

1 **An energy-dependent earthquake**
2 **moment-frequency distribution**

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Abstract

The magnitude-frequency distribution (MFD) of many earthquake catalogs is well described by the Gutenberg-Richter (GR) law, or its tapered version (TGR). This distribution is usually extrapolated to any subsets of the space-time window covered by the catalog. However, some empirical observations and logical thoughts may raise doubts about the validity of this extrapolation. For example, according to the elastic rebound theory, we may assert that the probability of a strong shock to nucleate within a short time-interval in a small area \mathcal{A} just ruptured by another strong event, should be lower than that expected by GR (or TGR): a lot of energy has already been released, and it takes time to recover to the previous state. Here we put forward a space-time modification of the TGR, named TGRE (energy-dependent TGR), where the corner seismic moment becomes a time-varying energy function depending on: i) the conceivable strongest shock that may nucleate in \mathcal{A} ; ii) the time elapsed since the last strong earthquake resetting the elastic energy in \mathcal{A} to a residual value; iii) the rate of the energy recovering, linked to the recurrence time of the fault(s) involved. The model also verifies an invariance condition: for large space-time windows the occurrence of a strong shock doesn't affect significantly the whole elastic energy available, i.e., the TGRE becomes the TGR. The model is simple and rooted in clearly stated assumptions. To evaluate its reliability and applicability, we apply it to the Landers sequence in 1992. As expected by TGRE, we find that the MFD close to the fault system interested by the mainshock (Mw7.3) differs from that of earthquakes off-fault, showing a lower corner magnitude. We speculate that TGRE may be profitably used in operational earthquake forecasting, and explains the empirical observation that strongest aftershocks nucleate always outside the

33 **Introduction**

34 The Gutenberg-Richter (GR) law (Gutenberg and Richter (1944)) and its tapered
35 version (TGR) (Kagan (2002a,b)) are the most used magnitude-frequency distributions
36 (MFD) at quite different space-time windows, such as, for example, in operational
37 earthquake forecasting models (Jordan et al. (2011); Marzocchi et al. (2017); Omi
38 et al. (2018); Michael et al. (2019)). The validity of the (T)GR rests on the assumption
39 that the magnitude of an earthquake is independent from the past seismicity for any
40 dimension of the space-time window. Although this assumption seems appropriate
41 when looking at large spatiotemporal domains, its validity at small space-time scales
42 conflicts with some empirical findings, for which the largest triggered events occur
43 outside the fault of the strong triggering earthquake (van der Elst and Shaw (2015);
44 Stallone and Marzocchi (2019)).

45 Conceptually, this empirical observation could be explained in the framework of
46 the elastic rebound theory (Reid (1911)), for which one strong earthquake decreases
47 significantly the elastic energy available in the fault that generates the shock, and
48 it takes time to recover it. This means that the probability of a strong shock to
49 nucleate in the same area where another strong earthquake just occurred within a
50 short-time window, has to be lower than that predicted by the (T)GR law. Conversely,
51 if we consider a larger spatial scale, the occurrence of a single shock does not affect
52 significantly the elastic energy available in the area, so it is expected that the (T)GR
53 keeps holding. Besides the empirical evidence, we notice that the existence of a possible

54 variability of MFD stems from recent operational earthquake forecasting models (Field
55 et al. (2017a,b)), based on faults system which can produce reliable forecasts only when
56 the MFD is changed in space.

57 In this paper we put forward a space-time dependent model, which describes the
58 MFD of earthquakes that nucleate in small space-time areas, taking into consideration
59 the elastic energy released by the past seismicity in that area. The use of a small space-
60 time dimension marks the difference with very recent studies on a similar argument
61 (Marsan and Tan (2020)), and with past analyses on the definition of the maximum
62 magnitude expected in fixed (long) time windows (Zöller et al. (2013)). The model
63 introduces a time-varying corner seismic moment in the TGR law, which results from
64 the level of elastic energy that is currently available to be released in the space-time
65 area of interest. We name the model TGRE to explicitly reflect the dependence of
66 the MFD on the elastic energy available. In a nutshell, TGRE inhibits the nucleation
67 of large earthquakes in the area that just experienced a significant release of elastic
68 energy.

69 An alternative approach to model the space-time variability of the elastic energy
70 available is based on quantifying space-time variations of the b -value parameter in the
71 GR law (Gulia and Wiemer (2019)). For instance, a larger b -value diminishes the
72 probability of large earthquakes, but they still remain possible (e.g., if we keep fixed
73 the rate of M4+ earthquakes, increasing the b -value from 1.0 to 1.2 diminishes the
74 M7+ rate of a factor of about 4). Empirical evidence seems to show that this chance is
75 maybe lower, because large aftershocks nucleate almost exclusively in the outer regions
76 of the mainshock zone (van der Elst and Shaw (2015)). The model that we put forward

77 in this study is likely more suitable to explain such empirical evidence.

78 In the first part of this paper we describe the theoretical aspects of the model: we
79 explicitly derive its formulation and that of the time-varying corner seismic moment
80 with respect to which it is conditioned; we also discuss the stability conditions in
81 comparison with that of the classical GR model. In the second part we analyze the
82 Landers earthquake sequence, started on June 28, 1992, with a Mw7.3 event, with a
83 dual purpose: i) to find empirical evidence corroborating the existence of space-time
84 variability of the MFD; ii) to test if the proposed TGRE model better describes the
85 data than the space-time independent TGR model.

86 **The energy-dependent MFD model (TGRE)**

87 For the sake of mathematical simplicity, the TGRE is built in terms of seismic moment
88 instead of magnitude; the transition from one to the other can be easily made by
89 applying Kanamori (1977)'s relationship $m = \frac{2}{3} \log M - 10.73$, where M stays for
90 seismic moment (in dyne \times cm) and m for the corresponding moment magnitude.
91 Such a notation will be adopted in this paper hereafter; furthermore, owing to the
92 unambiguous relationship above, we will use the acronym MFD also for the seismic
93 moment-frequency distribution. The MFD Tapered Pareto GR (TGR) law introduced
94 by Kagan (2002a,b) reads:

$$\Phi_{TGR}(M) = \Phi_{GR}(M) \cdot \exp \left\{ \frac{M_{min} - M}{M_c} \right\}, \quad (1)$$

95 where $\Phi_{GR}(M) = \left(\frac{M}{M_{min}}\right)^{-\beta_k}$ is the GR-distribution, $\beta_k = \frac{2}{3}b$ -value, M_{min} is the com-
 96 pleteness threshold and M_c is the corner seismic moment, which is the value such that
 97 events with a higher seismic moment are less likely than what expected by the de-
 98 creasing exponential distribution. The tail of the GR law is therefore forced to decay
 99 stronger in the TGR model, the decay itself being controlled by the M_c value which
 100 is assumed as a fixed parameter, typically estimated through the maximum likelihood
 101 technique (Kagan and Schoenberg (2001)).

102 In this paper we introduce the TGRE model for earthquakes which nucleate inside
 103 an arbitray portion \mathcal{A} of the fault (the generalization to a volume is straightforward).
 104 The TGRE model relaxes the hypothesis that M_c is a fixed parameter, and it allows
 105 it to vary as a function of the amount of energy E currently available in \mathcal{A} , i.e., $M_c \equiv$
 106 $M_c(E, t)$, where t is the time since the last earthquake which resets the energy in \mathcal{A}
 107 to a minimum value. This function $M_c(E, t)$ has to consider the past earthquakes that
 108 nucleated in \mathcal{A} , as well as the earthquakes that involved \mathcal{A} in their rupture nucleated
 109 somewhere else (we use the term “participation” hereafter as in Parsons et al. (2018)).
 110 In this way, the TGRE model inhibits a second strong shock to nucleate in a small
 111 area that has been involved in a strong earthquake recently, but it does not prevent
 112 this area to participate to the rupture of another big event which may nucleate nearby,
 113 along the same fault(s) involved. It follows that the nucleation MFDs in two nearby
 114 small areas may be different, but still influenced by the reciprocal seismicity. For the
 115 sake of simplicity, hereafter we will omit to specify the dependence on t in the notation
 116 of $M_c(E)$.

117 We also constrain the model to respect a sort of “invariance condition”, i.e., the

118 TGRE turns back to the classical TGR at large spatiotemporal scales. Of course, the
119 specific choice of considering a time-varying corner seismic moment is not the only one
120 possible to introduce an energy-dependence in the MFD, but it is justified in terms of
121 easily practical use and testing; any other way of including a direct dependence on
122 the energy can be proposed, provided that a higher complexity must be worth for a
123 better reliability of the model, and coherence with previous pieces of evidence is needed.
124 In the following subsections we define both the time-varying corner seismic moment
125 $M_c(E)$, and the explicit distribution of the TGRE model.

126 **Time-varying corner seismic moment $M_c(E)$**

127 Here we propose a formulation of $M_c(E)$ based on two main concepts. First, the
128 relevant quantities controlling the earthquake nucleation in \mathcal{A} are: the strongest earth-
129 quake that can eventually nucleate in \mathcal{A} , and the most recent past earthquake which
130 resets the available energy to the residual minimum value. Specifically, the elastic en-
131 ergy in \mathcal{A} is reset when this area participates to an earthquake which nucleates outside,
132 or when an earthquake nucleates inside and generates a fractured area larger than \mathcal{A} .
133 In other words, the resetting event must have a seismic moment $M \geq M_A$, where M_A
134 is the seismic moment of an earthquake with area equal to \mathcal{A} . To determine if an
135 event has involved this area, we check if at least part of \mathcal{A} falls in the CircleArea with
136 the earthquake epicenter. The relative diameter (as well as M_A) may be computed
137 through any proper RuptureLength-MomentMagnitude relationship such as in Wells
138 and Coppersmith (1994), Papazachos et al. (2004), or Allen and Hayes (2017).

139 Second, the elastic energy available in \mathcal{A} scales with times and it is related to M_c . In

140 elasticity theory, $E \propto \sigma^2$, where E is the elastic energy accumulated as a consequence
 141 of the applied stress σ ; since the stress rate due to plate tectonics can be considered
 142 a constant value (that is, $\sigma \propto t$), it follows that $E \propto t^2$. The link between elastic
 143 energy available and seismic moment is instead more controversial. In a general way,
 144 $E \propto M_c$ holds only if the static stress drop of earthquakes is independent from the
 145 magnitude. This hypothesis is still matter of intense debate (Ide and Beroza (2001);
 146 Kanamori and Brodsky (2004); Oth et al. (2010)). In this paper we assume that elastic
 147 energy available and seismic moment are proportional. At the same time we stress that
 148 a similar TGRE may be built adopting a different form of $M_c(E)$, which takes into
 149 account a different hypothesis.

150 Going into the detail, we define the following parameters.

- 151 a) M_c^* is the maximum corner seismic moment, i.e., the seismic moment of the
 152 conceivable strongest shock that may nucleate in \mathcal{A} . It is actually the corner
 153 seismic moment M_c adopted in the classical TGR distribution (1). We propose
 154 that M_c^* could be related, even though not necessarily, to the length of the longest
 155 fault included in the area: for instance, it can be obtained from any proper
 156 RuptureLength-MomentMagnitude relationship, such as those proposed in Wells
 157 and Coppersmith (1994), Papazachos et al. (2004), or Allen and Hayes (2017).
- 158 b) t_0 is the occurrence time of the earthquake which has reset the elastic energy in
 159 \mathcal{A} , i.e., the past earthquake at which \mathcal{A} participated.
- 160 c) $M_{c,0}^*$ sets the minimum value for the corner seismic moment which is achieved after
 161 the occurrence of a resetting earthquake in \mathcal{A} . In general, $M_{c,0}^* = \rho \cdot M_c^*$, where
 162 $\rho < 1$ indicates the fraction of elastic energy that is available after the resetting

163 event. The value of ρ , or equivalently of $M_{c,0}^*$, may be either set theoretically,
 164 for instance by analyzing the stress rotation (Hardebeck and Okada (2018)), or
 165 empirically, by analyzing one or more stacked similar earthquake sequences.

166 d) ν is a parameter connected to the recurrence time of the longest fault involved
 167 in \mathcal{A} , and it controls the velocity of convergence to the maximum value M_c^* after
 168 a resetting event.

169 In “Application to real earthquakes: the Landers sequence”, subsection “Setting
 170 parameters and assumptions”, we describe some practical choices of these parameters.
 171 Still, we stress again that the choices are not prescriptive for the TGRE’s application;
 172 different $M_c(E)$ parameterizations, assumptions and parameters can be used.

173 According to the above concepts and definitions, we define the time-varying energy
 174 function as

$$M_c(E) = M_{c,0}^* + (M_c^* - M_{c,0}^*) [\nu(t - t_0)]^\alpha \quad (2)$$

175 bounded to the values $(t - t_0) \leq \frac{1}{\nu}$, which translates in $(t - t_0) \leq \tau$ when the coefficient
 176 of variation of the interevent times between consecutive earthquakes (*CoV*) is zero, i.e.,
 177 τ is the recurrence time between earthquakes. This restriction guarantees indeed that
 178 $M_c(E) \in [M_{c,0}^*, M_c^*]$ when $(t - t_0) \in [0, \tau]$, a requirement that is deducible from the
 179 argument above. The dependence on time of the corner seismic moment is therefore
 180 expressed with respect to the time elapsed since the resetting event, and the seismic
 181 moments multiplication term allows us to account for the energy reloading process;
 182 while, $M_{c,0}^*$ is added to ensure that the available energy will not fall below its minimum
 183 value, even immediatly after the resetting event, that is, when $t - t_0 \sim 0$. In this paper,

184 according to the proportionality between elastic energy and seismic moment, we set
185 $\alpha = 2$ ($\alpha = 1$ if the seismic moment is assumed to increase linearly with time).

186 The temporal trend of $M_c(E)$, as well as its sensitivity to the parameters, can be
187 observed in Fig. 1, whose plots are obtained by considering two parameters among
188 $(M_c^*, M_{c,0}^*, \nu)$ fixed, the third varying; for an easier interpretation, we also display
189 magnitude values instead of seismic moments. An overall increasing trend is shown
190 in all the plots. As the intuition suggests, the time-varying corner seismic moment
191 approaches more rapidly its maximum when ν becomes larger: the lower the recurrence
192 time of the fault, the faster M_c^* can be reached. The increasing velocity of $M_c(E)$ is
193 also faster as M_c^* is higher, whereas it does not change with $M_{c,0}^*$. This is because the
194 influence of the latter on the taper's trend can be appreciated only within a short-
195 time interval since the resetting event (less than 1 year in our example), being $M_c(E)$
196 controlled mainly by M_c^* and ν at just larger scales: that's why the x -axes in plot c)
197 are cut at 1 year after the reset, otherwise the difference would not have been visible.
198 When focusing on the entire time window, we observe instead that the influence of ν
199 and M_c^* on $M_c(E)$ is a bit stronger. However, Fig. 2 highlights that in the short-term,
200 the time-varying corner seismic moment does not substantially depend on these two
201 values.

202 To be thorough, we add that $M_c(E)$ could be also interpreted as a random variable
203 whose distribution takes the cue from the stress level adopted in the stress release
204 model (Vere-Jones (1978, 1988); Wang et al. (1991); Zheng and Vere-Jones (1991);
205 Xiaogu and Vere-Jones (1994)). In fact, $M_c(E)$ could consist in a deterministic term
206 of accumulated energy, linked to the elapsed time since the resetting event, and a

207 stochastic term of energy released by each single past earthquake, which is distributed
 208 according to TGR. Nevertheless, so as to gain in easily applicability and reliability
 209 testing, we assume here that $M_c(E)$ is a deterministic function of time, as defined
 210 in (2).

211 **The mathematical description of the TGRE model**

212 The TGRE model we propose for earthquake seismic moments is simply obtained by
 213 including the time-varying corner seismic moment $M_c(E)$ previously derived, into the
 214 TGR cumulative distribution (1), i.e.,

$$\Phi_{TGRE}(M) = \left(\frac{M}{M_{min}} \right)^{-\beta_k} \exp \left\{ \frac{M_{min} - M}{M_c(E)} \right\}, \quad (3)$$

215 where $M_c(E) \in [M_{c,0}^*, M_c^*]$ is defined in (2) with $\alpha = 2$. Fig. 3 shows $\Phi_{TGRE}(M)$ as a
 216 function of $M_c(E)$.

217 If we consider a large spatial domain composed by many faults, and many cells \mathcal{A} ,
 218 the occurrence of one or a few large earthquakes may reset only a limited number of cells
 219 \mathcal{A} . This means that for the whole large spatial domain, $M_c(E) \equiv M_c^*$, acknowledging
 220 the spatial invariance condition. The temporal invariance condition is instead satisfied
 221 by construction in fact, equation (2) gives M_c^* for $t - t_0 \rightarrow \tau$.

222 One obvious application of the TGRE model is in operational earthquake forecasting
 223 (OEF; Jordan et al. (2011)). It is expected to solve the main conundrum of existing
 224 OEF models (Marzocchi et al. (2017); Omi et al. (2018); Michael et al. (2019)), for
 225 which the probability of a large aftershock is exactly where the mainshock occurred.

226 For example, the ETAS-TGRE (i.e., ETAS with TGRE) rate would be

$$\lambda(t, x, y, M|\mathcal{H}_t) = \left[\lambda_0(x, y) + \sum_{\{i|t_i < t\}} \lambda_{tr}(t - t_i, x - x_i, y - y_i; M_i) \right] p_{TGRE}(M|M_c(E)), \quad (4)$$

227 where \mathcal{H}_t is the past history up to time t , i.e., the past earthquakes $\{(t_i, x_i, y_i, M_i); t_i <$
 228 $t\}$; $\lambda_0(x, y)$ is the rate of the background events; $\lambda_{tr}(t - t_i, x - x_i, y - y_i; M_i)$ is the rate
 229 of the triggered events; $p_{TGRE}(M|M_c(E))$ is the TGRE probability density function
 230 for the seismic moment that is calculated in x, y at the time t ; finally, $M_c(E) \equiv$
 231 $M_c(E, x, y, s)$ is linked to the elastic energy available in x, y after a time s since the
 232 last resetting earthquake. In this framework, the TGRE may be applied to both the
 233 background and triggered earthquakes as in the classical ETAS model.

234 In the ETAS-TGRE setting, it is also interesting to investigate how the shift of
 235 the TGRE taper influences the computation of the branching ratio, which we recall
 236 being the average number of aftershocks triggered by an arbitrary event (Zhuang et al.
 237 (2012)). As for the TGR law, the branching ratio of the TGRE model can be derived

238 as

$$\eta_{TGRE} = \kappa + \kappa \alpha_k e^{\frac{M_{min}}{M_c(E)}} \left(\frac{M_{min}}{M_c(E)} \right)^{\beta_k - \alpha_k} \Gamma \left(-\beta_k + \alpha_k, \frac{M_{min}}{M_c(E)} \right), \quad (5)$$

239 where $\Gamma(s, t) = \int_t^\infty x^{s-1} e^{-x} dx$ is the upper incomplete Gamma function (Bateman
 240 (1953); Temme (1996); Spassiani (2020)), and κ, α_k are the parameters of the produc-
 241 tivity law $\varrho(\cdot)$ expressed in terms of the seismic moment through Kanamori (1977)'s

242 relationship, i.e., $\varrho(M) = \kappa \left(\frac{M}{M_{min}} \right)^{\alpha_k}$. In Fig. 4 we show that η_{TGRE} increases with
 243 the time-varying corner seismic moment $M_c(E)$, indicating that if the taper moves
 244 to the left as a consequence of a great amount of energy just released, the average
 245 number of aftershocks triggered by a generic event is reduced: in fact, an event with
 246 a lower seismic moment will generate a lower number of aftershocks. The plot shows
 247 that the increasing behavior is faster as the difference $\beta_k - \alpha_k$ is lower: in the case
 248 of the classical ETAS-GR it has to be $\beta_k > \alpha_k$ for the process not to explode, but
 249 this condition becomes unnecessary for the ETAS-TGRE model, so as for ETAS-TGR
 250 (Spassiani (2020)). As usual, the stability of the ETAS-TGRE process is guaranteed
 251 by imposing $\eta_{TGRE} < 1$; finally, when $\beta_k > \alpha_k$ it holds $\eta_{TGRE} < \eta_{GR}$, therefore in this
 252 case our model's stability conditions are even less restrictive than those of ETAS-GR.

253 **Application to real earthquakes: the Landers sequence**

254 In this section we test the hypothesis of the space and time independence of MFD, and
 255 then we show how the TGRE model works in practice. To do that, we consider the
 256 Landers earthquake sequence, which started with a magnitude Mw7.3 event occurred
 257 on June 28, 1992, in Southern California. The seismic catalog for such a sequence is
 258 rich enough to allow a statistically significant data-model comparison. Furthermore,
 259 the fault segment that generates the initial earthquake is well-defined in this case, as
 260 a detailed mapping of the slip distribution is available: in our analysis we focus on the
 261 fault segments that certainly slipped during the Mw7.3 event, as shown in Madden and
 262 Pollard (2012) and hereafter called “Landers fault”.

263 Seismic data for the analysis have been taken from the online available Uniform Cal-

264 ifornia Earthquake Rupture Forecast, version 3 (UCERF3) earthquake catalog, which
265 covers the entire California Region from July 1769 to April 2010 and includes events
266 with $\text{mag} \geq 4$ before 1894 and $\text{mag} \geq 2.5$ after this year (Field et al. (2013)). The data
267 relative to the Landers fault have instead been taken from the California Reference
268 Fault Parameter Database (CRFPD)-UCERF2 system, that is easily accessible online
269 (Field et al. (2009)) and does not present substantial differences with respect to the
270 UCERF3 regarding the faults involved by the Landers rupture. For the websites, see
271 “Data and Resources”.

272 In particular, in this application we test whether the MFDs inside and outside the
273 rupture that generates the Landers earthquake come from the same distribution, as it
274 would be expected in the case of space-time independence. Then, we apply the TGRE
275 model to the on-rupture earthquakes, and we quantify the difference of the reliability
276 of the TGR and TGRE models through the log-likelihood ratio test. The red stripe in
277 panels *a*) of figures from 5 to 8 shows the rupture on the Landers fault. The stripe has
278 a thickness of about 10 km (considering ± 5 km around the latitude of each segment
279 fault point). The analysis is conducted in the following four time intervals: 29 June -
280 6 July 1992, 29 June - 29 July 1992, 29 June - 29 September 1992, 29 June 1992 - 29
281 June 1993, that is, respectively 1 week, 1 month, 3 months and 1 year since the day
282 after the Mw7.3 resetting event.

283 **Setting parameters and assumptions**

284 The first step is to define \mathcal{A} , which sets the spatial resolution of the analysis. We
285 consider the case in which \mathcal{A} covers the whole fault rupture of the Mw7.3 earthquake

286 (red stripes in panels *a*) of figures from 5 to 8). For this tutorial application we set the
287 parameters of the TGRE model as follows.

288 i) M_c^* corresponds to $m_c^* = 7.59$, as proposed in Kagan et al. (2010) for active
289 continents.

290 ii) After the resetting Landers earthquake, no other resetting earthquake occurred
291 in \mathcal{A} in the time interval considered, therefore t_0 corresponds to June 28, 1992.

292 iii) $M_{c,0}^*$ is estimated through a grid search; specifically, we search the $m_{c,0}^*$ in the
293 set $[4, 4.1, \dots, 6]$ which maximizes the likelihood ratio in favour of TGRE in the
294 first week of data. As shown in Fig. 9, we find $m_{c,0}^* = 4.3$. Of course, more
295 sophisticated procedures to estimate $M_{c,0}^*$ are possible, but we argue that the
296 results are stable for reasonable variations of this parameter. In particular, the
297 log-likelihood ratio remains well above zero (TGRE explains the data better than
298 TGR) for $4.1 \leq m_{c,0}^* = 4.3 \leq 4.8$. Then, in Table 1 we show also that the $M_{c,0}^*$
299 estimated in the first week of data brings to a superiority of TGRE with respect
300 to TGR also for other time windows (1 month, 3 months, and 1 year) (see section
301 "Results" for more details).

302 iv) $\nu = \frac{1}{\tau(1-2cov)}$, where the recurrence time $\tau = 250$ years is rescaled accounting for
303 the covariance coefficient $CoV = 0.3$.

304 The results are illustrated in the next section. To check their stability and the
305 sensitiveness of the model, besides using different $M_{c,0}^*$, we perform the analysis also
306 for other possible values of the parameters M_c^* and τ . The details are reported in the
307 caption of Table 1. We anticipate that the results are not significantly modified, in

308 agreement with what shown in Fig. 2 previously discussed.

309 Results

310 Results are illustrated in figures from 5 to 8, respectively for 1 week, 1 month, 3 months
311 and 1 year since the day after the resetting Landers event. The space-time windows in
312 which the analysis is performed are shown in the map of panels *a*), where the on-rupture
313 seismicity of \mathcal{A} (red dots) is reported inside the red stripe.

314 In panels *b*) of figures from 5 to 8, we show the results of the null hypothesis of
315 having the same MFD inside and outside the ruptured area. In particular, we plot
316 the earthquake cumulative numbers of events inside (in red) and outside (in dark
317 blue) \mathcal{A} in different time windows. In each of the four temporal intervals, red and dark
318 blue step functions are clearly different, and the two-samples Kolmogorov-Smirnov test
319 (Massey (1951)) confirms the rejection of the null hypothesis that the data are drawn
320 from the same continuous distribution, with a p - *value* much smaller than the 1%
321 significance level chosen before carrying out the analysis. We stress that these results
322 are completely independent of the modeling, as they are obtained by considering only
323 earthquake data. At the same time, these results support the main motivation of this
324 work, i.e., empirical data support the hypothesis of different MFDs on- and off-rupture
325 just after a large shock.

326 Panels *c*) in all the figures from 5 to 8 show the goodness-of-fit of the TGRE and
327 TGR models with respect to the earthquake data inside \mathcal{A} . Specifically, we plot the
328 TGR model in black versus the TGRE one in yellow, orange, green and blue respectively
329 for 1 week (Fig. 5), 1 month (Fig. 6), 3 months (Fig. 7) and 1 year (Fig. 8) since the

330 reset. We also show 1000 simulations of 1000 magnitudes each, obtained both with
 331 $m_c^* = 7.59$, that is drawn from a TGR (light gray cones), and with the new corner
 332 magnitudes $m_c(E)$ obtained for the TGRE model (light yellow, orange, green and blue
 333 cones for the four temporal intervals considered). Results show that within one week
 334 since the Landers earthquake, the TGRE corner seismic moment is reduced to a value
 335 $\sim M_{c,0}^*$ corresponding to the minimum energy, and after that it increases with the
 336 energy reloading process.

337 In all the four cases, the TGRE model gives visually a better fit to earthquake data
 338 than the TGR model: the red step functions representing the recorded magnitudes are
 339 almost completely contained in the non-gray cones, indicating our model's capability
 340 to better reproduce the time evolution of the real seismicity occurred in \mathcal{A} that just
 341 experienced the strong resetting Landers event. We argue that this general observation
 342 is independent from the choice on $M_{c,0}^*$, because of the clear bending in the MFD of
 343 the earthquakes inside \mathcal{A} .

344 We explore further the suitability of TGRE calculating the likelihood ratio for the
 345 nested TGR and TGRE models (King (1998)). The likelihood ratio is a measure of
 346 how much the TGRE is supported by the data with respect to TGR. In particular, the
 347 log-likelihood function

$$\log L(\theta) = N\beta_k \log M_{min} + \frac{NM_{min} - \sum_{i=1}^N M_i}{\theta} + \sum_{i=1}^N \log \left(\frac{\beta_k}{M_i} + \frac{1}{\theta} \right) - \beta_k \sum_{i=1}^N \log M_{min} \quad (6)$$

348 is the same for both the models, and it represents the TGRE when $\theta = M_c(E)$, and
 349 the TGR when $\theta = M_c^*$. In Table 1 we show the difference $\log L(M_c(E)) - \log L(M_c^*)$
 350 between the two log-likelihoods, computed for the four space-time windows considered

351 above. The results in the first row of Table 1 are relative to the earthquake data
 352 used in figures from 5 to 8, panels *c*), i.e., with $m_c^* = 7.59$, $m_{c,0}^* = 4.3$ and $\tau = 250$
 353 years. As anticipated in the previous section, to verify the stability of the results as a
 354 function of these parameters, we calculate the likelihood ratio also for different m_c^* and
 355 τ (see the first three columns in Table 1); for all these cases, we found that $m_{c,0}^* = 4.3$
 356 maximizes the likelihood ratio in favour of TGRE in the first week of data. The
 357 results of this stability test are shown in the rows from the second on. Borrowing the
 358 terminology adopted by Kass and Raftery (1995) for the Bayes factor, we may say that
 359 the evidence in favour of TGRE with respect to TGR is, most of the times, “substantial”
 360 and “strong”. As expected, this evidence lowers only in some cases for a long temporal
 361 window considered, but it still remains always > 0 , showing a superiority of TGRE
 362 with respect to TGR independently from the parameters. In general, the overall first-
 363 increasing-then-decreasing trend of the log-likelihood differences when moving to longer
 364 time periods is what expected, as a trade-off between the number of events and the
 365 recharging of the elastic energy of the system.

366 Finally, we find that the results remain stable also when considering completeness
 367 thresholds m_{min} higher than 2.5, or when removing the first few days just after the
 368 resetting Mw7.3 event, in which m_{min} may be higher than in the following days. As a
 369 matter of fact, any problem in the completeness magnitude should have equally affected
 370 the MFD of both events inside and outside \mathcal{A} , leaving the difference between the two
 371 distributions unchanged.

372 Discussion and Conclusions

373 Basic physical principles and empirical evidence suggest that MFD can vary with space
374 and time. To this purpose, in this paper we have proposed the energy-varying seismic
375 moment-frequency model TGRE for earthquake nucleation, which depends on the elas-
376 tic energy currently available in an area \mathcal{A} of interest. This model acknowledges the
377 elastic rebound theory and justifies the observation that the largest triggered earth-
378 quakes nucleate always outside the fault section which has just generated a large shock.
379 In a different perspective, the model may also describe quantitatively an intermittent
380 criticality state that is tuned by the available elastic energy. In other words, the state
381 of self-organized criticality (SOC) – advocated to explain the power law distribution
382 of the seismic moments at large spatiotemporal scales (Bak and Tang (1989); Sornette
383 and Sornette (1989)) – changes in intermittent criticality when zooming on small space-
384 time windows which have been recently involved by a large earthquake, indicating that
385 a fault system approaches and retreats from a critical state by turns (Ben-Zion et al.
386 (2003); Bowman and Sammis (2004); Bebbington et al. (2010)).

387 The TGRE distribution is obtained as a modification of Kagan’s TGR law, in which
388 the corner seismic moment is a time-varying energy function, i.e., it is linked to the
389 proxy of the amount of energy available in \mathcal{A} . The TGRE model is conceptually simple
390 and it depends on a few parameters: i) the corner seismic moment M_c^* , which is loosely
391 related to the strongest event that may nucleate in \mathcal{A} ; ii) the temporal occurrence of
392 the last large earthquake resetting the elastic energy in \mathcal{A} to a residual value; iii) the
393 rate of the energy recovering, which depends on the recurrence time of the fault(s)
394 involved; iv) $M_{c,0}^*$, which is the minimum value for the corner seismic moment which is

395 achieved after the occurrence of a resetting earthquake in \mathcal{A} . In other words, the TGRE
396 right tail $M_c(E)$ abruptly moves to $M_{c,0}^*$ just after the occurrence of a strong (resetting)
397 event, and then it slowly recovers to the long-term value; in practice, the model inhibits
398 the nucleation of a large triggered earthquake in segments that recently experienced a
399 large shock. An interesting feature of TGRE is that it verifies an invariance condition:
400 as the dimension of the selected space-time window becomes larger, it converges to the
401 TGR law with a limiting corner seismic moment M_c^* .

402 The TGRE has been designed purposely simple (depending on a few clear physical
403 parameters), acknowledging that understandability (and usability) is inversely propor-
404 tional to the complexity of a model. Like for any other model, it contains (more or less
405 explicit) subjective choices, but we think that these choices are less subjective than
406 ignoring the empirical evidence that strong triggered earthquakes do not nucleate in
407 the vicinity of a fault just ruptured by another strong event, like assumed in the (T)GR
408 model. Note that this empirical evidence can be hardly explained by space-time vari-
409 ability of the b -value of the GR law, which would lower, not inhibit, the triggering of
410 large earthquakes on a fault that has just slipped.

411 Despite its simplicity, we have shown that TGRE may explain well the statistically
412 significant difference of the MFDs relative to on- and off-rupture seismicity for the
413 Landers sequence, and that the results are stable for possible variations of the param-
414 eters. In particular, TGRE outperforms TGR for different values of $m_{c,0}^*$, showing the
415 strongest difference for $m_{c,0}^* = 4.3$. Further studies will be necessary to reduce uncer-
416 tainties on this value. For now, we just notice that the results seem to indicate that it
417 is more important the fact that we allow the corner seismic moment to vary in space

418 and time, rather than the details about the model’s parameters choice. That said, we
419 underline that the TGRE reliability (like for any other model) and the comparison with
420 alternative models (e.g., models based on space-time variations of the b -value) have to
421 be evaluated only through prospective tests. For this model, prospective tests will be
422 carried out in the framework of the ongoing european RISE project, which supports the
423 Collaboratory for the Study of Earthquake Predictability (CSEP) network activities in
424 Europe (for the websites, see “Data and Resources”; (Zechar et al., 2010; Schorlemmer
425 et al., 2018)) .

426 Finally, we suggest that the implementation of the TGRE may offer some benefits
427 for operational earthquake forecasting (OEF) models, because it overcomes one of the
428 conundrum of the best performing current clustering models (Taroni et al. (2018)), for
429 which the likelihood of a large earthquake is exactly where another large earthquake
430 has just occurred. This conundrum has been also identified as one of the main reasons
431 for the instability of the forecasts produced by the UCERF3-ETAS model, which has to
432 impose a space-time variability of the MFD to solve the problem (Field et al. (2017b)).
433 At the same time, TGRE may also provide a different explanation of recent empirical
434 evidence relative to variations of the b -value before and after large earthquakes close to
435 faults (Gulia and Wiemer (2019)). In particular, although it is worth remarking that
436 the meaning of the b -value is questionable for a distribution that is not exponential,
437 such as the TGR, if the corner magnitude gets closer to the completeness threshold
438 (even though the slope remains the same), the b -value necessarily increases (Marzocchi
439 et al. (2020)).

440 More in general, since the use of a proper MFD may have a large impact on the

441 earthquake predictability, we hope that this paper will stimulate further thoughts on
442 this issue.

443 **Data and Resources**

444 The data used in this study are available at <http://www.wgcep.org/ucerf3>, [https://](https://pubs.usgs.gov/of/2013/1165/)
445 pubs.usgs.gov/of/2013/1165/ and <https://pubs.usgs.gov/of/2007/1437/> (last
446 access, January 2019). Finally, for the RISE and CSEP projects, see respectively [www.](http://www.rise-eu.org)
447 [rise-eu.org](http://www.rise-eu.org) and https://scec.usc.edu/scecpedia/CSEP_Working_Group (last ac-
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Table 1: Difference between TGRE and TGR log-likelihoods in bold, and $m_c(E)$ values in brackets.

	1 week	1 month	3 months	1 year
$m_c^* = 7.59^a$ $\tau = 250$ y. ^e $m_{c,0}^* = 4.3$	3.16 (4.301)	3.51 (4.32)	2.76 (4.43)	1.04 (4.96)
$m_c^* = 7.53^b$ $\tau = 100$ y. ^f $m_{c,0}^* = 4.3$	3.15 (4.305)	3.36 (4.38)	1.85 (4.69)	0.26 (5.4)
$m_c^* = 7.5^c$ $\tau = 1000$ y. ^g $m_{c,0}^* = 4.3$	3.16 (4.3)	3.53 (4.301)	2.65 (4.31)	2.62 (4.4)
$m_c^* = 8.0^d$ $\tau = 500$ y. ^h $m_{c,0}^* = 4.3$	3.16 (4.301)	3.51 (4.32)	2.75 (4.43)	1.02 (4.96)

^a $m_c^* = 7.59$ as in Kagan et al. (2010).^c $m_c^* = 7.5$ like for S. Jacinto fault (Salisbury et al. (2012)).^e $\tau = 250$ like for Northern S. Andreas fault.^g $\tau = 1000$ as in Sieh et al. (1993).^b $m_c^* = 7.53$ by Wells and Coppersmith (1994)'s relations.^d $m_c^* = 8$ close to that for Northern S. Andreas fault.^f $\tau = 100$ like for S. Jacinto fault.^h $\tau = 500$ close to a mean value.

List of figure captions

655 **Fig. 1:** Time-varying corner seismic moment $M_c(E) = M_{c,0}^* + (M_c^* - M_{c,0}^*) [\nu(t - t_0)]^2$
656 as a function of the elapsed time $t - t_0$ between the event (t, M) and the resetting one
657 (t_0, M_0) . Panels *a*), *b*) and *c*) are obtained respectively for: fixed $(M_c^*, M_{c,0}^*)$ – varying
658 ν , fixed $(\nu, M_{c,0}^*)$ – varying M_c^* , fixed (ν, M_c^*) – varying $M_{c,0}^*$. The latter is obtained
659 for a shorter $t - t_0$ interval, because here the differences of the corner seismic moment
660 function can be appreciated: $M_c(E)$ substantially would not change over a longer tem-
661 poral interval, being M_c^* predominant over $M_{c,0}^*$. Magnitude values are shown in place
662 of seismic moments for an easier interpretation of the figure.

663

664 **Fig. 2:** Surface plots of the time-varying corner seismic moment $M_c(E) = M_{c,0}^* +$
665 $(M_c^* - M_{c,0}^*) [\nu(t - t_0)]^2$ as a function of the time elapsed since the reset $t - t_0$ and: the
666 parameter ν with fixed M_c^* in the first line panels, viceversa in the second line. The
667 minimum corner seismic moment is set at $m_{c,0}^* = 4.5$ in each panel.

668

669 **Fig. 3:** Survival function of the TGRE model for several values of the available energy
670 (corner seismic moment), corresponding to the $M_c(E)$ indicated in the legend, in a
671 log-log scale.

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673 **Fig. 4:** TGRE branching ratio (5) versus its time-varying corner seismic moment,
674 for several values of the difference $\beta_k - \alpha_k$.

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676 **Fig. 5:** TGR vs TGRE analysis relative to the considered area \mathcal{A} , covering the Landers

677 segment fault, as shown in panel *a*) (the colored lines with circles represent the nearby
 678 segment faults). The temporal interval here is 29 June 1992 - 6 July 1992, that is,
 679 within 1 week since the day after the Landers resetting earthquake. The number of
 680 events contained in this spatiotemporal window is 437 (red dots in panel *a*)). Panel *b*)
 681 contains the earthquake cumulative numbers of events inside \mathcal{A} (in red), and outside
 682 it (in dark blue). Finally, in panel *c*) we compare the fit to the data of the TGR model
 683 in black, and the TGRE model in yellow, obtained respectively with $m_c^* = 7.59$ and
 684 $m_c(E) = 4.301$ (the latter derived from equation (2) with $\alpha = 2$). These corner mag-
 685 nitudes are used also to obtain 1000 simulations of 1000 TGR- and TGRE- distributed
 686 seismic moments, respectively, which are plotted as light gray and light yellow cones.
 687 The data (red step functions) almost completely fall into the TGRE cone. Magnitude
 688 values are shown in place of seismic moments for an easier interpretation.

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690 **Fig. 6:** The same as Fig. 5, but relative to the temporal interval 29 June 1992 -
 691 29 July 1992, that is, within 1 month since the day after the Landers resetting earth-
 692 quake. The number of events contained in this spatiotemporal window is 739. The
 693 color used for the TGRE model is orange, and $m_c(E) = 4.32$.

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695 **Fig. 7:** The same as Fig. 5, but relative to the temporal interval 29 June 1992 -
 696 29 September 1992, that is, within 3 months since the day after the Landers resetting
 697 earthquake. The number of events contained in this spatiotemporal window is 926.
 698 The color used for the TGRE model is green, and $m_c(E) = 4.43$.

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700 **Fig. 8:** The same as Fig. 5, but relative to the temporal interval 29 June 1992 -
701 29 June 1993, that is, within 1 year since the day after the Landers earthquake. The
702 number of events contained in this spatiotemporal window is 1120. The color used for
703 the TGRE model is blue, and $m_c(E) = 4.96$.

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705 **Fig. 9:** Difference between TGRE and TGR log-likelihoods versus the minimum mag-
706 nitude $m_{c,0}^*$ achieved after the reset.

Figures

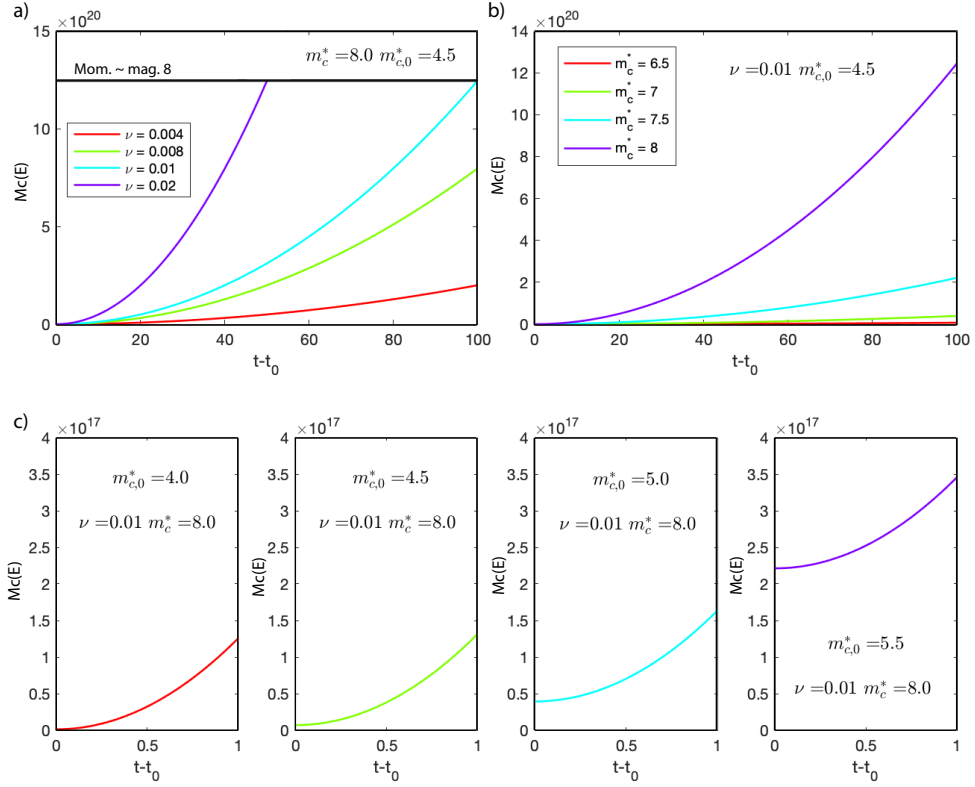


Figure 1: Time-varying corner seismic moment $M_c(E) = M_{c,0}^* + (M_c^* - M_{c,0}^*) [\nu(t - t_0)]^2$ as a function of the elapsed time $t-t_0$ between the event (t, M) and the resetting one (t_0, M_0). Panels a), b) and c) are obtained respectively for: fixed $(M_c^*, M_{c,0}^*)$ – varying ν , fixed $(\nu, M_{c,0}^*)$ – varying M_c^* , fixed (ν, M_c^*) – varying $M_{c,0}^*$. The latter is obtained for a shorter $t-t_0$ interval, because here the differences of the corner seismic moment function can be appreciated: $M_c(E)$ substantially would not change over a longer temporal interval, being M_c^* predominant over $M_{c,0}^*$. Magnitude values are shown in place of seismic moments for an easier interpretation of the figure.

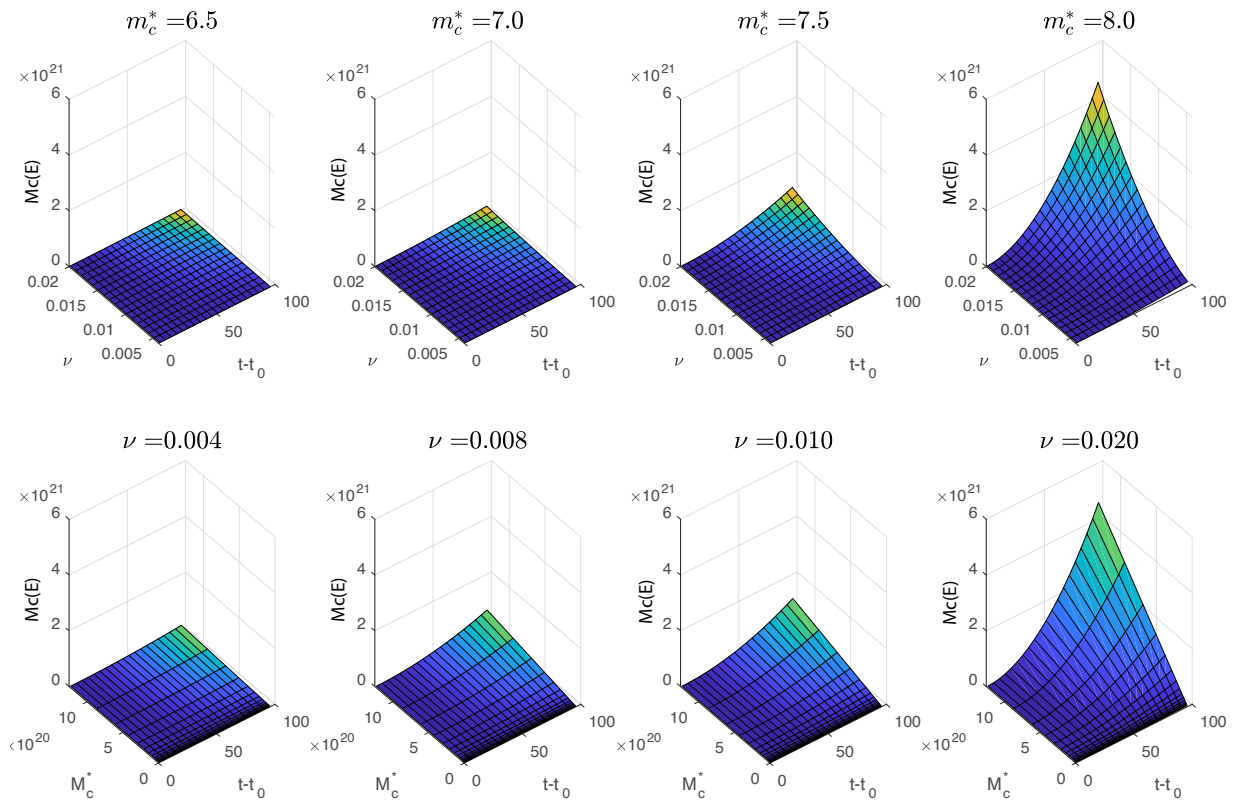


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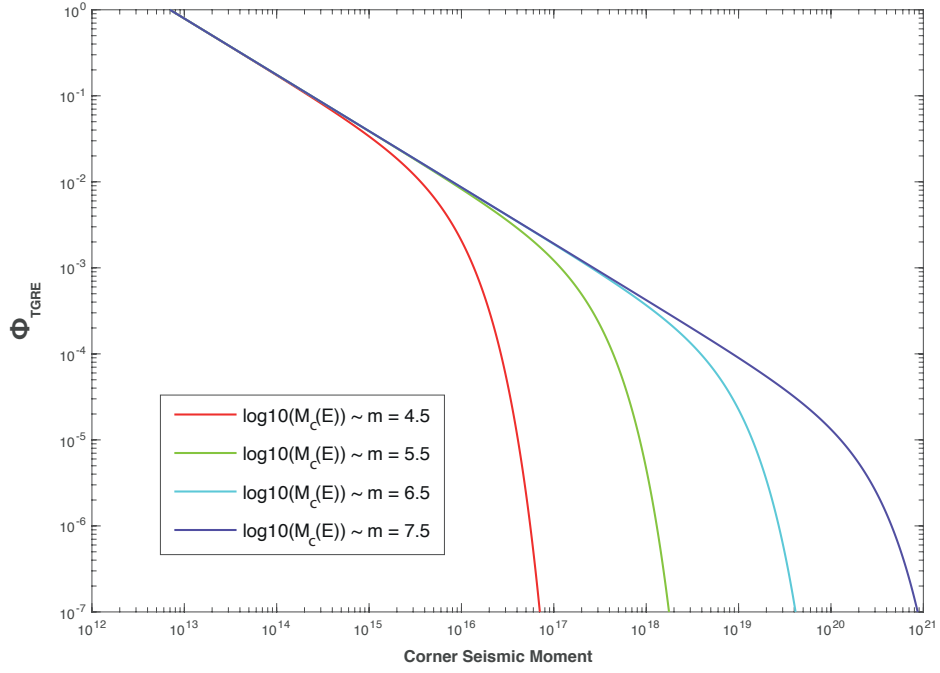


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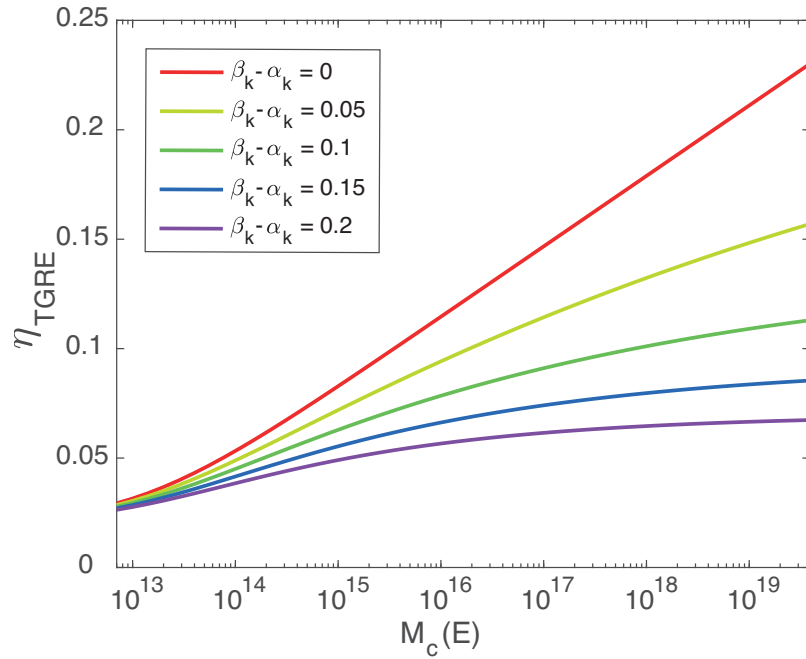


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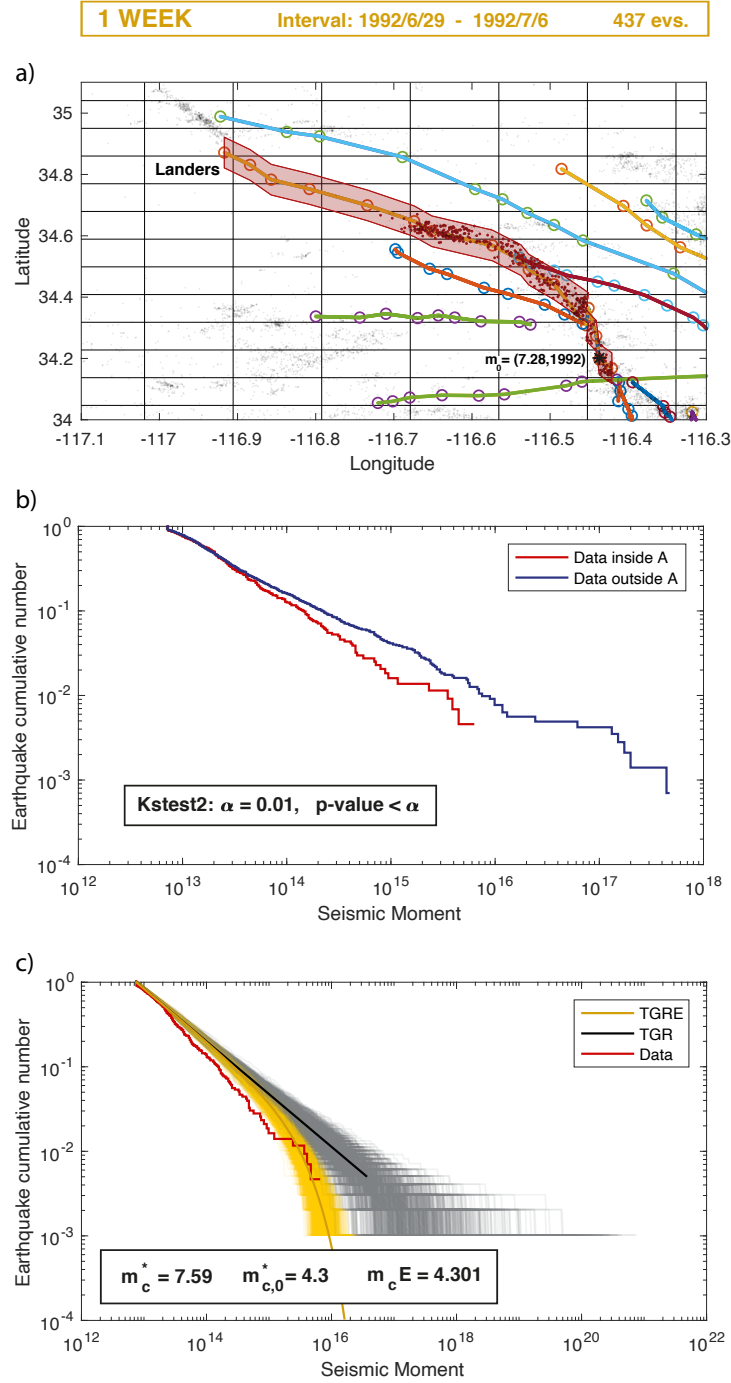


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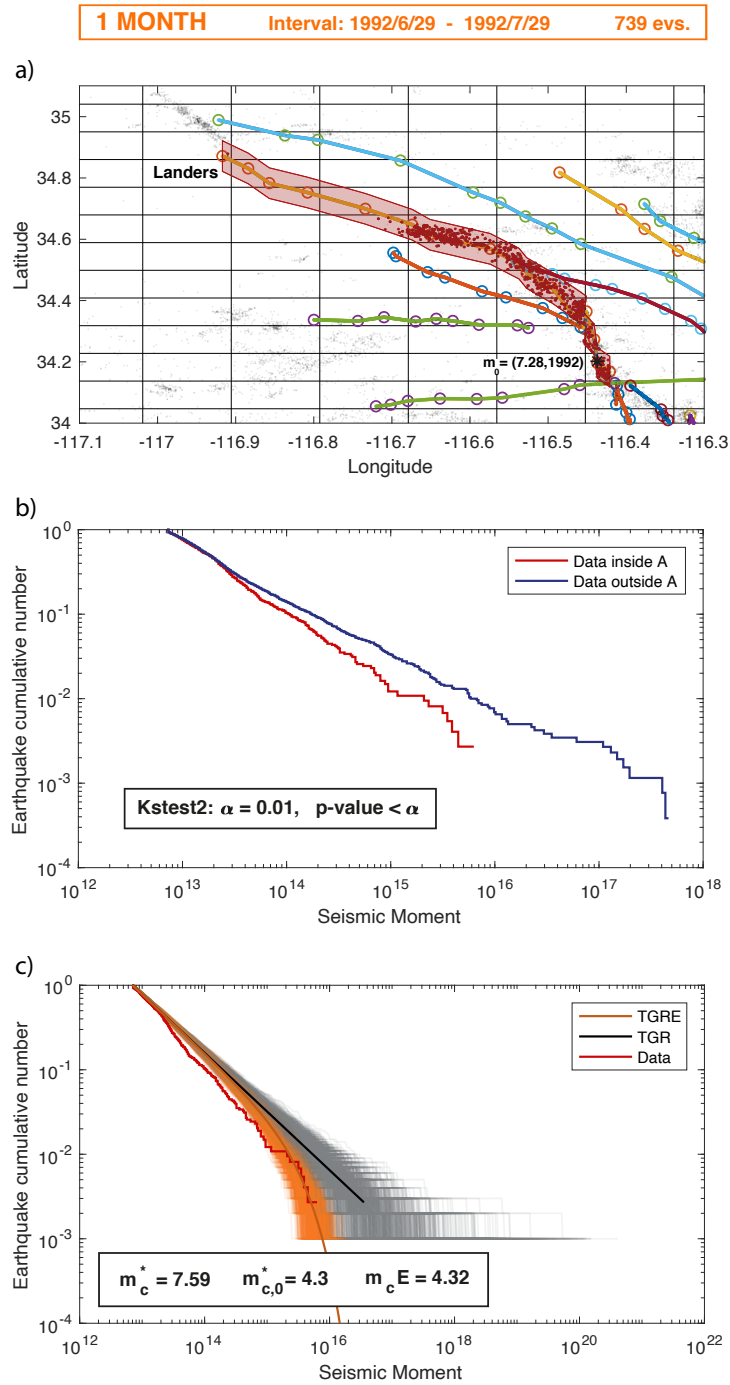


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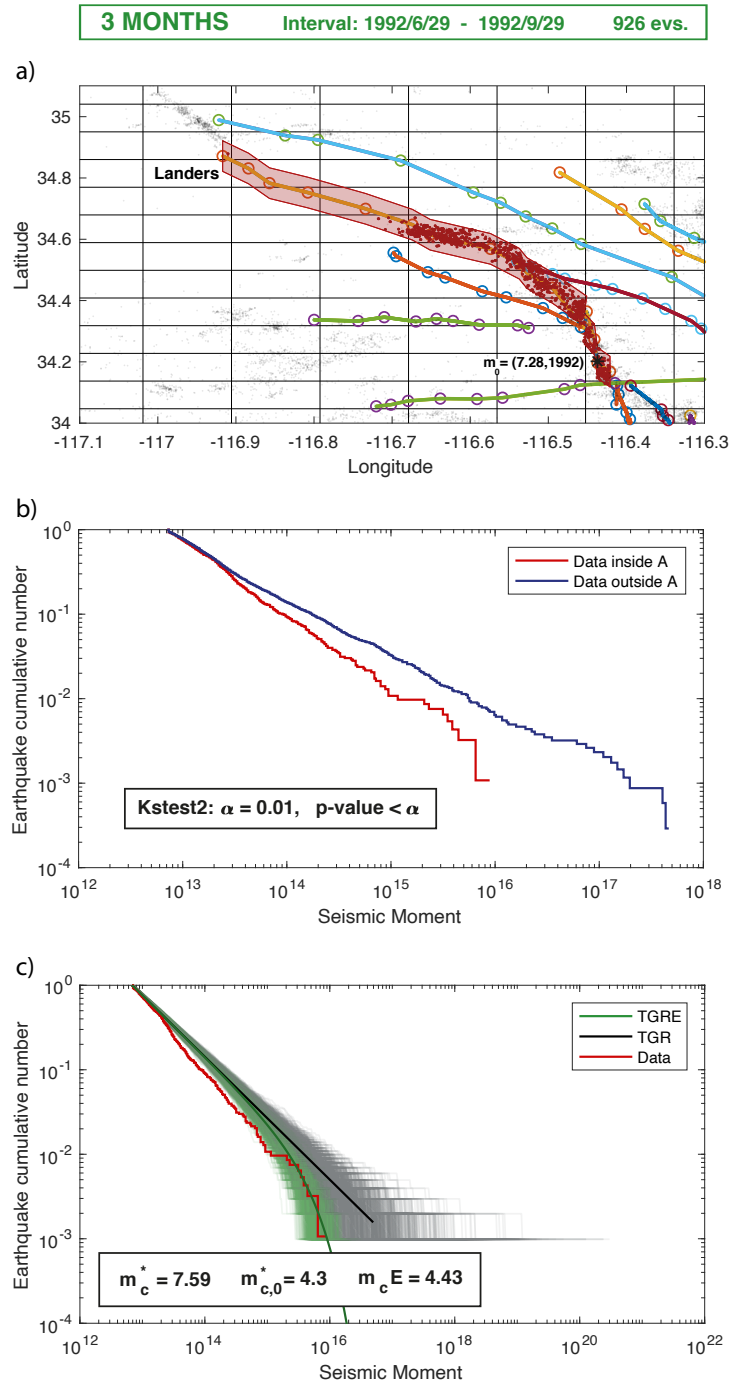


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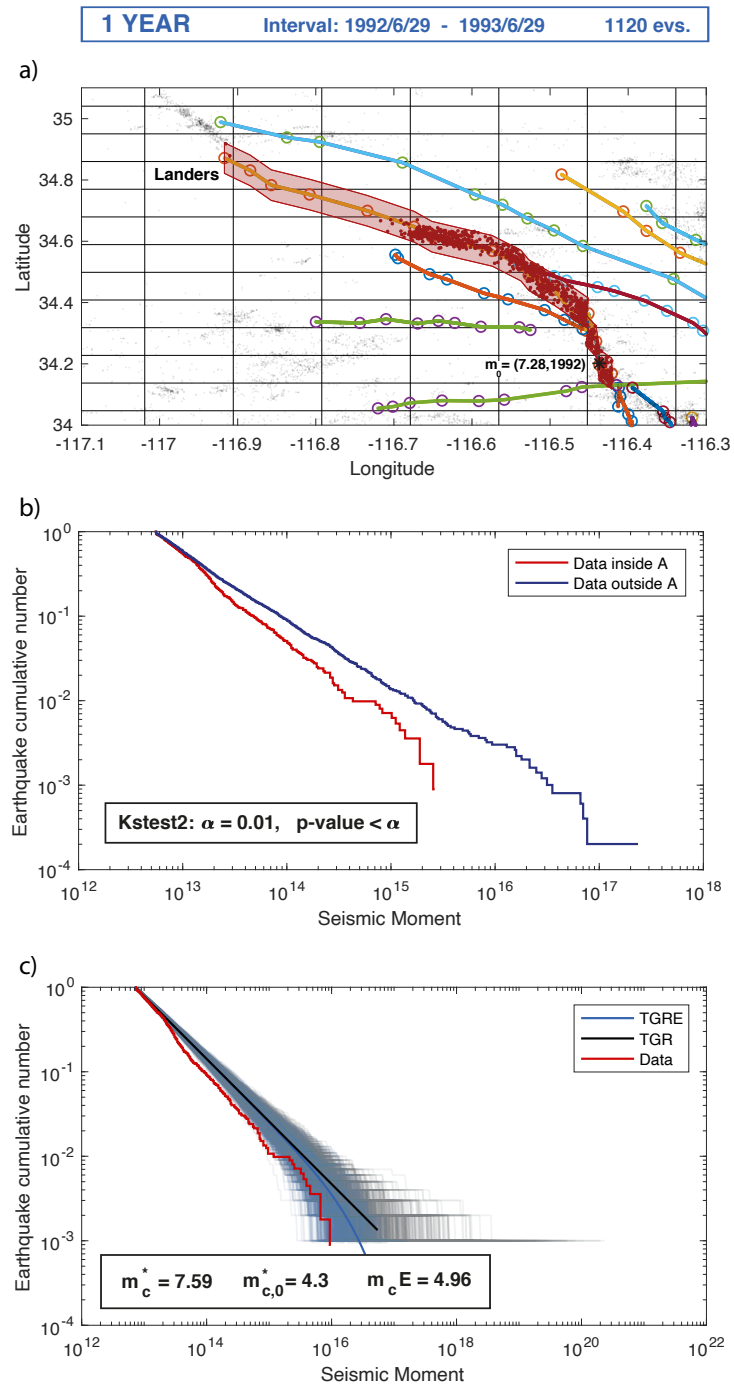


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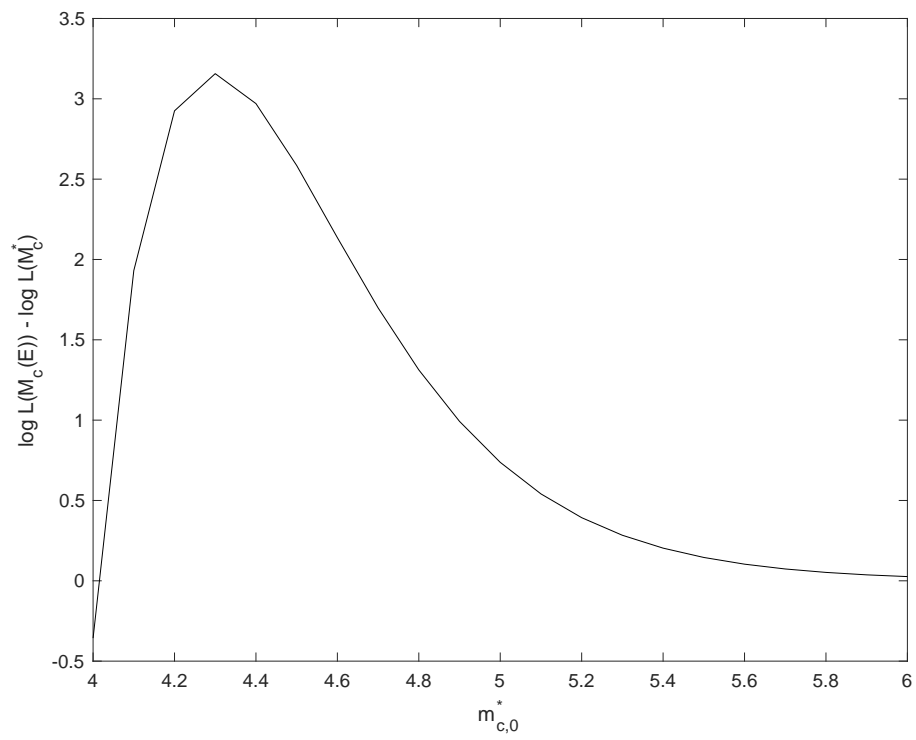


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