3 4	1	The unexpected explosive sub-Plinian eruption of Calbuco volcano (22–23 April 2015;
5 6	2	southern Chile): triggering mechanism implications
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#### Abstract

Plinian-type eruptions are extremely hazardous, producing pyroclastic fallout and flows extending many kilometres from the vent. The most commonly invoked eruption trigger for Plinian-type eruptions is the intrusion of fresh magma, generally associated with precursory ground deformation and seismicity days/weeks before eruption. Closed-system internal triggering has also been proposed, such as protracted crystallisation of magma, which can produce a build-up of exsolved volatiles and thus pressurise the system prior to eruption. On 22-23 April 2015 Calbuco volcano, Chile, produced a sub-Plinian eruption with <3 hours seismic precursory activity and no clear deformation signals in the preceding months. Here, we show that petrological and geochemical evidence do not support a hypothesis of eruption triggering due to pre-eruptive intrusion of fresh magma, but instead are consistent with an internal trigger. We found that basaltic andesitic magma was stored at depths between 8 and 12 km (i.e. 230-320 MPa) beneath Calbuco volcano before the 2015 eruption. The stored magma had an initial temperature of 900-950 °C, was water-saturated (5.5-6.5 wt.% H<sub>2</sub>O) and formed phenocrysts of titanomagnetite, orthopyroxene, clinopyroxene and plagioclase cores  $(An_{78-93})$ . Gradual cooling of the magma chamber produced thermal gradients and magma convection, evidenced by plagioclase overgrowth rims (An<sub>58-77</sub>) and blocky microlites (25-250  $\mu$ m). Our interpretation is that this continuing crystallisation induced second boiling and an over-pressurisation of the system, leading to the rapid onset of the 2015 eruption. Petrological and geochemical evidence therefore shows that a closed-system magma chamber can evolve into a highly explosive eruption with very little precursory warning, posing a challenge for current volcano monitoring paradigms. We propose that internal triggering should be carefully considered as a mechanism for unexpected sub-Plinian eruptions, prompting a potential revision of existing hazard management strategies.

#### **1. Introduction**

Plinian-type eruptions are the largest explosive volcanic events in terms of their mass discharge rate (intensity) and erupted magma volume (magnitude) (Walker, 1980). The temporal recurrence of sub-Plinian eruptions  $(0.1 - 1.0 \text{ km}^3 \text{ ejecta volume}, >10 \text{ km plume height})$  is about every year on Earth, while Plinian eruptions  $(1.0 - 10.0 \text{ km}^3 \text{ ejecta volume}, >20 \text{ km})$ plume height) occur about every decade, and therefore together they produce significant volcanic hazards (Newhall and Self, 1982). The explosivity of Plinian-type eruptions generally requires volatile-rich and high viscosity magmas. Usually, these types of eruptions are associated with evolved calc-alkaline (andesite to rhyolite; Castro and Dingwell, 2009; Castro et al., 2013) and alkaline (trachyte to phonolite; Signorelli et al., 1999) magmas, although there are some rare exceptions of basaltic Plinian eruptions (e.g. Houghton et al., 2004; Walker et al., 1984).

The triggering mechanisms for large explosive eruptions are difficult to constrain and are the subject of considerable debate (Caricchi et al., 2014; Malfait et al., 2014; Gregg et al., 2015; Gudmundsson, 2016). These can be broadly grouped into five overlapping categories: (i) Injection of volatile-rich magmas into a more evolved magma chamber, which is the most commonly invoked triggering mechanism (Blake, 1981; Williams and Self, 1983; Walker et al., 1984; Pallister et al., 1992; de Silva et al., 2008; Sigmundsson et al., 2010; Wehrmann et al., 2016; Cassidy et al., 2016); (ii) Magma mixing following mafic magma injection (Sparks et al., 1977; Leonard et al., 2002; Stock et al., 2012; Macías et al., 2017); (iii) generation of overpressure due exsolution of volatiles via second boiling due to protracted crystallisation (Stock et al., 2016; Tramontano et al., 2017); (iv) perturbation of magmatic reservoirs close to a critical state by tectonic earthquakes (Linde and Sacks, 1998); (v) sector collapse of a volcanic edifice (e.g. Mount St Helens in 1980; Lipman and Mullineaux, 1981).

Most volcanic eruptions are preceded by a period of volcanic unrest, marked by particular patterns of seismicity, gas emission, and ground deformation providing opportunities for eruption forecasting (Gorshkov and Dubik, 1970; Chouet et al., 1994; Chouet, 1996; Harlow et al., 1991; Sigmundsson et al., 2010; Aiuppa et al., 2010; Tarasewicz et al., 2014; White and McCausland, 2016; Riveira et al., 2017). The alert period for most eruptions is measured in weeks to months, where an increase in seismic activity may signal a greater likelihood of an eruption. The alert period can be as short as a few hours before eruption onset, leaving insufficient time to draw up actions for minimizing volcanic risk (Castro et al., 2013).

Improved knowledge of the pre-eruptive conditions and triggering mechanisms of Plinian-type events is of paramount importance for volcanologists to provide better short-term forecasts of eruption onset and hence mitigate hazardous situations (Sparks, 2003; Segall, 2013; Tramontano et al., 2017). The aim of this study is to better understand the conditions and mechanisms leading to the onset of Plinian-type eruptions, especially those that are preceded by only short periods of volcanic unrest. We focus on a case study, using petrological, geochemical and seismic data to constrain the pre- and syn-eruptive conditions of the 22-23 April 2015 sub-Plinian eruption of Calbuco volcano, southern Chile (Supplementary Fig. 1).

After 54 years of quiescence since its last sub-Plinian eruption in 1961, Calbuco volcano erupted on 22 April 2015 at 21:05 GMT (18:05 local time) near the cities of Puerto Montt, Puerto Varas and Ensenada, inhabited by a total of ~260,000 people. Calbuco volcano has been monitored by two seismometers since 2009, and a tiltmeter is located at 4 km west of the summit. The eruption started suddenly with only ~3 hours of precursory seismic activity. Pre-eruptive deformation was not detected by the tiltmeter nor observed in InSAR interferograms up to one day before the eruption (SERNAGEOMIN, 2015; Valderama et al., 2015; Delgado et al., 2017). Delgado et al. (2017) discussed that pre-eruptive deformation may have not been detected if it was lower than the background noise. Only co-eruptive deformation was detected 

during the eruption (Delgado et al., 2017). The first eruptive phase generated an eruptive column which reached a height of ~15 km above the crater level and lasted ~1.5 hours (SERNAGEOMIN, 2015; Romero et al., 2016; Van Eaton et al., 2016; Castruccio et al., 2016; Pardini et al., 2018). After a pause of 5.5 hours, a second and more energetic eruptive phase started on 23 April at 04:08 GMT (01:08 local time) and lasted for 6 hours, developing an eruptive column >15 km above the crater level (SERNAGEOMIN, 2015; Romero et al., 2016; Van Eaton et al., 2016; Castruccio et al., 2016). 

The total erupted volume has been estimated in four articles as 0.27 km<sup>3</sup> (Romero et al., 2016), 0.38 km<sup>3</sup> (Castruccio et al., 2016), 0.48 km<sup>3</sup> (Pardini et al., 2018, assuming 1000 kg/m<sup>3</sup> density) and 0.56 km<sup>3</sup> (Van Eaton et al., 2016). Of this total volume, 38% was erupted during the first phase of the eruption and 62% during the second phase. The eruption has been classified as a sub-Plinian type event, according to the pyroclastic fall deposit (Romero et al., 2016; Castruccio et al., 2016). The magma bulk composition was basaltic andesite (Romero et al., 2016). 

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 eruption with minimal eruption precursors, posing a significant challenge for risk managers.
 The purpose of this paper is to examine the processes that may produce such eruptions.

<sup>275</sup> 276 115

<sup>277</sup><sub>278</sub> 116 **2. Methods** 

280 117 2.1 Fieldwork

Fall deposits at proximal and medial distances from the vent were sampled during fieldwork on 2-10 June 2015. Four main tephra units have been recognised within the eruptive sequence (units A, B, C, and D; see Fig. 1). Here we present analyses of basaltic andesite scoriae collected in two of the proximal fall deposits: DF-2, ~6 km from the vent, and DF-10, ~4 km from the vent (Romero et al., 2016). 

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#### 1 124 2.2 Scanning Electron Microscopy (SEM) and textural analysis

Back-scattered electron (BSE) images were collected using a JEOL JSM-6390LA FE-SEM
at the School of Earth and Environmental Sciences, University of Manchester (UK), using an
acceleration voltage of 15 kV and beam current of 10 nA.

Textural analysis on BSE images was performed using ImageJ software. 2D image analysis was performed to measure the area of each phase (crystal phases, matrix glass and bubbles) in scoriae of the four units of the tephra fall deposits. The crystal fraction  $\phi$  of each phase was calculated on a vesicle-free basis:

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$$\phi = \frac{A_{xtal}}{A_r}$$

<sup>321</sup> 133 where  $A_{xtal}$  is the area of the crystal phase and  $A_r$  is the reference area used for the image <sup>323</sup> 134 analysis of natural samples (105 mm<sup>2</sup> per each sample). The sizes of crystal phases were also <sup>325</sup> 135 measured using ImageJ software. All crystal fraction data are provided in Table 1.

Crystal size distribution (CSD) analysis provides quantitative information on relationships between crystal population density and crystal length for a population of crystals. Calbuco scoriae were used to analyse plagioclase CSD in order to quantify the nucleation events occurred before the eruption and to estimate the magma residence time based on the relationship between plagioclase population density, crystal sizes, and experimentally determined plagioclase growth rates. Crystal dimensions and abundances of each size population were recovered through analysis of BSE images using ImageJ software. The relationship between crystal population density and crystal length for a population of crystals was then obtained using CSD Corrections 1.6 (Higgins 2000, 2002). The linear relation provides estimates of timescales of magmatic processes, as the slope of the correlation is equal to -1/(growth rate  $\times$  residence time) (Higgins 2000). 

2.3 X-Ray Fluorescence (XRF) Bulk rock compositions of juvenile material were measured at the School of Earth and Environmental Sciences of the University of Manchester by X-ray fluorescence (XRF) spectrometry. The PAN analytical spectrometer, characterised by wavelength dispersive (WDXRF) systems, was used to analyse the bulk major and trace element compositions of the scoria clasts. This instrument uses a Rh anode X-ray tube and is configured for geochemical analysis of rocks (Potts and Webb, 1992). The precision is better than 5% for all major elements and better than 10% for trace elements. 2.4 Electron microprobe analysis (EMPA) Major element concentrations of minerals, plagioclase-hosted melt inclusions (MIs) and matrix glasses were analysed using a Cameca SX100 at Centro Nazionale di Ricerca of Firenze (Italy) and a JEOL JXA-8530F field emission electron microprobe at the Photon Science Institute, University of Manchester. For both instruments, the operating conditions were as follows: 15 kV accelerating voltage, 10 nA beam current, and beam diameter of 10 or 5 µm (the latter for microlites). Sodium and potassium were measured first to minimise loss by volatilisation. Calibration standards were albite for Na, periclase for Mg, corundum for Al, fayalite for Fe, tephroite for Mn, apatite for P, sanidine for K, wollastonite for Ca and Si and rutile for Ti. The uncertainty is between 1 to 5% for major elements and up to 10% for elements at concentration <0.2 wt%. 2.5 Laser ablation ICP-MS Trace element concentrations in minerals (plagioclase, orthopyroxene, clinopyroxene), plagioclase-hosted MIs and matrix glasses were determined by Laser Ablation Inductively 

Coupled Plasma Mass Spectrometry (LA-ICP-MS) at the Department of Physics and Geology, University of Perugia (Italy). The instrumentation consisted of a Teledyne/Photon Machine G2 LA device equipped with a Two-Volume ANU HelEx 2 cell coupled with a Thermo Scientific<sup>TM</sup> guadrupole-based iCAP O ICP-MS. Analyses were performed using a circular laser beam with a diameter of 20 µm, frequency of 8 Hz, and a laser density on the sample surface of 3.5 J/cm<sup>2</sup>. The instrument was calibrated using the reference material NIST SRM-610, and <sup>29</sup>Si as an internal standard. The USGS reference material BCR2-G was analysed as an unknown in order to monitor precision and accuracy (Rocholl, 1998), which are better than 10% for all elements under these operating conditions (Petrelli et al., 2016). 

The acquisition of measured transient signals allows the identification of melt inclusions in
trystals (Longerich et al., 1998). Raw data were carefully screened for such inclusion signals
and these were then manually removed.

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444 185 2.6 Raman spectroscopy

Raman spectra were collected from exposed melt inclusions hosted in plagioclase, orthopyroxene and clinopyroxene phenocrysts, using a Thermo Scientific<sup>TM</sup> DXR<sup>TM</sup>xi Raman Imaging Microscope at the School of Earth Sciences, University of Bristol (UK). The spectrometer is equipped with a 532 nm doubled Nd:YVO4 DPSS, a 900 lines  $\times$  mm<sup>-1</sup> grating, and a magnetic stage. Spectra were acquired between 100 cm<sup>-1</sup> and 4000 cm<sup>-1</sup> using a  $100 \times$ objective, 25 µm confocal pinhole, and 3 mW of laser power on the sample at a depth of ~6  $\mu$ m. The acquisition time was set to 20 × 3 sec in order to maximise the signal-to-noise ratio.

The acquired spectra were corrected for temperature and excitation line effects according to Long (1977). For the silicate region, a background subtraction was applied following the strategy of Di Genova et al. (2016), where a cubic spline was fit through two intervals devoid of peaks (100-250 and 1250-1500 cm<sup>-1</sup>, respectively). A cubic baseline between 2750-3100 and 

3750-3900 cm<sup>-1</sup> was applied to the water region. A set of standards of dacitic composition (HO series from Di Genova et al., 2017) was used to estimate the water content of the melt inclusions. This was calculated using the ratio between the band area of the water (HW) and silicate (LW) region (i.e. internal calibration). The root mean square error (RMSE) of the water content measured using the internal calibration is 0.15 wt.% (Di Genova et al., 2017). 

<sup>488</sup> 203 **3. Results** 

<sup>490</sup> 204 *3.1 Tephra fall deposits* 

The eruptive sequence is subdivided in four units: A, B, C, and D. The sequence starts with unit A (showing a reverse grading; Fig.1), which was deposited during the first phase of the eruption (Romero et al., 2016; Castruccio et al., 2016). Units B, C and D were deposited during the second phase of eruption (Castruccio et al., 2016). At site DF10 (Fig. 1) unit A consists of two types of lapilli-size scoriae; one is lowdensity, vesicular and light grey to light brown in colour (74 vol.%), and the other is high density, poorly vesicular and grey to brown in colour (23 vol.%). Lithics are also present (3 vol.%). The sequence continues with units B and C of the second eruptive phase, which are normally graded and ungraded, respectively. Units B and C are composed of brown scoriae with sizes from coarse lapilli to bomb-sized. At site DF10 (Fig. 1) both units B and C consist of low-density, vesicular and light brown scoriae (60 vol.%); high-density, poorly vesicular brown scoriae (37 vol.%); and lithics (3 vol.%). Finally, the uppermost unit D is ungraded and composed of high-density, poorly vesicular lapilli scoriae (71 vol.%); low-density, vesicular lapilli scoriae (24 vol.%); and lithics (4 vol.%). 

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522 219 *3.2 Textures* 523

The mineral assemblage in the scoriae consists of plagioclase (Pl), orthopyroxene (Opx),
 clinopyroxene (Cpx), titanomagnetite (Ti-Mag) and olivine (Ol) (Figs. 1 and 2). Scoria clasts

are characterised by isolated phenocrysts and glomeroporphyritic textures (Fig. 1). The groundmass is characterised by microlites of plagioclase, clinopyroxene, orthopyroxene and titanomagnetite, plus glass (Figs. 1 and 2). In general, phenocrysts have euhedral blocky and prismatic shapes (Fig. 1). Microlites are characterised by euhedral, hopper, swallowtail and skeletal morphologies with prismatic to acicular habits (Figs. 1 and 2).

The plagioclase crystal fraction  $\phi_{Pl}$  is constant throughout the four units of the tephra fall deposit:  $\phi_{Pl}$  ranges between 0.23 and 0.27 for phenocrysts, in particular, the crystal fraction of plagioclase cores is ~0.10, whereas  $\phi$  of plagioclase overgrowth is ~0.15. The microlite crystal fraction ranges from 0.17 to 0.21 (Table 1), in which the crystal fraction of blocky microlites is  $\leq 0.10$  with respect to the analysed sample. Crystal size distribution (CSD) analysis shows 3 populations of plagioclase crystals (Fig. 3 and Supplementary Table 1). Plagioclase phenocrysts are characterised by a normally zoned core and an overgrowth rim with oscillatory zoning (Fig. 4) and their sizes range between 300  $\mu$ m and 1.5 mm. Oscillatory zoning shows internal resorption and growth morphologies (Fig. 4), sieve textures (Fig. 4e) and patchy zoning textures (Fig. 4e, f). Plagioclase microlites vary from 1 to 250 µm in size. Two textures can be distinguished in plagioclase microlites: those with sizes between 25 and 250 µm have euhedral prismatic shapes with occasional oscillatory zoning (blocky microlites), while those with sizes <15 µm are elongated, acicular crystals with hopper and swallowtail shapes (skeletal microlites). The larger blocky plagioclase microlites have CSD slopes that range between -12.44 and -21.72, while the smaller skeletal microlites have CSD slopes between -221.69 and -319.85 (Supplementary Table 1). 

Pyroxene crystals (Opx and Cpx) are less abundant than plagioclase crystals and  $\phi_{Px}$  ranges from 0.07 to 0.14 ( $\phi_{Px-phenocryst}$ : 0.07-0.09,  $\phi_{Px-microlite}$ : <0.01-0.04; Table 1). Pyroxene phenocryst sizes range between 200 and 700 µm, whereas the dimensions of microlites are between 1 µm and <130 µm. 

Titanomagnetite phenocrysts are commonly found within glomerocrysts and their crystal fraction is constant at  $\phi_{Ti-Mgt}=0.01$  through the four units of the tephra fall deposit (Table 1). Titanomagnetite phenocrysts are  $\sim 120$  to  $\sim 300 \ \mu m$  in diameter, while microlites are  $< 100 \ \mu m$ . In the groundmass,  $\phi_{\text{Ti-Mgt}}$  is <<0.01. Olivine crystals are typically found in glomerocrysts or as inclusions in the cores of orthopyroxene phenocrysts. Olivine is an accessory phase and  $\phi_{Ol}$  is <0.01. Olivine crystals are characterised by anhedral habits, which indicate resorption and disequilibrium conditions. Melt inclusions are hosted in plagioclase, orthopyroxene and clinopyroxene phenocrysts. Plagioclase-hosted MIs are characterised by spherical, oblate and polygonal shapes (Fig. 4e, f) and have diameters between 30 to 100 µm. Ortho- and clinoyroxene-hosted inclusions are much smaller, with diameters generally  $<40 \,\mu$ m, and are often irregularly shaped. 3.3 Bulk rock, MIs and glass compositions The bulk rock composition of the juvenile material is basaltic andesitic containing 55-56 wt.% SiO<sub>2</sub> (Fig. 5; Supplementary Table 2). There is no statistically significant compositional difference between the four units of the tephra fall deposit. Melt inclusions (MIs) hosted in orthopyroxene and clinopyroxene phenocrysts were difficult to characterise chemically as their small sizes (smaller than those hosted in plagioclase) precluded microprobe analysis. MIs hosted in plagioclase phenocrysts, mostly entrapped in external layers (Fig. 4e, f), have compositions from andesitic to low-silica dacitic, with SiO<sub>2</sub> between 58 and 62 wt.% (Fig. 5; Supplementary Table 3). Only plagioclase-hosted MIs were 

chemically characterized, as their sizes allowed us to analyse by EPMA and LA-ICP-MS.
Plagioclase-hosted MIs were corrected for post-entrapment crystallisation (PEC) following the
method of Neave et al. (2017), incrementally adding equilibrium plagioclase back into the
inclusions until their MgO-Al<sub>2</sub>O<sub>3</sub> systematics were consistent with a liquid line of descent

connecting bulk rock and glass compositions. Trace elements were corrected using partition coefficients collated Geochemical Earth Reference from the Model database (https://earthref.org/GERM/); volatiles were assumed to be perfectly incompatible. The mean PEC correction applied was 13.8% (range 1-30%).

The matrix glass ranges from andesite to dacite (SiO<sub>2</sub> between 61 and 65 wt.%) (Fig. 5; Supplementary Table 3). The full range of matrix glass compositions is found within each unit of the tephra fall deposit. Harker-style diagrams show that Al<sub>2</sub>O<sub>3</sub>, FeO, MgO and CaO decrease as SiO<sub>2</sub> increases (Fig. 5a, b, c, d), highlighting a chemical evolution from bulk rock to residual melt, passing through the corrected MI compositions. The decrease in FeO and MgO (Fig. 5b, c) may be due to the crystallisation of orthopyroxene, clinopyroxene and oxides, while the variations in CaO content (Fig. 5d) are primarily controlled by plagioclase crystallisation. The observed variability in Na<sub>2</sub>O (Fig. 5e) could be a real magmatic feature partially controlled by plagioclase crystallisation, or could be due to slight Na<sub>2</sub>O loss during EPMA analysis. However, EPMA analyses were carefully performed in order to avoid Na<sub>2</sub>O loss. Our PEC correction has little effect on the Na<sub>2</sub>O content of the inclusions and does not decrease the variability in measured Na<sub>2</sub>O contents (Fig. 5e). Figure 5f shows that K<sub>2</sub>O is enriched in the residual melt as it is incompatible in plagioclase, orthopyroxene and clinopyroxene. Trace element data also show an enrichment of incompatible elements (e.g., Rb, Y, Nb, Ba and La) in the matrix glass, consistent with residual melt differentiation obtained with the fractional crystallisation model (Neumann et al., 1954; Petrelli et al., 2005) (Fig. 6). Sr contents are more variable than other trace elements (Fig. 6b) because Sr is slightly compatible in plagioclase and the partitioning of Sr between plagioclase and melt is a function of the anorthite (An) content (Blundy and Wood, 1991; Aigner-Torres et al., 2007). 



Plagioclase phenocrysts typically exhibit a normally zoned core and an oscillatory zoned overgrowth rim (Fig. 4). Plagioclase cores are characterised by An contents between 78 and 93 mol.% (Fig. 7a). The An content of the overgrowth rims varies from 58 to 77 mol.% (Fig. 7a). The overgrowth rims, particularly the external rims in contact with the melt ( $An_{61-75}$ ), have compositions similar to the microlites  $(An_{63-70})$  (Fig. 7a). Therefore, the transition between the core and the oscillatory zoned overgrowth rim is characterized by an important decrease of An content, however, this transition is not associated with a significant variation of FeO and MgO (Fig. 8). Therefore, concentrations of FeO and MgO in plagioclase are approximately constant from the core to the overgrowth rim (through the oscillatory zoning).

Orthopyroxene phenocrysts are slightly compositionally heterogeneous with En content between 63 and 73 mol.% (En<sub>65-73</sub>, Fs<sub>24-32</sub>, Wo<sub>2-4</sub>; Fig. 7b). Clinopyroxene phenocrysts are uniformly augitic at En<sub>41-45</sub>, Fs<sub>10-15</sub>, Wo<sub>41-45</sub> (Fig. 7b). Olivine crystals have compositions ranging from Fo<sub>72</sub> to Fo<sub>66</sub>, and are normally zoned (Fig. 7c).

<sup>11</sup> 311 3.5  $H_2O$  in MIs

Using Raman spectroscopy, we estimated the water contents of 109 MIs hosted in orthopyroxene (43), clinopyroxene (15), and plagioclase (51) phenocrysts (Fig. 9a). In the 43 Opx-hosted MIs from units A, B, C and D, the H<sub>2</sub>O contents range from 2.5 wt.% to 5.8 wt.% (Fig. 9a; Supplementary Table 4). In the 15 Cpx-hosted MIs, the H<sub>2</sub>O contents range from 2.5 wt.% to 6 wt.% (Fig. 9a). In the 51 Pl-hosted MIs, the H<sub>2</sub>O contents range from 3.6 wt.% to 6.7 wt.% (Fig. 9a). Figure 9 shows that the H<sub>2</sub>O contents are mostly constant in pyroxene-hosted MIs from units A, B and C. Pyroxene-hosted MIs from unit D have slightly lower measured H<sub>2</sub>O contents, but are not significantly different from the earlier three units within the uncertainty of the measurements. Plagioclase-hosted MIs are characterised by near-uniform H<sub>2</sub>O contents across all units of the tephra fall deposit (Fig. 9a). 

The range of measured H<sub>2</sub>O contents in our melt inclusions (Fig. 9a) could indicate that diffusive re-equilibration of H<sub>2</sub>O occurred between the MIs and their external environment (e.g. Massare et al., 2002). However, the plagioclase- and pyroxene-hosted inclusions are often associated with large bubbles, which could be indicative of heterogeneous trapping of fluid and melt. Numerical models suggest that the glass phase in heterogeneously trapped inclusions is less susceptible to post-entrapment modification of volatile contents (Steele-Macinnis et al., 2017); therefore we consider it reasonable to assume that the range of the highest measured H<sub>2</sub>O contents is between 5.5 and 6.5 wt.% in plagioclase- and pyroxene-hosted MIs and is representative of the highest pre-eruptive melt water content (Fig. 9a). 4. Discussion 4.1 Pre-eruptive conditions and timescales of the Calbuco sub-Plinian eruption 4.1.1 Pre-eruptive temperatures and pressures Pre-eruptive temperature (T) conditions of the Calbuco magma reservoir have been calculated using the mineral-melt thermometers of Putirka (2008) for orthopyroxene, clinopyroxene and plagioclase. The plagioclase-liquid hygrometer-thermometer of Waters and Lange (2015) was used to better constrain the range of temperatures under which plagioclase cores was able to grow. Detailed results from thermometric calculations and tests for mineral-melt equilibrium are presented in Supplementary Table 5. All measured clinopyroxene compositions were tested for equilibrium against the mean bulk rock, plagioclase-hosted MIs (PEC-corrected MIs) and mean matrix glass compositions from units A-D. We evaluated equilibrium on the criterion  $K_D=0.28\pm0.08$  (Putirka, 2008). All the clinopyroxene crystals analysed in this study passed the equilibrium criterion with the less evolved MI compositions with  $0.25 \le K_D \le 0.26$ . Clinopyroxene crystallisation temperatures were calculated considering 5.5 and 6.5 wt.% H<sub>2</sub>O as the highest pre-eruptive melt water content 

<sup>829</sup> 347 (Fig. 9a). The clinopyroxene-MI liquid equilibration temperatures range between 961 and 988
<sup>831</sup> 348 °C (Eqn. 33; Putirka, 2008) (Fig. 10a).

All analysed orthopyroxene compositions were tested for equilibrium against the mean bulk rock, plagioclase-hosted MIs (PEC-corrected MIs) and mean matrix glass compositions. All the orthopyroxene crystals analysed passed the equilibrium criterion  $K_D=0.29\pm0.06$  (Putirka, 2008) with the less evolved MI compositions. Orthopyroxene crystallisation temperatures were then calculated assuming magmatic H<sub>2</sub>O contents of 5.5 and 6.5 wt.% (Fig. 9a). Calculated orthopyroxene-liquid equilibration temperatures range between 970 and 987 °C (Fig. 10a). These temperatures are similar to those obtained for clinopyroxene, suggesting that ortho- and clinopyroxene were grown under the same magmatic conditions (Fig. 10a). 

Further evidence to constrain pre-eruptive conditions was obtained using plagioclase-liquid thermometry. Plagioclase compositions were tested for equilibrium with the PEC-corrected MI and the mean matrix glass compositions and only equilibrium pairs with  $K_D(An-Ab)^{Pl-liq} =$ 0.10±0.05 at T <1050 °C were considered. Plagioclase phenocryst cores with An<sub>78-93</sub> were calculated to be in equilibrium with the less evolved MI compositions (with 5.5 and 6.5 wt.% H<sub>2</sub>O) at temperature between 927 and 963 °C (Fig. 10a). Plagioclase blocky microlites (with sizes from 25 to 250  $\mu$ m; An<sub>63-70</sub>) are also in equilibrium with the melt inclusion and matrix glass compositions at temperatures between 900 and 926 °C (Fig. 10a). 

The plagioclase-liquid hygrometer-thermometer (Waters and Lange, 2015) recovers and confirms that the plagioclase phenocryst cores  $(An_{84-93})$  are in equilibrium with the less evolved MI compositions at temperatures between 900 and 940 °C (Fig. 10a) and H<sub>2</sub>O between 5.4 and 6.1 wt. % (this amount is consistent with the maximum amount of  $H_2O$  measured in MIs). 

Lange, 2015) is  $\pm 12$  °C. This implies that the pre-eruptive temperature of the phenocrysts overlap within this degree of error. 

We note that plagioclase overgrowth rim compositions (characterised by oscillatory zoning) may result from disequilibrium crystallisation due to high undercooling, magma convection and thermal mixing, therefore temperatures were not calculated from these compositions.

Pre-eruptive conditions of the Calbuco magma reservoir also have been calculated using Rhyolite-MELTS software (version 1.2) (Ghiorso et al. 2012; Ghiorso and Gualda, 2015). The range of temperatures obtained from thermometers was used together with the highest amount of H<sub>2</sub>O measured in MIs (5.5 to 6.5 wt.% of H<sub>2</sub>O). The bulk rock composition was used for the simulations. In accordance with the  $H_2O$  solubility in basaltic andesitic melts predicted by Moore et al. (1998), the water-saturated conditions of this basaltic andesite magma with 5.5 and 6.5 wt.% of H<sub>2</sub>O would be reached at pressures between 230 and 320 MPa (Fig. 9b). Our simulations investigated a wider range of pressures in order to constrain the possible pre-eruptive P and T conditions where oxide, orthopyroxene, clinopyroxene and plagioclase phenocrysts were able to form (Supplementary Table 6). Particularly, pre-eruptive pressures and temperatures were constrained by replicating the same mineral assemblage and crystal volume fraction of natural samples (Supplementary Table 6). Rhyolite-MELTS simulations show that the observed crystal fraction of oxide, orthopyroxene, clinopyroxene and plagioclase cores were formed at 260-320 MPa (~10-12 km, in excellent agreement with pressures calculated from dissolved H<sub>2</sub>O contents) and temperatures between 940 and 950 °C (Supplementary Table 6). Considering the total phenocryst volume fraction, therefore, taking into account the plagioclase overgrowth rims and the blocky microlites, the pre-eruptive pressures predicted from Rhyolite-MELTS range between 230 and 320 MPa, whereas temperature ranges between 900 and 950 °C (Figure 10b; Supplementary Table 6). This means that the Calbuco magma was probably water-saturated before the eruption (Fig. 9b). 

Furthermore, the pre-eruptive temperatures obtained with Rhyolite-MELTS simulations indicate an overall cooling of the magma reservoir before the eruption.

The pre-eruptive temperatures of the magma obtained in this study (900 to 950 °C) are comparable with those estimated by Morgado et al. (2019) (910 to 970 °C). The pre-eruptive pressures obtained here (230 to 320 MPa) are higher than the 140 to 260 MPa estimated by Morgado et al. (2019). The pressures estimated by Morgado et al. (2019) are inconsistent with the H<sub>2</sub>O content of MIs (5.5 to 6.5 wt.%), as volatile solubility models would predict lower melt  $H_2O$  concentrations at these pressures. Therefore, only the highest pressures estimated by Morgado et al. (2019) are compatible with our measured H<sub>2</sub>O contents and with pressure-dependent H<sub>2</sub>O solubility models for basaltic andesitic and andesitic melt compositions (Moore et al., 1998). 

4.1.2 Closed vs open system 

Scoriae textures and plagioclase CSDs show that at least three events of crystal nucleation and growth occurred, producing in sequence: (1) phenocrysts (titanomagnetite, orthopyroxene, clinopyroxene and plagioclase cores; Figs 1 and 3); (2) overgrowth rims around plagioclase, and blocky microlites (plagioclase, clinopyroxene and orthopyroxene) with sizes from ~25 to ~250  $\mu$ m (Figs 1, 2 and 3); and (3) skeletal microlites <15  $\mu$ m (Figs 2 and 3). The distinction between the first and second events is based on the different chemical compositions of the plagioclase cores and the overgrowth rims and blocky microlites. On the other hand, the compositional similarity between overgrowth rims ( $An_{58-77}$ ) and blocky microlites ( $An_{63-70}$ ) suggests that they crystallised together under the same magmatic conditions (Fig. 10b). 

The normal zoning of the plagioclase cores is in agreement with an initial slow cooling of the magma in a closed system; however, water-saturated decompression also induces the crystallisation of more sodic compositions (Blundy and Cashman, 2001), causing a progressive 

- decrease of An content within the crystallising mineral (plagioclase core). The oscillatory
  zoning in plagioclase crystals shows resorption morphologies (Fig. 4) and sieve and patchy
  zoning textures (Fig. 4e, f), which may relate to different processes:
- (i) open-system processes, such as mafic magma recharge and chemical mixing (Ginibre et 1014
  al., 2002; Ruprecht and Wörner, 2007; Ginibre and Wörner, 2007; Ginibre et al., 2007);

1017 426 (ii) closed-system processes (closed-system for magma but not volatiles), such as
1018 1019 427 decompression, heat transfer, and thermal mixing in response to a thermal perturbation,
1020 1021 428 associated with a magma chamber overturn (Ginibre et al., 2002; Ruprecht and Wörner, 2007;
1023 1024 429 Ginibre and Wörner, 2007; Ginibre et al., 2007).

It is important to note that our chemical data show FeO and MgO remain constant whilst An content decreases from plagioclase cores to oscillatory zoned overgrowth rims (Fig. 8). The observation of resorption and growth morphologies (Fig. 4c), sieve textures and patchy zoning 1030 432 textures (Fig. 4e, f) in plagioclase crystals, combined with constant FeO and MgO contents 1032 433 1034 434 suggest that the oscillatory zoning was not caused by magma recharge and consequent magma mixing, but rather represents closed-system differentiation due to thermal mixing during magma convection (thermal gradients within the chamber) or decompression events (Ginibre et al., 2002; Ruprecht and Wörner, 2007; Ginibre and Wörner, 2007; Ginibre et al., 2007). We highlight that the evidence from zoned crystals supports a closed system before the 2015 eruption. Furthermore, no mafic enclaves and macroscopic magma heterogeneities are observed 1047 440 in the erupted products, indicating that magma intrusion immediately before the eruption was 1049 441 unlikely.

The crystallisation sequence is confirmed by whole-rock, melt inclusion and glass major element compositional trends, which are consistent with melt differentiation without magma mixing. Mass balance calculations computed using Petrograph software (Petrelli et al., 2005) show that the Calbuco bulk composition (basaltic andesite) can evolve to dacite through ~53% 

crystallisation plagioclase, orthopyroxene, clinopyroxene titanomagnetite of and (Supplementary Table 7). Trace element differentiation trends have also been modelled in agreement with Neuman et al. (1954) using Petrograph. The melt fraction was varied from 0.47 to 1 in agreement with textural analysis. Figure 6 shows that the observed trace element 1072 449 1074 450 variations in the Calbuco erupted products can be successfully described by closed-system melt 1076 451 differentiation. 

1078 452 

## 1080 453 *4.1.3 Multidisciplinary constraints*

Assuming a lithospheric pressure gradient and crustal density of 2.7 g/cm<sup>3</sup>, our calculated pressures of 230-320 MPa correspond to a magma reservoir located 8.5 to 12 km beneath Calbuco volcano. Although this is a wide depth range, the shallower part of this range is close 1089 457 to the recorded depths of pre-eruptive seismicity. Indeed, between 1 January and 21 April 2015, 147 volcano-tectonic (VT) earthquakes were recorded and located at depths of 6.5-8 km 1091 458 1093 459 beneath Calbuco (SERNAGEOMIN, 2015; Valderrama et al., 2015) (Figs 10a and 11; Supplementary Table 8). This activity marked a very small increase in energy when compared to the Calbuco seismic baseline. The low magnitude and frequency of earthquakes, and the absence of other coeval geophysical anomalies was not considered to be indicative of renewed unrest (Valderrama et al., 2015). On 22 April at 18:11 GMT (15:11 local time), 3 hours before the eruption, a seismic swarm consisting of ~140 VT events was recorded (SERNAGEOMIN, 1106 465 2015; Valderrama et al., 2015) at depths of ~7 km. The hypocentral depth of the pre-eruptive VT swarms is likely to indicate the rooftop of the magma chamber before the eruption (Figs 1108 466 1110 467 10a and 11). A magma reservoir located 8 to 11 km below the summit has also been inferred by modelling the co-eruptive subsidence recorded by Interferometric Synthetic Aperture Radar (InSAR) during the 22-23 April 2015 eruption (Delgado et al., 2017). Our pressures predicted by Rhyolite-MELTS simulations are consistent with these independent observations. 

## 4.1.4 Timescales of crystal growth

Plagioclase residence times were estimated by combining CSD slopes and experimental growth rates, as residence time is equal to  $-1/(\text{growth rate} \times \text{CSD slope})$ . We used the maximum ( $1.32 \times 10^{-2}$  mm/h) and the minimum ( $1.50 \times 10^{-3}$  mm/h) plagioclase growth rates estimated by Shea and Hammer (2013) for basaltic andesitic melts. We also assume a slow growth rate for plagioclase of  $3.96 \times 10^{-5}$  mm/h from a series of experimental growth rates, obtained from basaltic melts (Agostini et al., 2013; Arzilli et al., 2015), in order to take into account the slower crystallisation kinetics of plagioclase phenocrysts and zoned blocky microlites. Calculated residence times for plagioclase phenocrysts range from several days to ~8 months (Supplementary Table 1). For blocky microlites (25 to 250 µm) the residence time ranges between a few hours to  $\sim 2$  days (estimated with the faster growth rates), consistent with crystallisation experiments performed by Shea and Hammer (2013) in which plagioclase microlites were able to form over 12 to 48 hours. The maximum residence time estimated for blocky microlites is 2 months (Supplementary Table 1). Therefore, these timescales suggest that the population of blocky microlites (25-250 µm) may represent a final, crystallisation event before the onset of the eruption, and therefore before syneruptive decompression could begin. Furthermore, we suggest that skeletal microlites of plagioclase and pyroxene with size <15 µm (Fig. 3) were grown in syn-eruptive conditions during rapid magma ascent, i.e. during decompression-induced crystallisation (Blundy and Cashman, 2001), as their residence time ranges between a few minutes to ~3 hours. This is consistent with the time passed between the last seismic swarm and the eruption. This crystallisation event produced a further evolution of the residual melt towards dacitic compositions (measured in the matrix glass; Fig. 5).

Since all the samples analysed in the present study show similar textural, chemical and mineralogical features, it is possible to infer that the magma erupted during the first and second phases of the sub-Plinian eruption experienced the same pre-eruptive conditions.

4.2 Triggering of the Calbuco sub-Plinian eruption 1192 498

The Calbuco sub-Plinian eruption (22-23 April 2015) occurred with just three hours warning from geophysical data. This poses a clear challenge for volcanologists and risk managers: what are the conditions under which a volcanic system may produce a sudden unexpected violent eruption? In the Calbuco 2015 case, the pre-eruptive conditions and the timescales of the processes, constrained using petrological, geochemical and seismic data, allow us to compare and contrast three different triggering mechanisms: 

a) Second boiling induced by crystallisation prior to eruption due to the thermal mixing and 1207 505 convection within the magma chamber, i.e. an internal triggering mechanism in a closed 1209 506 system;

b) Gradual, slow decompression of the system and outgassing (e.g. Scandone, 1996; Ruprecht and Wörner, 2007; Ginibre and Wörner, 2007; Shea et al., 2009), which could have begun at least several hours/days before the sub-Plinian eruption, i.e. open system for volatiles but not magma; 

c) Injection of hot magma into the magma chamber (as suggested by Morgado et al., 2019) 1224 513 and magma mixing, which may be considered an archetype of the external triggering mechanism for explosive eruption, i.e. open system for volatiles and magma. 1226 514

4.2.1 Second boiling induced by protracted crystallisation 

The An, FeO and MgO profiles of plagioclase crystals (Fig. 8) indicate that the oscillatory zoning may have been produced by protracted resorption-regrowth events, caused by thermal 

mixing due to magma convection in a closed system (in agreement with Ginibre et al., 2002; Ruprecht and Wörner, 2007; Ginibre and Wörner, 2007; Ginibre et al., 2007). Rhyolite-MELTS 1247 521 simulations show that  $0.10 \le 0.15$  of high-anorthite plagioclase (An<sub>80-76</sub>) can be formed at 930-950 °C and 250-320 MPa (Fig.10b; Supplementary Table 6) from the basaltic andesitic 1249 522 liquid (Supplementary Table 2). This may represent the initial conditions of the system that induced the crystallization of plagioclase cores, as  $\phi$  and the An content are comparable to those analysed in the natural samples (Supplementary Table 3 and Supplementary Table 6). Rhyolite-MELTS simulations show that a further  $0.10 \le 0.25$  of low-anorthite plagioclase (An<sub>70-57</sub>) can be formed, in addition to the plagioclase core crystal content, at temperatures 1260 527 between 900 and 930 °C (Fig. 10b). Temperatures between 900 and 930 °C may represent the conditions where plagioclase oscillatory zoning and blocky microlites (25-250 µm) formed, as  $\phi$  and the An content are similar to those observed in the natural samples (Table 1 and Supplementary Table 3). Therefore, the oscillatory zoning observed within the plagioclase crystals may be the result of thermal mixing and magma convection driven by a thermal gradient of ~30 °C within the chamber. The protracted crystallisation of phenocrysts and 1275 534 blocky microlites (25-250 µm) may have induced second boiling (Blake, 1984; Tait et al., 1989; Fowler and Spera, 2008, 2010; Degruyter and Huber, 2014; Stock et al., 2016). In this 1277 535 scenario, the thermal convection of the magma may have favoured the formation of plagioclase oscillatory zoning, resorption morphologies (Fig. 4) and sieve and patchy zoning textures (Fig. 4e, f). A gradual cooling of the overall magmatic system and an increase in the total crystal content may have increased magma viscosity sufficiently to slow convection. We highlight that viscosity can increase by ~4 orders of magnitude (calculated using models of Giordano et al., 2008 and Vona et al., 2011) during an increase in crystal volume fraction from 15 to 45 vol% (considering all the crystal phases at pre-eruptive conditions). Following this stage, further 1292 542 1294 543 crystallisation and pressurisation from second boiling led to the eruption. The seismic activity 

produced shortly before eruption is likely to represent the pressure build-up prior to rapid magma ascent in the conduit. In the case of Calbuco, this interpretation is in agreement with the short-lived precursory seismic activity that occurred on 22 April, only  $\sim 3$  hours before the 1308 547 eruption, at depths of 6.5-8 km (SERNAGEOMIN, 2015; Valderrama et al., 2015; Fig. 11). 1310 548 Periodic seismic swarms at near-constant depths beneath a volcano may indicate cooling and solidification of a stationary magma (Chouet, 1979). Furthermore, no significant pre-eruptive 1312 549 deformation was detected up to one day before the eruption from the tiltmeter and InSAR interferograms (Valderrama et al., 2015; Delgado et al., 2017). We highlight that the presence of abundant exsolved volatiles in the closed-system magma chamber would increase magma compressibility, attenuating deformation signals associated with pressure increase (Johnson et al., 2000; Delgado et al., 2017). 

The original basaltic andesitic magma may have been evolved to a state in which eruption 1325 555 was triggered, without the need for an external triggering mechanism. Volatile exsolution is a 1327 556 key internal trigger for eruption-inducing over-pressurisation (Blake, 1984; Tait et al., 1989; Fowler and Spera, 2008, 2010; Degruyter and Huber, 2014; Tramontano et al., 2017). We note that Delgado et al. (2017) consider volatile exsolution in a closed system as a much more likely trigger mechanism for the Calbuco eruption than dike intrusion, based on the incongruence displayed by synthetic models, which account for a dipping dike to predict co-eruptive interferograms. This is also confirmed by the satellite imagery model developed by Pardini et 1342 563 al. (2018), which shows that magmatic volatiles were already exsolved before the onset of the Calbuco eruption. 1344 564

To investigate the potential overpressure generated by second boiling, we performed a mass balance calculation (following the model developed by Tait et al., 1989) on a magmatic body with fixed volume. For this calculation, we assumed isothermal conditions and a closed magmatic system, i.e. no mass is exchanged with the surrounding rocks. We considered a 

magmatic body at ~10 km (270 MPa), with a temperature of 950 °C, containing 6.0 wt.% magmatic H<sub>2</sub>O. The system is then forced to crystallise, causing a decrease in the mass of the melt in the magmatic body and a concentration of dissolved volatiles in the melt phase. This, in turn, increases the saturation pressure of the dissolved volatiles and, as soon as the saturation 1367 572 pressure exceeds the pressure of the reservoir, volatiles start to exsolve and pressurise the 1369 573 system. From the mass balance calculation we found that, once the magma reservoir is pressurised by exsolved volatiles, for each additional 1% of crystallisation there is a 2 MPa increase in pressure. The overpressure at which failure occurs depends on the tensile strength of the surrounding rocks, but typically ranges from a few MPa to 10-20 MPa (Tait et al., 1989). Therefore, crystallisation of 10 vol.% is sufficient to over-pressurise the system and trigger eruption. This also signifies that crystallisation and the resulting second boiling could theoretically lead a "silent" magma stored in the crust beneath the volcano towards conditions 1384 580 1386 581 favourable for the triggering of a large explosive eruption on timescales of hours to days. 

While this mechanism for eruption triggering has been proposed by several authors (Stock et al., 2016), this is, to our knowledge, the first occurrence where an internal triggering process may have been recognised in a sub-Plinian eruption at the Earth's surface. Recent eruptions from several continental margin and island arc volcanoes have been characterised by short-term (i.e. hours) precursory seismic activity, including the sub-Plinian eruptions of Redoubt 1989 (Alaska; Nye et al., 1994), Galunggung 1982 (Java; Bluth et al., 1994), Reventador 2002 (Ecuador; Hall et al., 2004), Cordón Caulle 2011-2012 (Chile; Castro et al., 2013) and the 1401 588 Plinian eruption of Chaitén 2008 (Chile; Castro et al., 2009; Pallister et al., 2013). We therefore 1403 589 1405 590 highlight that the internal triggering mechanism proposed for the 2015 Calbuco sub-Plinian eruption may be a common short-onset triggering mechanism for Plinian-type eruptions in continental and island arc volcanotectonic settings. Further studies should be performed in order to consider internal triggering mechanisms as potential causes for Plinian events. 

4.2.2 Slow decompression A second hypothesis of a possible triggering mechanism is the slow decompression of the magma (Shea et al., 2009) before the explosive eruption. This mechanism is supported by experimental evidence but not from geophysical observations, as a pre-eruptive decompression 1428 598 in the days prior to eruption would have produced much more significant precursory signals (Scandone 1998; Francis et al. 1993). Focusing first on the supporting evidence we highlight that the mineral assemblage in the scoriae is more in agreement with the results obtained by decompression experiments performed by Shea and Hammer (2013) than those obtained from cooling experiments. Particularly, the absence of amphibole in the microlite assemblage may be related to decompression: in fact, amphibole forms more readily during cooling and tends to destabilize during decompression (Rutherford and Hill, 1993; Shea and Hammer, 2013). 1443 605 Furthermore, resorption-regrowth morphologies (Fig. 4c), and sieve and patchy zoning textures 1445 606 (Fig. 4e, f) in plagioclase phenocrysts, associated with significant An variations and constant FeO and MgO contents (Fig. 8), may suggest several decompression events and magma degassing should have occurred before the eruption (in agreement with Ginibre et al., 2002; Ruprecht and Wörner, 2007; Ginibre and Wörner, 2007; Ginibre et al., 2007). Our CSD and residence time analysis (Supplementary Table 1), in agreement with the experiments of Shea and Hammer (2013), shows that blocky microlites (25 to 250  $\mu$ m) may be formed in a few hours/days. However, this would require a slower decompression than can be provided by the 1460 613 1462 614 eruption itself. Furthermore, there is no evidence for slow magma ascent and degassing, with associated seismic activity and deformation prior to the eruption (Delgado et al., 2017). 1464 615 4.2.3 Injection of fresh magma into the magma chamber and magma mixing 

Injection of fresh magma into the magma chamber may be included amongst the possible triggering mechanisms of the Calbuco eruption. Morgado et al. (2019) suggested that localised heating, consistent with an injection of hot magma, could have triggered the 2015 Calbuco eruption. Data reported by Morgado et al. (2019) show that the injection of hot magma 1485 621 increased the temperature of only a small portion of the base of the magma reservoir by 70-200 1487 622 °C (reaching up to 1070 °C), whilst the remaining magma was not thermally affected. This conclusion is based on Fe-Ti oxide thermometry obtained from just one sample extracted from a pyroclastic density current deposit (Morgado et al., 2019). If a sudden injection of fresh magma into an evolved magma reservoir was the trigger of the eruption, textures of the samples in the overall eruption deposit might be expected to record evidence of magma mixing and mingling. However, the textural, chemical and mineralogical features of the scoriae show no clear evidence of mixing or mingling; mafic enclaves are absent; and no evidence such as 1502 629 1504 630 compositional layering (Cioni et al., 1995) has been found in the volcanic deposits. Therefore, 1506 631 if injection of mafic magma occurred, efficient magma mixing is required to homogenize the magma composition. Injection of mafic magma may have occurred months/years before the eruption, but we do not see any evidence (e.g. mingling textures) of a sudden magma recharge before the eruption. Indeed, the significant An variations and the constant FeO and MgO contents across plagioclase oscillatory zones (Fig. 8) are not likely to have formed through compositional mixing, but rather they result from thermal mixing combined with an overall 1519 637 cooling of the system and/or decompression (Ginibre et al., 2002; Ruprecht and Wörner, 2007; Ginibre and Wörner, 2007; Ginibre et al., 2007). 1521 638 5. Conclusion and implications 

Prior to the 2015 sub-Plinian eruption of Calbuco, a basaltic andesite magma was stored at depths between 8 and 12 km (i.e. 230-320 MPa) beneath the volcano. We estimate that the depths between 8 and 12 km (i.e. 230-320 MPa) beneath the volcano. We estimate that the 

initial temperature of the magma was between 900 and 950 °C and that crystallisation occurred under H<sub>2</sub>O-saturated conditions (5.5-6.5 wt.% H<sub>2</sub>O). Crystallisation occurred under these conditions, producing phenocrysts of titanomagnetite, orthopyroxene, clinopyroxene and 1544 646 plagioclase cores (An<sub>78-93</sub>). Plagioclase overgrowth rims (An<sub>58-77</sub>) and blocky microlites (from 25 to 250 µm) grew during a second crystallisation event, which was caused by thermal mixing 1546 647 and magma convection within the chamber. A lack of geophysical signals precludes a slow decompression explanation for the blocky microlites. The rapid onset of a sub-Plinian eruption occurred with no evidence of magma intrusion, magma mixing and precursory deformation, and was, instead, caused by protracted crystallisation and second boiling over-pressurisation. Such a triggering mechanism explains how violent sub-Plinian eruptions may occur with very little warning from geophysical monitoring. Ephemeral geophysical and seismic signals followed by unexpected violent eruptions should be more systematically studied in order to carefully evaluate the risks posed by apparently peaceful volcanoes. 1563 655 1565 656 Acknowledgements We gratefully acknowledge funding support from the RCUK NERC DisEqm project 1578 662 (NE/N018575/1). The research leading to these results has received funding from the European 1580 663 Research Council under the European Union's Seventh Framework Programme (FP/2007-2013) / ERC Grant Agreement n. 279802 and from ERC Consolidator Grant 612776-CHRONOS. We thank Paul Lythgoe and Alastair Bewsher for their assistance with X-ray fluorescence analysis. We thank T. Milian for support during the writing of the manuscript. We 

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2222	907	Figure captions
2224	908	Figure 1. Stratigraphy of the tephra fall deposit from the 2015 Calbuco eruption, and textural
2220 2227 2228	909	features of scoriae. Left: schematic stratigraphy (not to scale) of the DF10 proximal deposit
2229 2230	910	showing the four units of the tephra fall deposit (re-drawn after Romero et al., 2016). Unit A is
2231 2232	911	the tephra fall deposit of the first phase of the eruption on 22 April 2015. Units B, C and D are
2233 2234	912	the deposits of the second phase on 23 April 2015. Bulk textures of typical scoria clasts from
2235 2236	913	each unit are shown in BSE images at the centre, while details of the groundmass (plagioclase
2237 2238 2239	914	and pyroxene microlites and the glassy matrix) are shown on the right.
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Figure 2. Textural features of glassy matrices in scoriae from units A, B, C and D showing
skeletal plagioclase and pyroxene microlites <15 μm.</li>

919 Figure 3. Crystal size distributions (CSDs) of Calbuco scoriae derived by CSDCorrections
920 1.6 software (Higgins 2000, 2002).

Figure 4. BSE images of plagioclase zoning textures. (a) Plagioclase phenocryst characterised by a large core with normal zoning and an overgrowth rim with oscillatory zoning. (b-c-d) Plagioclase phenocrysts characterised by a large overgrowth rim with oscillatory zoning, in which several resorption and regrowth layers can be observed. Red arrows in (c) indicate the last resorption event. (e-f) Overgrowth rims can be characterised by sieve textures (e) and patchy zoning textures (e-f), in which melt inclusions can be hosted.

Figure 5. Harker-style diagrams of major element compositional variability in products from the 2015 Calbuco eruption: bulk composition of juvenile scoriae, plagioclase-hosted melt inclusions, and matrix glass. Plagioclase-hosted melt inclusions were corrected for postentrapment crystallisation following the method of Neave et al. (2017). The diagram reports chemical data collected for each unit (A, B, C and D).

Figure 6. Trace element variability measured in the bulk rock, plagioclase-hosted melt inclusions, and matrix glass of products from the 2015 Calbuco eruption. Melt differentiation trends driven by fractional crystallisation are modelled assuming the mean bulk rock composition of juvenile scoriae as the starting composition. The melt fraction was varied from 0.47 to 1. For the fractional crystallisation model the bulk partition coefficient ( $D^b$ ) of Zr ranges

between 0.1 and 0.15,  $D^{b}_{Rb}$  is 0.06 (a),  $D^{b}_{Sr}$  is between 1.2 and 1.75 (b),  $D^{b}_{Y}$  is 0.25 (c),  $D^{b}_{Nb}$  is 0.05 (d),  $D^{b}_{Ba}$  is 0.4 (e) and  $D^{b}_{La}$  is 0.1 (f) (Fujimaki et al., 1984; Dunn and Sen, 1994; Aigner-Torres et al., 2007). The ticks along the fractional crystallisation path indicate the melt fraction (each tick is 0.1). 2313 944 2315 945 Figure 7. (a) Anorthite contents of plagioclase crystals in the scoriae of the 2015 Calbuco eruption, showing the compositions of phenocrysts (cores, overgrowth with oscillatory zones and external rims) and microlites. Kernel density distributions for the plagioclase populations are reported as solid line. (b) Compositions of pyroxene crystals in Calbuco scoriae, Di: diopside; Hd: hedenbergite; En: enstatite; Fs: ferrosilite. (c) Compositions of olivine crystals in Calbuco scoriae. All diagrams report mineral compositions from each tephra fall unit (A, B, C 2328 951 and D). 2330 952 2332 953 Figure 8. Electron microprobe quantitative point analysis profiles across plagioclase crystals from core to rim. Profile for An (mol.%), FeO (wt.%) and MgO (wt.%) are reported in the <sup>2336</sup> 955 figure: (a-b) plagioclase #8 of unit A; (c-d) plagioclase #12 of unit C; (e-f) plagioclase #6 of unit D. Figure 9. (a) H<sub>2</sub>O dissolved in melt inclusions (MI) hosted in orthopyroxene, clinopyroxene, and plagioclase crystals, measured by Raman spectroscopy. Kernel density distributions for H<sub>2</sub>O dissolved in MIs are reported. (b) H<sub>2</sub>O solubility in basaltic andesitic and andesitic melts 2347 960 as a function of pressure, calculated following Moore et al. (1998). Basaltic andesite is the bulk composition erupted during the 2015 Calbuco eruption; andesite is the composition of plagioclase-hosted melt inclusions and represents the pre-eruptive melt composition. 

Figure 10. Pre-eruptive conditions of the 2015 Calbuco eruption. (a) The orthopyroxene-liquid thermometer (Putirka, 2008) was used to estimate the crystallisation temperatures of orthopyroxene phenocrysts. Pre-eruptive temperatures for cores and blocky microlites were obtained using the plagioclase-liquid thermometer of Putirka (2008) and the plagioclase-liquid hygrometer-thermometer of Waters and Lange (2015). The range of pre-eruptive pressures was 2372 969 predicted with Rhyolite-MELTS software (version 1.2) (Ghiorso et al. 2012; Ghiorso and Gualda, 2015), investigating the highest amount of H<sub>2</sub>O (5.5 to 6.5 wt.% H<sub>2</sub>O) measured in pyroxene- and plagioclase-hosted MIs and the temperatures obtained from mineral-liquid thermometers. P2008: Putirka, 2008. W&L2015: Waters and Lange, 2015. The 1<sub>o</sub> calibration uncertainty on the Putirka (2008) thermometer is ±48 °C. The Waters and Lange (2015) hygrometer has a  $1\sigma$  temperature uncertainty of  $\pm 12$  °C. The red horizontal bar shows the 2387 976 pressure range of the major volcanotectonic (VT) earthquakes that preceded the eruption. (b) Rhyolite-MELTS simulations indicate that the plagioclase crystal fraction ( $\phi_{Pl}$ ) ranges between 2389 977 0.10 and 0.35 at the pre-eruptive conditions, in agreement with  $\phi_{Pl}$  calculated in the natural scoriae produced from the 2015 Calbuco sub-Plinian eruption. The plagioclase liquidus in H<sub>2</sub>O-saturated basaltic andesite magma has been obtained from Shea and Hammer (2013). 

2398 981

Figure 11. Localization depths of major volcanotectonic (VT) earthquakes at Calbuco 2400 982 reported in monthly (diamond), daily (dots), and bi-weekly (square) reports from 2402 983 2404 984 SERNAGEOMIN (available at http://www.sernageomin.cl/) during 2015. The colour bar 2406 985 indicates the number of events. The volcano was monitored by five seismometers, with the closest station 5 km to the west. 



# In (population density)













olivine Fo mol.%







![](_page_49_Figure_0.jpeg)

P (MPa)

![](_page_50_Figure_0.jpeg)

![](_page_51_Picture_0.jpeg)

Supplementary Figure 1. Calbuco volcano located in the Southern Andes of Chile. (a) Location map of Calbuco volcano (red triangle) and major tectonic features such as the Chile trench and Liquiñe-Ofqui Fault Zone (LOFZ) are reported together with other stratovolcanoes (green triangles). The 5 mm isopach of the 2015 eruption (Romero et al., 2016) is represented by the brown shadow, showing the areal distribution of the tephra fall deposit. Major cities are indicated as gridded polygons. (b) Photograph of Calbuco volcano during the 22 April 2015 eruption (first pulse) as seen from the southwest (photograph by Roger Smith).

Sample	Sample Texture with crystals >250 μm			Groundmass with crystals <250 μm			Total phases			
	φPl	φPx (Opx+Cpx)	φTi-Mag	φPl microlite	φPx microlite	φgl	φPl	φPx (Opx+Cpx)	φTi-Mag	φgl
DF10A	0.24	0.09	0.01	0.21	0.04	0.40	0.46	0.14	0.01	0.40
DF10B	0.24	0.09	0.01	0.19	0.03	0.44	0.43	0.12	0.01	0.44
DF10C	0.23	0.07	0.01	0.17	<<0.01	0.52	0.40	0.07	0.01	0.52
DF10D	0.27	0.07	0.01	0.20	0.04	0.40	0.47	0.11	0.01	0.40

Table 1: Crystal fraction of Calbuco scoriae of the DF10 proximal deposit.

note: the crystal fraction ( $\phi$ ) of each phase is calculated on a vesicle-free basis. Texture with crystals >200 $\mu$ m includes phenocrysts (generally from ~250 to ~1500  $\mu$ m). Groundmass with crystals <250  $\mu$ m includes microlites (from 1 to ~250  $\mu$ m) and patches of glass (residual melt). In the groundmass  $\phi$ Ti-Mag is <<0.01. The crystal fraction of olivine crystals is <0.01 and it is included in  $\phi$ Px. Pl: plagioclase; Px: pyroxene; Opx: orthopyroxene; Cpx: clinopyroxene; Ti-Mag: titanomagnetite; gl: glass.

0	The second	Slope				
Sample	Intercept (mm <sup>-4</sup> ) <sup>a</sup>	(mm <sup>-1</sup> ) <sup>a</sup>	t <sub>r</sub> min (h)	t <sub>r</sub> mean (h)	t <sub>r</sub> max (h)	t <sub>r</sub> max (days)
unit A-250 µm <plg<1500 td="" µm<=""><td>4.62</td><td>-3.81</td><td>20</td><td>175</td><td>5608</td><td>234</td></plg<1500>	4.62	-3.81	20	175	5608	234
unit A - 25 µm <plg<250 td="" µm<=""><td>7.48</td><td>-12.44</td><td>6</td><td>54</td><td>1718</td><td>72</td></plg<250>	7.48	-12.44	6	54	1718	72
unit A - 1 µm <plg<15 td="" µm<=""><td>13.7</td><td>-319.85</td><td>0.2</td><td>2</td><td>67</td><td>3</td></plg<15>	13.7	-319.85	0.2	2	67	3
unit B - 250 μm <plg<1500 td="" μm<=""><td>4.55</td><td>-4.04</td><td>19</td><td>165</td><td>5289</td><td>220</td></plg<1500>	4.55	-4.04	19	165	5289	220
unit B - 25 μm <plg<250 td="" μm<=""><td>8.38</td><td>-12.90</td><td>6</td><td>52</td><td>1656</td><td>69</td></plg<250>	8.38	-12.90	6	52	1656	69
unit B - 1 µm <plg<15 td="" µm<=""><td>15.82</td><td>-221.69</td><td>0.3</td><td>3</td><td>96</td><td>4</td></plg<15>	15.82	-221.69	0.3	3	96	4
unit C - 250 μm <plg<1500 td="" μm<=""><td>4.39</td><td>-4.03</td><td>19</td><td>165</td><td>5302</td><td>221</td></plg<1500>	4.39	-4.03	19	165	5302	221
unit C - 25 µm <plg<250 td="" µm<=""><td>8.73</td><td>-17.95</td><td>4</td><td>37</td><td>1190</td><td>50</td></plg<250>	8.73	-17.95	4	37	1190	50
unit D - 250 μm <plg<1500 td="" μm<=""><td>4.62</td><td>-3.91</td><td>19</td><td>170</td><td>5465</td><td>228</td></plg<1500>	4.62	-3.91	19	170	5465	228
unit D - 25 µm <plg<250 td="" µm<=""><td>9.83</td><td>-21.72</td><td>3</td><td>31</td><td>984</td><td>41</td></plg<250>	9.83	-21.72	3	31	984	41
Experimental growth rates			SH2013 (mm/h)	SH2013 (mm/h)	Ag2013- Ar2015 (mm/h)	
			1.32E-02	1.50E-03	3.96E-05	

**Supplementary Table 1.** Crystal size distribution (CSD) parameters and residence times of Calbuco rocks from CSD measurements.

Note: aIntercept and slope were calculated from a linear regression of the population density distributions, which are derived by CSD Corrections 1.6 program (Higgins 2000, 2002). Residence times  $(t_r)$  are calculated through the equation:  $t_r = (-1/\text{growth rate} \times \text{slope})$ . The minimum, mean and maximum residence times are  $t_r$  min,  $t_r$  mean and  $t_r$  max, respectively. Crystal growth rates obtained from Shea and Hammer (2013; *SH2013*) are used to estimate the minumum and the mean residence times. The maximum residence time is obtained considering the slowest growth rates estimated from Agostini et al. (2013; *Ag2013*) and Arzilli et al. (2015; *Ar2015*) in order to take into account the crystallization kinetics of plagioclase phenocrysts and zonations. Growth rates obtained from experiments are italicized.

**Supplementary Table 5.** Pre-eruptive temperatures calculated using geothermometers (Putirka, 2008) and the hygrometer-thermometer (Waters and Lange, 2015).

Unit of the tephra fall deposit	Cpx-liquid geothemometer	H <sub>2</sub> O (wt.%)	Eqn. 33 (°C)	K <sub>D</sub> (Fe-Mg) 0.28+0.08
unit A	MI - Cpx	5.5	975-985	0.26
unit A	MI - Cpx	6.5	964-974	0.26
unit B	MI - Cpx	5.5	973-985	0.26
unit B	MI- Cpx	6.5	961-973	0.26
unit C	MI - Cpx	5.5	973-988	0.25 -0.26
unit C	MI - Cpx	6.5	962-977	0.25 - 0.26
unit D	MI - Cpx	5.5	974-977	0.26
unit D	MI - Cpx	6.5	963-966	0.26

#### Geothermometer clinopyroxene-liquid

Note: equilibrium temperatures have been calculated using the clinopyroxene-liquid geothermometer of Putirka (2008). For the estimation of temperature Eqn. 33 (Putirka, 2008) has been used. The maximum amount of water dissolved in pyroxene- and plagioclase-hosted melt inclusions (5.5 wt.% and 6.5 wt.% of H<sub>2</sub>O respectively; Fig. 9a) was used to test clinopyroxene phenocrysts against plagioclase-hosted melt inclusions (MI) compositions. Test for equilibrium is based on Fe-Mg exchange coefficients. The Fe-Mg exchange coefficients yield  $K_D$ (Fe-Mg)<sup>Cpx-liq</sup> = 0.28±0.08 (Putirka, 2008). The Putirka (2008) thermometer model (Equation 33) predicts temperature to within ±48 °C.

Unit of the tephra fall deposit	Opx-liquid geothemometer	H <sub>2</sub> O (wt.%)	Eqn. 28b (°C)	K <sub>D</sub> (Fe-Mg) (0.29+0.06)
unit A	MI - Opx	5.5	984-987	0.24 - 0.31
unit A	MI - Opx	6.5	970-974	0.24 - 0.31
unit B	MI - Opx	5.5	984-987	0.26 -0.28
unit B	MI - Opx	6.5	970-974	0.26 - 0.28
unit C	MI - Opx	5.5	984-987	0.28 - 0.30
unit C	MI - Opx	6.5	970-974	0.28 - 0.30
unit D	MI - Opx	5.5	984-987	0.24 - 0.27
unit D	MI - Opx	6.5	970-974	0.24 - 0.27

### Geothermometer orthopyroxene-liquid

Note: equilibrium temperature have been calculated using the orthopyroxene-liquid geothermometer (Putirka, 2008). The maximum amount of water dissolved in pyroxene- and plagioclase-hosted melt inclusions (5.5 wt.% and 6.5 wt.% of H<sub>2</sub>O respectively; Fig. 9a) was used to test orthopyroxene phenocrysts against plagioclase-hosted melt inclusions (MI) compositions. Test for equilibrium is based on Fe-Mg exchange coefficients. The Fe-Mg exchange coefficients yield  $K_D$ (Fe-Mg)<sup>Opx-liq</sup> = 0.29±0.06 (Putirka, 2008). The Putirka (2008) thermometer model (Equation 28b) predicts temperature to within ±48 °C.

Unit of the tephra fall deposit	Pl-liquid geothemometer	H <sub>2</sub> O (wt.%)	Eqn. 23 (°C)	K <sub>D</sub> (Ab-An) T<1050 (0.10+0.05)
unit A	MI- Pl phenocrysts (core)	5.5	962-963	0.06
unit A	MI- Pl phenocrysts (core)	6.5	929-9630	0.056
unit A	MI- Pl microlite	6.5	915-926	0.11 - 0.12
unit A	RM- Pl microlite	6.5	900-902	0.08 - 0.09
unit B	MI- Pl phenocrysts (core)	5.5	961-963	0.06 - 0.08
unit B	MI- Pl phenocrysts (core)	6.5	929-931	0.06- 0.08
unit C	MI- Pl phenocrysts (core)	5.5	962-963	0.06 - 0.08
unit C	MI- Pl phenocrysts (core)	6.5	930-931	0.06 - 0.08
unit D	MI - Pl phenocrysts (core)	5.5	959-963	0.07 - 0.10
unit D	MI - Pl phenocrysts (core)	6.5	927-930	0.07 - 0.10

#### Geothermometer plagioclase-liquid

Note: equilibrium temperatures have been calculated using the plagioclase-liquid geothermometer (Putirka, 2008). The maximum amount of water dissolved in pyroxene- and plagioclase-hosted melt inclusions (5.5 wt.% and 6.5 wt.% of H<sub>2</sub>O respectively; Fig. 9a) was used to test plagioclase phenocrysts against MI compositions. The maximum amount of water dissolved in plagioclase-hosted melt inclusions (6.5 wt.% H<sub>2</sub>O; Fig. 9a) was used to test plagioclase microlites against residual melt (RM) compositions. Test for equilibrium is based on Ab-An exchange coefficients. The Ab-An exchange coefficients yield  $K_D(Ab-An)^{Pl-liq} = 0.10\pm0.05$  (T<1050 °C) (Putirka, 2008). The cores of Pl phenocrysts (An<sub>78-93</sub>) are in equilibrium with the less evolved MI compositions. Plagioclase microlites (An<sub>63-70</sub>) are in equilibrium with the residual melt (RM) compositions. The Putirka (2008) thermometer model (Equations 23 and 24a) predicts temperature to within  $\pm 48$  °C.

Unit of the tephra fall deposit	Liquid-Pl	H <sub>2</sub> O (wt.%)	An (mol.%)	T (°C)
unit A	MI - Pl (C)	5.5 - 6.1	84 - 93	920 - 940
unit B	MI - Pl (C)	5.5 - 6.0	85 - 88	910 - 930
unit C	MI - Pl (C)	5.4 - 6.1	84 - 90	900 - 940
unit D	MI - Pl (C)	5.6 - 5.7	85 - 88	920 - 930

#### Hygrometer-thermometer plagioclase-liquid

Note: equilibrium temperatures have been calculated using the hygrometer-thermometer plagioclase-liquid (Waters and Lange, 2015). The maximum amount of water dissolved in pyroxene- and plagioclase-hosted melt inclusions (5.5 wt.% and 6.5 wt.% of H<sub>2</sub>O respectively; Fig. 9a) was used to test plagioclase phenocrysts against plagioclase-hosted melt inclusions (MI) compositions. The cores (C) of Pl phenocrysts (An<sub>84.93</sub>) are in equilibrium with the less evolved MI compositions at temperatures between 900 and 940 °C, where the amount of H<sub>2</sub>O ranges from 5.4 to 6.1 wt. %. Application of the hygrometer model as a thermometer recovers temperatures to within  $\pm 12$  °C, on average.

Supplementary Table 6. Rhyolite-MELTS simulations to investigate the pre-eruptive conditions of the Calbuco eruption.

Crystal fraction of Calbuco samples								
Ox	Px (Onx-Cnx)	Pl (core)	Pl (core + overgrowth rim)	Pl (core + overgrowth rim + blocky microlites)				
0.01	0.07 - 0.09	0.10 - 0.14	0.23 - 0.27	0.33 - 0.37				
	Comples	Ox         Px (Opx-Cpx)           0.01         0.07 - 0.09	Ox         Px (Opx-Cpx)         Pl (core)           0.01         0.07 - 0.09         0.10 - 0.14	Ox         Px (Opx-Cpx)         Pl (core)         Pl (core + overgrowth rim)           0.01         0.07 - 0.09         0.10 - 0.14         0.23 - 0.27				

#### Rhyolite-MELTS simulations

T (°C)	P (MPa)	H <sub>2</sub> O (wt.%)	φOx	фОрх	фСрх	φPl
900	320	6.45	0.03	0.1	0.02	0.27
900	310	6.35	0.03	0.1	0.02	0.28
900	300	6.25	0.03	0.1	0.02	0.29
900	290	6.1	0.03	0.1	0.02	0.3
900	280	6	0.03	0.1	0.02	0.31
900	270	5.9	0.03	0.1	0.02	0.31
900	260	5.8	0.04	0.1	0.02	0.32
900	250	5.7	0.04	0.1	0.02	0.33
900	240	5.6	0.04	0.1	0.02	0.34
900	230	5.5	0.04	0.11	0.02	0.35
910	320	6.45	0.03	0.09	0.02	0.23
910	310	6.35	0.03	0.09	0.02	0.24
910	300	6.25	0.03	0.09	0.02	0.25
910	290	6.1	0.03	0.1	0.02	0.26
910	280	6	0.03	0.1	0.01	0.27
910	270	5.9	0.03	0.1	0.01	0.28
910	260	5.8	0.04	0.1	0.01	0.29
910	250	5.7	0.04	0.1	0.01	0.30
910	240	5.6	0.04	0.1	0.01	0.31
910	230	5.5	0.04	0.1	0.01	0.32
920	320	6.45	0.03	0.08	0.02	0.19
920	310	6.35	0.03	0.08	0.02	0.21
920	300	6.25	0.03	0.08	0.02	0.22
920	290	6.1	0.03	0.08	0.01	0.23
920	280	6	0.03	0.09	0.01	0.24
920	270	5.9	0.03	0.09	0.01	0.25
920	260	5.8	0.04	0.09	0.01	0.26
920	250	5.0	0.04	0.09	0.01	0.27
920	240	5.6	0.04	0.09	0.01	0.28
920	230	5.5	0.04	0.09	0.01	0.29
930	320	6.45	0.03	0.07	0.02	0.16
930	310	6.35	0.03	0.07	0.02	0.17
930	300	6.25	0.03	0.07	0.02	0.18
930	290	6.1	0.03	0.07	0.01	0.19
930	280	6	0.03	0.07	0.01	0.20
930	270	59	0.03	0.08	0.01	0.21
930	260	5.8	0.03	0.08	0.01	0.22
930	250	5.7	0.03	0.08	0.01	0.23
930	240	5.6	0.03	0.08	0.01	0.25
930	230	5.5	0.04	0.08	0.01	0.26
940	320	6.45	0.03	0.05	0.02	0.11
940	310	6.35	0.03	0.05	0.02	0.12
940	300	6.25	0.03	0.05	0.02	0.13
940	290	6.1	0.03	0.06	0.01	0.14
940	280	6	0.03	0.06	0.01	0.15
940	270	5.9	0.03	0.06	0.01	0.17
940	260	5.8	0.03	0.06	0.01	0.18
940	250	5.7	0.04	0.07	0.01	0.19
940	240	5.6	0.04	0.07	0.01	0.21
940	230	5.5	0.04	0.07	0.01	0.22

950	320	6.45	0.03	0.04	0.02	0.10
950	310	6.35	0.03	0.04	0.02	0.10
950	300	6.25	0.03	0.04	0.02	0.10
950	290	6.1	0.03	0.04	0.02	0.10
950	280	6	0.03	0.04	0.01	0.11
950	270	5.9	0.03	0.05	0.01	0.12
950	260	5.8	0.03	0.05	0.01	0.13
950	250	5.7	0.03	0.05	0.01	0.15
950	240	5.6	0.04	0.06	0.01	0.17
950	230	5.5	0.04	0.06	0.01	0.18
960	320	6.45	0.03	0.03	0.02	0.07
960	310	6.35	0.03	0.03	0.02	0.07
960	300	6.25	0.03	0.03	0.02	0.07
960	290	6.1	0.03	0.03	0.02	0.07
960	280	6	0.03	0.03	0.02	0.07
960	270	5.9	0.03	0.03	0.01	0.07
960	260	5.8	0.03	0.03	0.01	0.09
960	250	5.7	0.03	0.03	0.01	0.11
960	240	5.6	0.03	0.04	0.01	0.12
960	230	5.5	0.03	0.04	0.01	0.13
970	320	6.45	0.03	0.02	0.03	0.04
970	310	6.35	0.03	0.02	0.03	0.04
970	300	6.25	0.03	0.02	0.02	0.04
970	290	6.1	0.03	0.02	0.02	0.04
970	280	6	0.03	0.01	0.01	0.05
970	270	5.9	0.03	0.01	0.01	0.05
970	260	5.8	0.03	0.01	0.01	0.05
970	250	5.7	0.03	0.01	0.01	0.05
970	240	5.6	0.03	0.02	0.01	0.07
970	230	5 5	0.03	0.02	0.01	0.08

Note: Pre-eruptive conditions of the Calbuco magma reservoir have been calculated using Rhyolite-MELTS software (Ghiorso et al. 2012; Ghiorso and Gualda, 2015). The range of temperatures obtained from thermometers (Supplementary Table 5) was used together with the maximum amount of  $H_2O$  measured in MIs (5.5 to 6.5 wt.% of  $H_2O$ ; Supplementary Table 4). The bulk rock composition was used for the simulations (Supplementary Table 2). Preeruptive pressures and temperatures were constrained by obtaining the same mineral assemblage and phenocryst volume fraction of natural samples. Rhyolite-MELTS simulations show that the observed crystal fraction of oxide, orthopyroxene, clinopyroxene and plagioclase cores were formed at 260-320 MPa and temperatures between 940 and 950 °C (bold blue). Considering also the plagioclase overgrowth rims and blocky microlites, the pre-eruptive pressures predicted from Rhyolite-MELTS range between 230 and 320 MPa, whereas temperature ranges between 900 and 950 °C (bold red).

	Bulk composition	Residual melt	Residuals	Removed mineral phases	Amount of subtracted mineral phases (wt.%)	Total amount of subtracted phases (wt.%)	Sum of the squares of residuals
SiO <sub>2</sub>	55.74	64.95	0.213	Plagioclase	67.07	53.13	0.67
TiO <sub>2</sub>	0.88	1.37	0.233	Orthopyroxene	19.97		
$Al_2O_3$	18.23	14.59	0.153	Clinopyroxene	7.25		
FeO	8.22	7.25	0.134	Ti-Magnetite	5.71		
MnO	0.13	0.14	0.012				
MgO	4.07	1.52	0.005				
CaO	7.64	4.00	0.080				
Na <sub>2</sub> O	4.17	4.54	-0.718				
K <sub>2</sub> O	0.75	1.29	-0.109				
$P_2O_5$	0.18	0.35	-0.000				
Total	100	100					

#### Supplementary Table 7: Mass balance calculation.

Note: results of mass balance calculation were obtained using the method of Stormer and Nicholls (1978). For the bulk rock composition was chosen one of the less evolved composition analyzed in the unit D (DF2D\_ves; Supplementary Table 2). The matrix glass is one of the most evolved melt analyzed in the unit D. Results of mass balance calculation show that the melt differentiation is a possible process in the Calbuco magma. For the calculation we used the least squares mass balance (Stormer and Nicholls, 1978) option in the PetroGraph software (Petrelli et al., 2005).

#### Reference

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