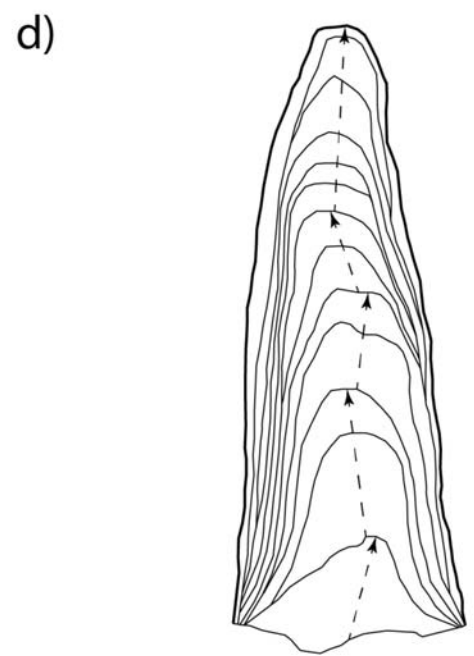
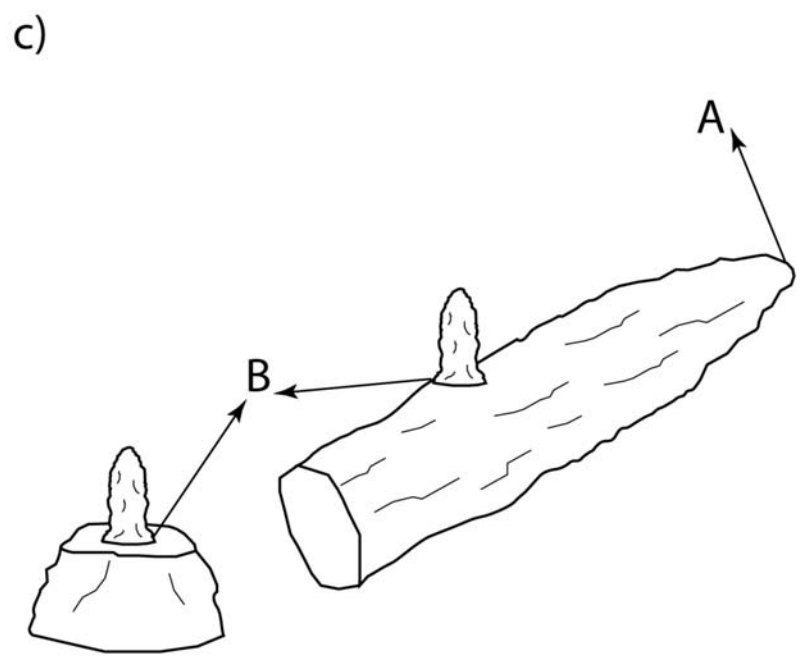
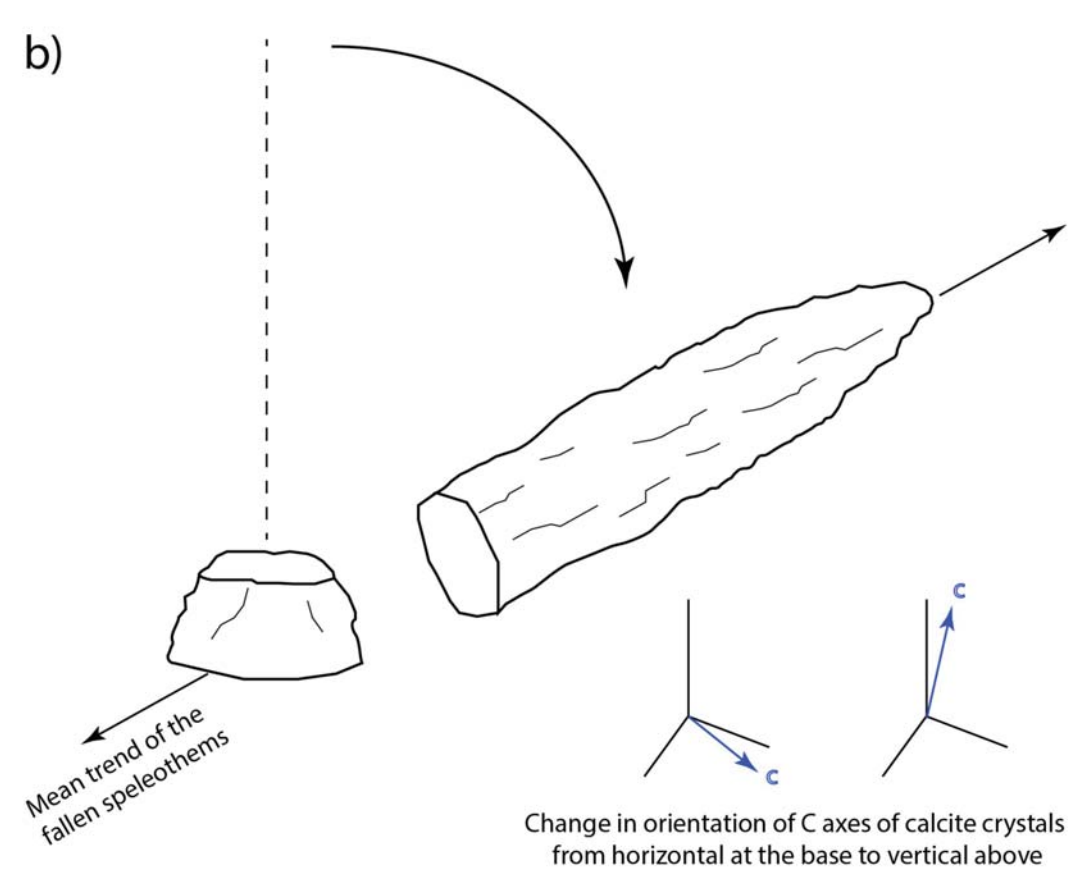
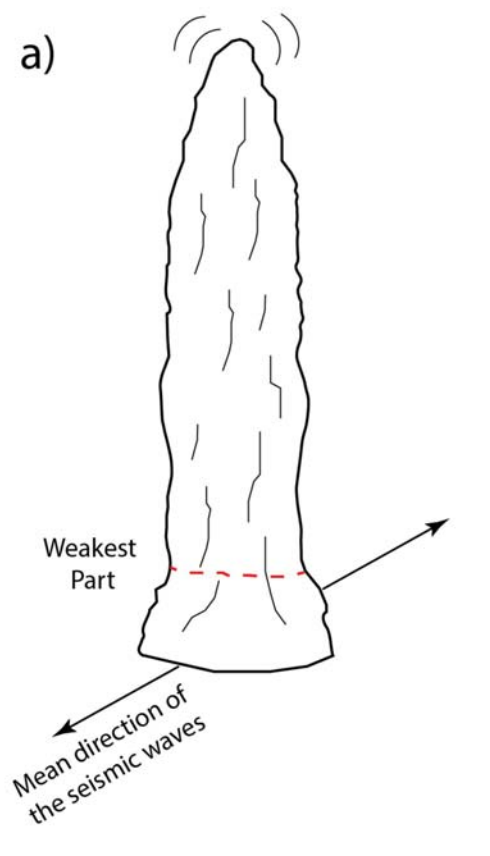


Figure 4.

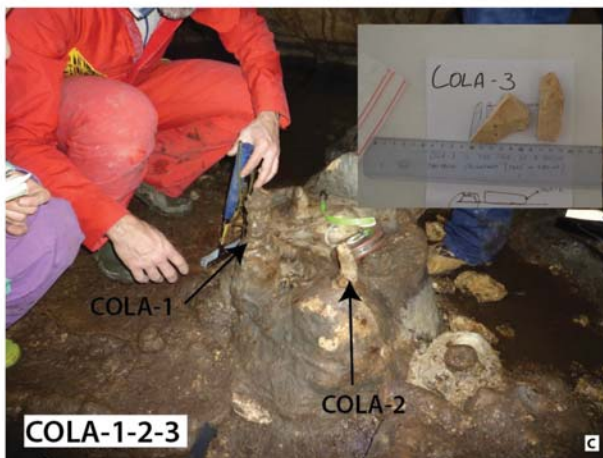
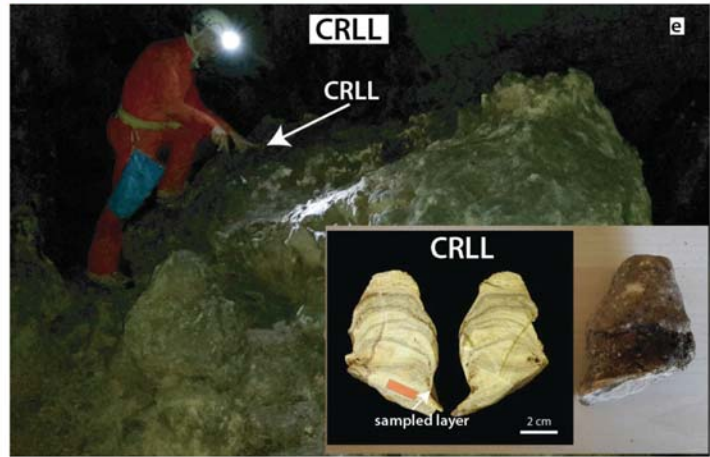
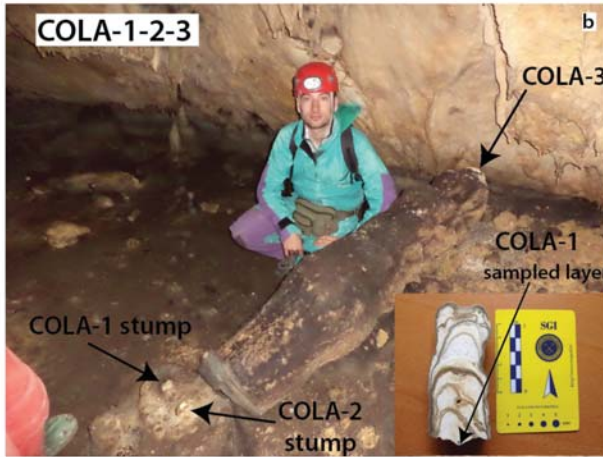
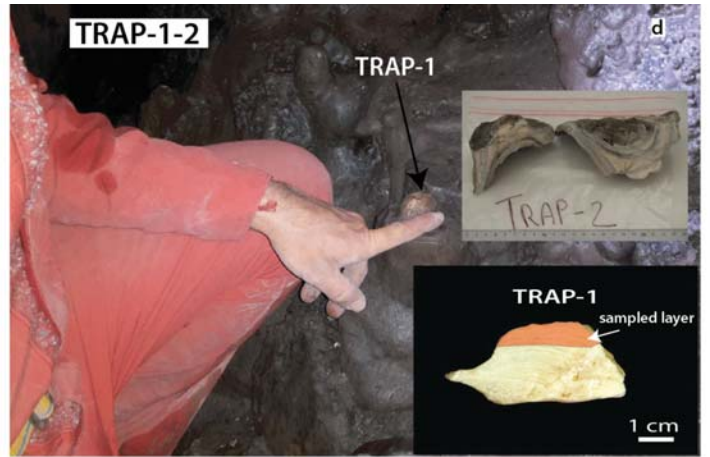


Figure 5.

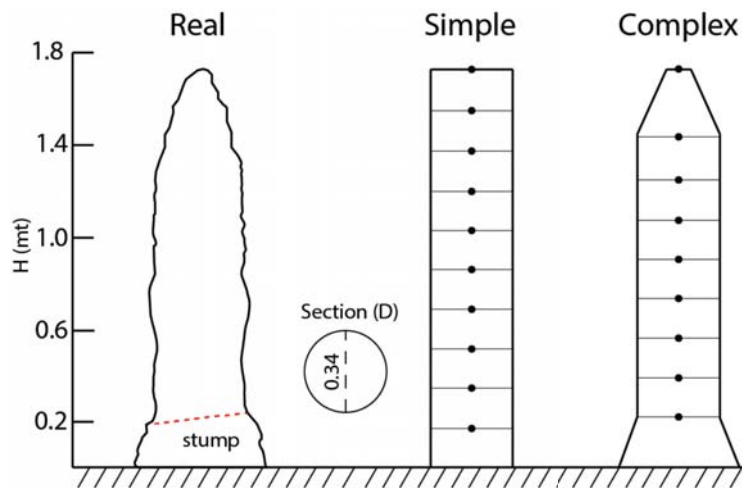


Figure 6.

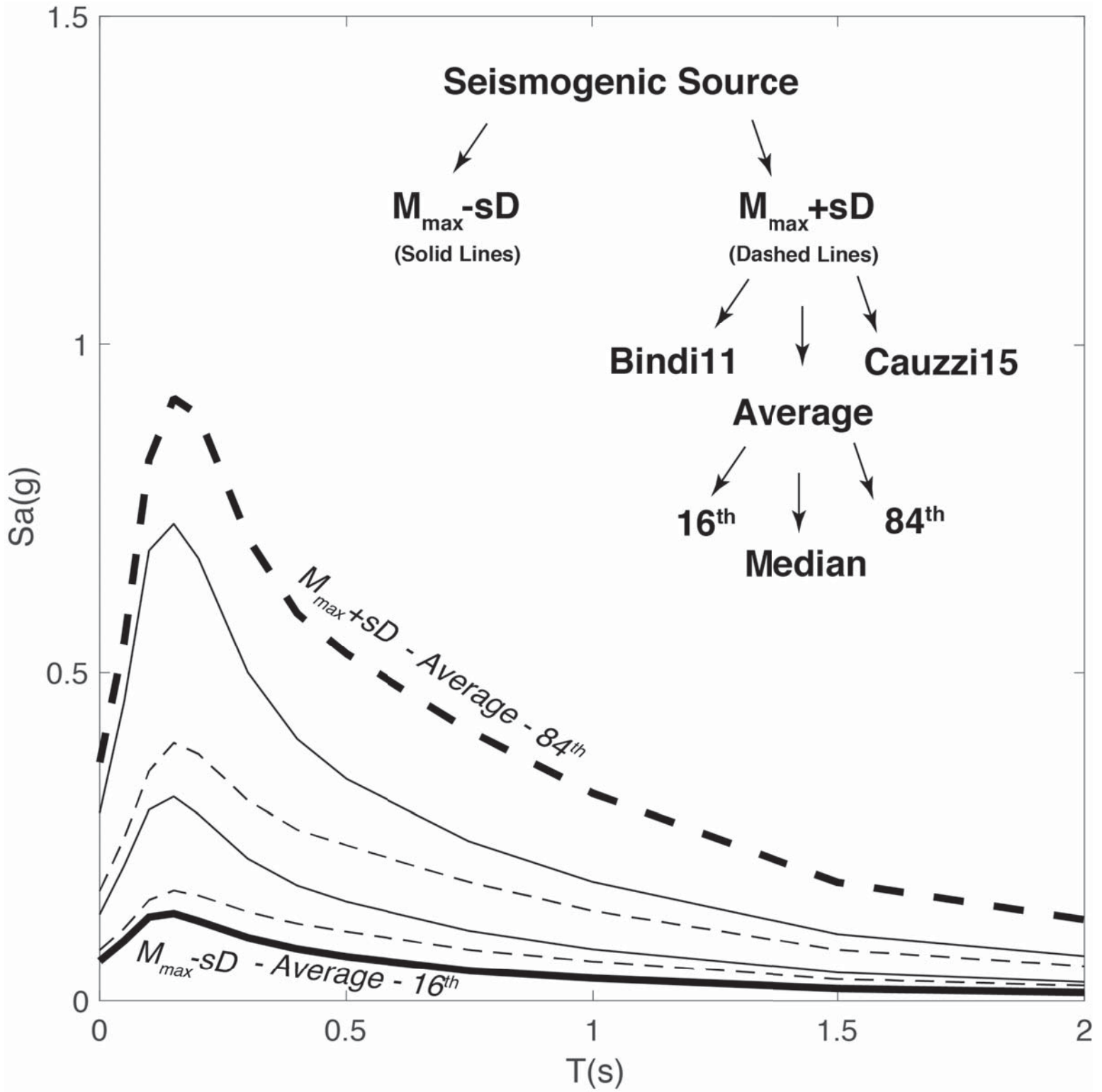


Figure 7.

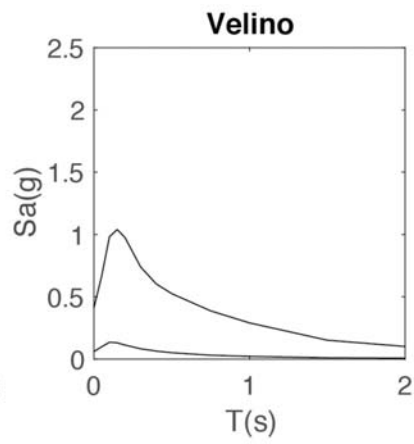
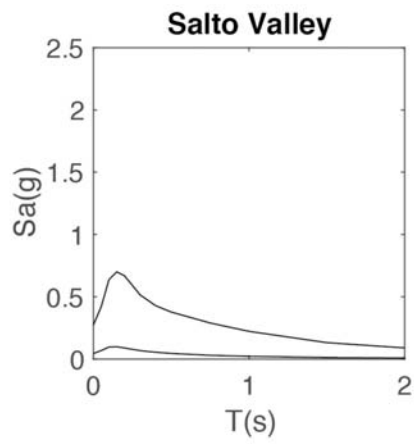
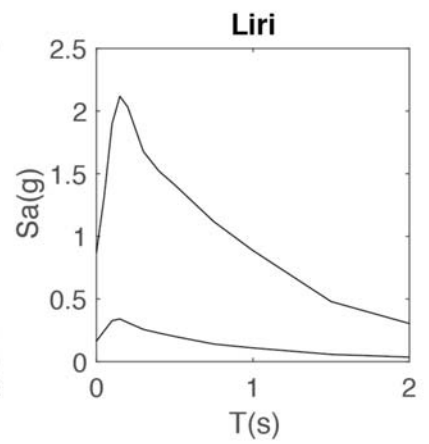
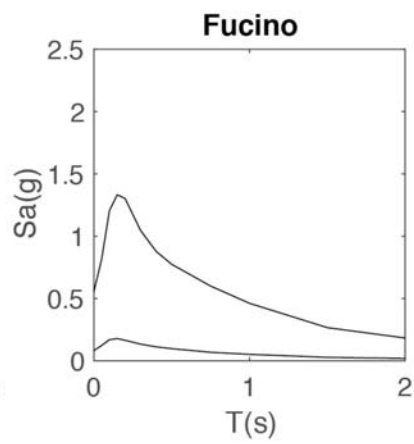
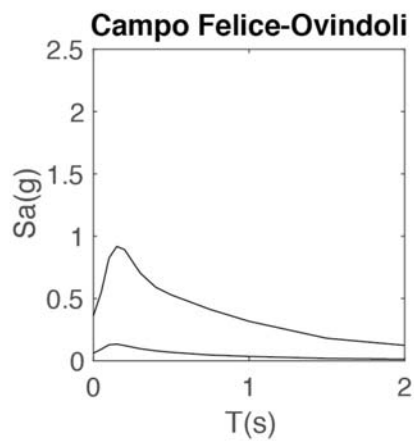


Figure 8.

COLA CAVE
FUCINO

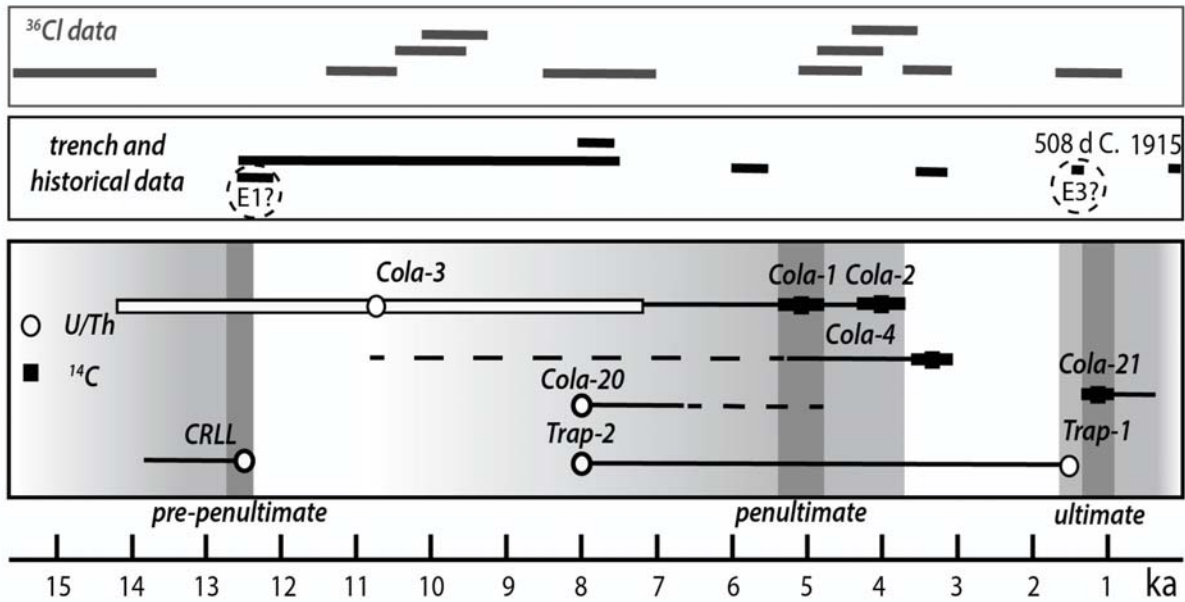


Figure 9.

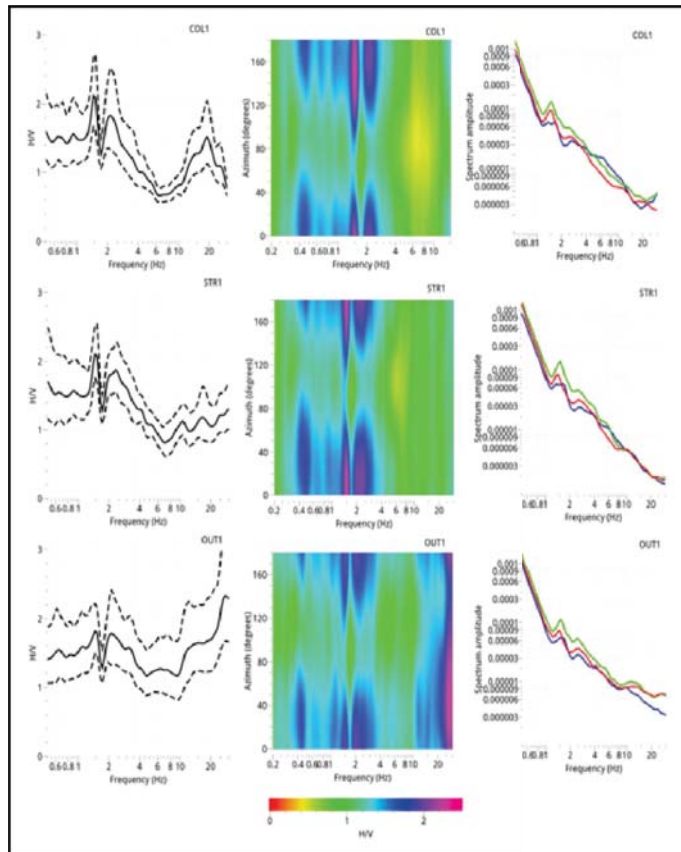


Figure 10.

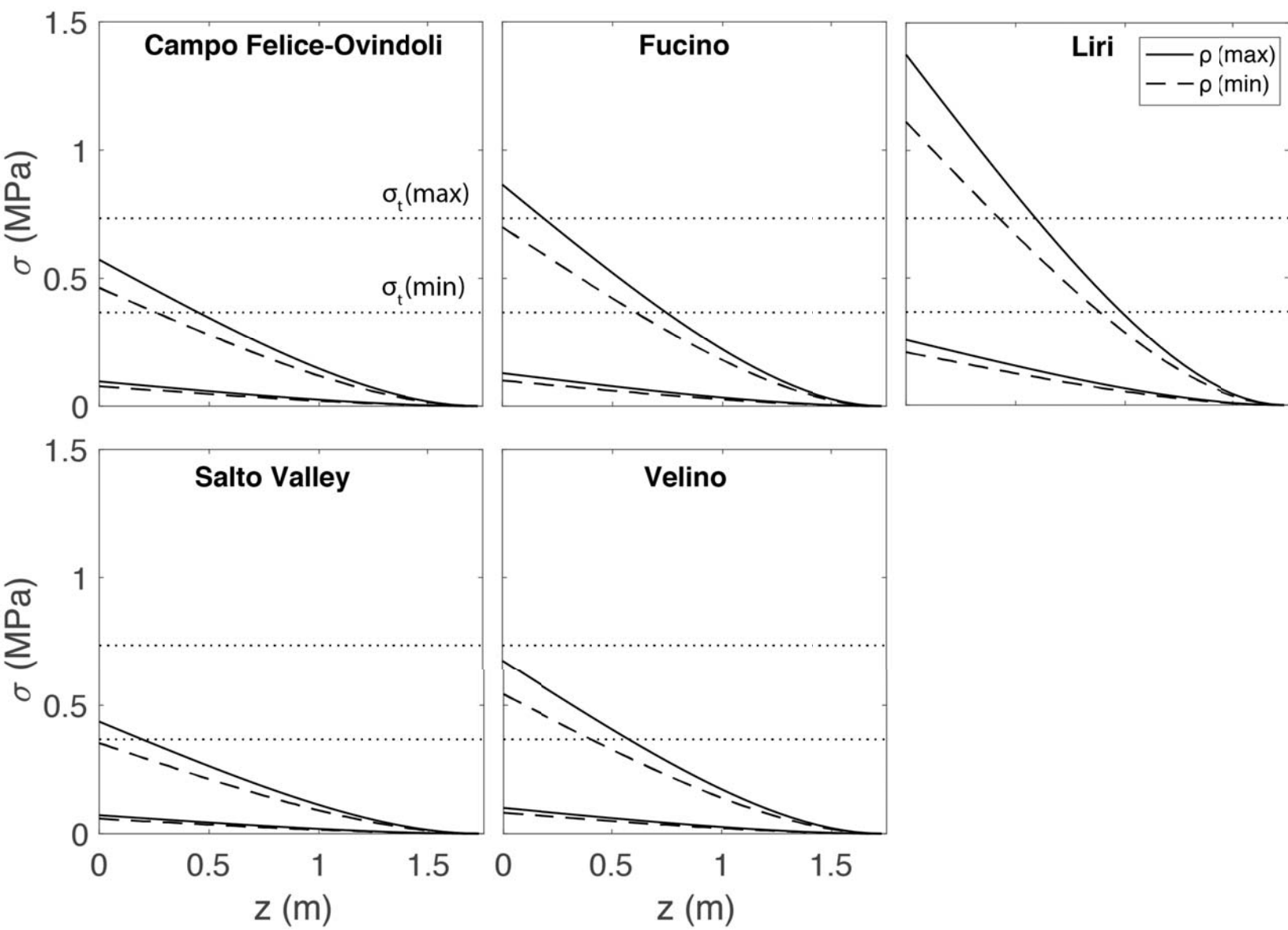


Figure 11.

