

Mantle wedge exhumation beneath the Dora-Maira (U)HP dome unravelled by local earthquake tomography (Western Alps)

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Abstract

The behaviour of the mantle wedge of continental subduction zones during exhumation of ultra-high pressure (UHP) rocks is still poorly understood. Here, we shed light on this issue by a detailed analysis of the velocity structure of the mantle wedge beneath the Dora-Maira (U)HP dome in the Western Alps, based on local earthquake tomography independently validated by receiver function analysis. Our results point to a composite structure of the mantle wedge above the subducted European lithosphere. The Dora-Maira (U)HP dome lays directly above partly serpentinized peridotites ($V_p \sim 7.5$ km/s; $V_p/V_s = 1.70-1.72$), documented from ~ 10 km depth down to the top of the eclogitized lower crust of the European plate. To the east, these serpentinized peridotites are juxtaposed against dry mantle peridotites of the Adriatic upper plate along an active fault rooted in the lithospheric mantle. We propose that peridotite serpentinization was due to fluids released from the subducting Alpine slab to the Adriatic mantle wedge. Seismic velocities point to a minor amount of eclogitic metasediments above the European slab, which suggests that the subduction wedge was largely exhumed during late Eocene transtension along the Western Alps subduction zone. Meanwhile, part of the Adriatic mantle wedge was also exhumed at shallow

27 crustal levels, to be finally indented under the Alpine metamorphic units in the early Oligocene. Our
28 results suggest that mantle-wedge exhumation may represent an important feature of the deep
29 structure of exhumed continental subduction zones. The deep orogenic levels here imaged by
30 seismic tomography may be exposed today in older continental subduction zones, where mantle
31 wedge serpentinites are commonly associated with continental (U)HP metamorphic rocks.

32 **Keywords:** continental subduction; ultra-high-pressure metamorphism; mantle wedge exhumation;
33 peridotite serpentinization; local earthquake tomography; Western Alps

34 **Highlights:**

- 35 - First geophysical evidence of mantle wedge exhumation during continental subduction
- 36 - High-resolution image of the seismic velocity structure of the Alpine mantle wedge
- 37 - Mantle wedge exhumation is coeval to continental UHP rock exhumation

38 **1. Introduction**

39 Exhumed ultra-high pressure (UHP) units bear compelling evidence of the interaction between
40 subducting plates and the overlying mantle wedge (Carswell and Compagnoni, 2003; Hacker et al.,
41 2006; Ferrando et al., 2009; Scambelluri et al., 2010; Deschamps et al., 2013). However, the role
42 played by the mantle wedge during UHP rock exhumation is still poorly understood. Some
43 numerical models point to a negligible mantle involvement during exhumation (Yamato et al.,
44 2008; Butler et al., 2013), whereas other models suggest that the mantle may be strongly involved
45 and may possibly follow the exhumation path of buoyant UHP rocks towards Earth's surface
46 (Schwartz et al., 2001; Petersen and Buck, 2015). A precise observational characterization of the
47 behaviour of the mantle wedge in (U)HP orogenic belts such as the European Alps, where the slab
48 structure is still largely preserved (Zhao et al., 2016a), may thus provide important pinpoints for a
49 better understanding of the mechanisms leading to the exhumation of UHP continental and mantle-

50 wedge rocks more generally (Agard et al. 2009; Guillot et al., 2009; Liou et al., 2009; Little et al.,
51 2011; Malusà et al., 2011; Warren, 2013).

52 Here, we shed light on this issue by a detailed analysis of the seismic velocity structure of the
53 mantle wedge beneath the Dora-Maira (U)HP dome in the Western Alps, where coesite, attesting
54 deep continental subduction (e.g., Gilotti, 2013), was first described three decades ago (Chopin,
55 1984). We exploit a comprehensive seismic dataset, also including anomalously deep earthquakes
56 (Eva et al., 2015), to derive a local earthquake tomography model of the analyzed mantle region,
57 which is then compared with the results provided by receiver function analysis along the CIFALPS
58 transect (China-Italy-France Alps seismic survey; Zhao et al., 2015). Our results suggest that part of
59 the mantle wedge that was metasomatized above the Alpine subduction zone, was subsequently
60 exhumed at shallow depth beneath continental (U)HP rocks now exposed at the surface. These
61 findings suggest that mantle wedge exhumation can be a prominent feature of the deep structure of
62 many orogenic belts, which should be integrated in future theoretical models of continental
63 subduction and UHP rock exhumation. Moreover, our results provide new interpretive keys to
64 understand the field relationships between mantle-wedge rocks and continental UHP rocks in older
65 orogenic belts, where deep crustal levels, here illuminated by geophysical investigations, are
66 possibly exposed by protracted erosional unroofing.

67 **2. Tectonic setting**

68 The Western Alps are the result of oblique subduction of the Alpine Tethys under the Adriatic
69 microplate since the Late Cretaceous, followed by continental collision between the Adriatic and
70 European paleomargins during the Cenozoic (Coward and Dietrich, 1989; Lardeaux et al., 2006;
71 Handy et al. 2010; Malusà et al., 2015). The orogenic wedge exposed along the CIFALPS transect
72 (X-X' in Fig. 1), in the southern Western Alps, mainly consists of rocks derived from the Piedmont
73 ocean-continent transition and from the adjoining European paleomargin (Lemoine et al., 1986;
74 Dumont et al., 2012). The external zone, exposed to the west of the Frontal Pennine Fault (FPF in

75 Fig. 1), includes the Pelvoux and Argentera basements and their deformed Meso-Cenozoic
76 sedimentary cover sequences (Ford et al., 2006), which record a transition from thin-skinned to
77 thick-skinned compressional tectonics during the Neogene (Schwartz et al., 2017). East of the
78 Frontal Pennine Fault, in the Alpine metamorphic wedge, the Briançonnais nappe stack (Br in Fig.
79 1) mainly consists of Upper Paleozoic to Mesozoic metasediments and underlying pre-Alpine
80 basement rocks that underwent subduction starting from the Paleocene, and were later exhumed in
81 the Eocene - early Oligocene (Malusà et al., 2002, 2005; Ganne et al., 2007; Lanari et al., 2014).
82 The Briançonnais nappe stack forms the core of the present-day Alpine fan-shaped structure
83 (Michard et al., 2004) that was overprinted by a dense network of extensional faults during the
84 Neogene (Sue et al., 2007; Malusà et al., 2009). The eastern part of the fan is formed by oceanic
85 metasediments of the Schistes lustrés complex (SL in Fig. 1; Lemoine et al., 1986; Lagabrielle and
86 Cannat, 1990), including boudinaged decametre-to-kilometre-sized ophiolitic bodies that were
87 deformed and metamorphosed during Alpine subduction under blueschist to transitional blueschist-
88 eclogite facies conditions (Agard et al., 2002; Tricart and Schwartz, 2006; Schwartz et al., 2009). A
89 ductile normal fault (Ballèvre et al., 1990) separates the Schistes lustrés complex from the Viso
90 metaophiolites (Vi in Fig. 1; Lombardo et al., 1978; Angiboust et al., 2012), representing major
91 imbricated remnants of the Tethyan oceanic lithosphere that were deformed and metamorphosed
92 under eclogite facies conditions during the Eocene (Duchêne et al., 1997; Schwartz et al., 2000;
93 Rubatto and Hermann, 2003). Another ductile normal fault (Blake and Jayko, 1990) separates the
94 Viso eclogites from the underlying stack of deeply subducted continental basement slices referred
95 to as the Dora-Maira (U)HP dome (DM in Fig. 1; Henry et al., 1993; Michard et al., 1993). This
96 dome includes the coesite-bearing Brossasco-Isasca eclogitic unit (black star in Fig. 1; Chopin et
97 al., 1991; Compagnoni and Rolfo, 2003), which is sandwiched between quartz-bearing eclogite
98 facies rocks, above, and blueschist facies metasediments, below (Avigad et al., 2003). Along the
99 boundary with the Po Plain, the CIFALPS transect crosses the southern tip of the Lanzo massif (La
100 in Fig. 1; Boudier, 1978; Piccardo et al., 2007), an eclogitized mantle slice separated from the Dora-

101 Maira dome by a vertical active fault system rooted in the upper mantle (Malusà et al., 2017), at the
102 southward prolongation of the Insubric Fault. The Lanzo massif consists of slightly serpentized
103 spinel plagioclase peridotites surrounded by a 3–5 km thick envelope of foliated serpentinites
104 (Müntener et al., 2004; Debret et al., 2013), and records a high-pressure metamorphic peak of early
105 Eocene age (Rubatto et al., 2008). Beneath the Po Plain, the complex transition zone between the
106 Adriatic upper plate and the Apennines, also involving rotated fragments of the Alpine orogenic
107 wedge (Maffione et al., 2008; Eva et al. 2015), is mainly covered by thick Cenozoic to Quaternary
108 sedimentary successions.

109 **3. Methods**

110 ***3.1. Building the database***

111 The local earthquake tomography presented in this work is largely based on the dataset collected
112 during the CIFALPS experiment (Zhao et al., 2016b), which was integrated by data recorded in the
113 same time interval by permanent seismic networks operating in Italy and France, and complemented
114 with select older events. The temporary network of the CIFALPS experiment (blue marks in Figure
115 1B) includes 46 broadband seismic stations deployed along a linear WSW-ESE transect from the
116 European foreland to the western Po Plain, and 9 additional stations installed to the north and to the
117 south of the main profile. Stations operated from July 2012 to September 2013, and were
118 specifically deployed for a direct comparison between receiver function and local earthquake
119 tomography. Stations located along the main profile were conceived for receiver function analysis
120 (Zhao et al., 2015). Their spacing ranges from ~5 km in the Western Alps mountain range to ~10
121 km in the European foreland and in the western Po Plain. Off profile stations were installed to
122 improve the crossing of seismic rays for local earthquake tomography.

123 The high number of recording stations along the main CIFALPS profile may increase the
124 computational burden during local earthquake tomography (e.g. in ray tracing) without a direct
125 improvement in the final resolution. However, it ensures a number of advantages. For example, any

126 potential loss of data due to station malfunctioning is easily recovered by adjacent instruments, and
127 doubtful data can be discarded without jeopardizing the quality of the dataset. In order to improve
128 the ray coverage and ensure ray crossing from any azimuth in the study volume, we added to the
129 dataset all published phase pickings recorded by permanent seismic stations operating in France and
130 Italy during the CIFALPS experiment (red marks in Figure 1B). We additionally considered few
131 events that occurred before the experiment to fill specific spatial gaps. This was the case of the
132 intermediate depth earthquakes that were useful to sample anomalies at the bottom of the study
133 volume. Because these earthquakes are relatively rare (Eva et al., 2015), only few events were
134 recorded during the CIFALPS experiment. In summary, 270 events on a total of 1088 events
135 utilized in this work were added as supplementary entries from datasets available at French and
136 Italian seismic networks; about 80% of the remaining events were merged with existing phase
137 pickings. The final P and S ray coverage is shown in Figure 2A.

138 ***3.2. Seismic tomography setup and reconstruction test***

139 We adopted the local earthquake tomography code SIMULPS (Thurber, 1983) for tomographic
140 analysis, in its version 14 that implements the ray tracer by Virieux (1991) to cope with models of
141 regional size. We subdivided the study volume into layers containing nodes, and used an initial
142 velocity model derived from previous seismic experiments over a larger area (Scafidi et al., 2009).
143 Several tests were performed for a correct choice of the inversion parameters, and classical damping
144 trade-off curves (Eberhart-Phillips, 1986) were computed to pick up the best values for P and S
145 velocities.

146 The resolution capability of the coupling between inversion setup and data was evaluated by
147 checkerboard and reconstruction tests. These tests were useful to choose an adequate geometry of
148 the starting model and evaluate the smearing due to the contrast between high and low velocity
149 anomalies. The reconstruction test was specifically conceived to test the potential impact of the
150 high-velocity Ivrea body, a long recognized tectonic feature associated to a positive gravimetric

151 anomaly (red dotted line in Fig. 1) and interpreted as a slice of Adriatic mantle emplaced at shallow
152 depth (Closs and Labrouste, 1963; Nicolas et al., 1990). We used a “stairwell” geometry to simulate
153 a high-velocity east-dipping layer along the CIFALPS profile (Fig. 2B) and test the resolution
154 capability of the coupling between seismic dataset and inversion setup. The same geometry after
155 interpolation by the algorithm used in SIMULPS is shown in Figure 2C. A comparison with Figure
156 2B shows that the interpolation process introduces a smoothing of the anomalies and a band of fake
157 colors around them. Figure 2D shows the reconstruction of the imposed stairwell structure based on
158 our seismic dataset. The inversion of synthetic data does not consider the resolution, and Figure 2D
159 only displays the reconstructed model as if it was completely resolved except for areas that were not
160 sampled (in white). As shown in the reconstruction test, the shape of the anomaly is well
161 reproduced, but the velocity of the first and second steps is lowered from ~ 8.0 km/s (blueish) to
162 about ~ 7.5 km/s (greenish), and weak vertical and horizontal periodic stripes of yellow color appear
163 at ~ 50 km depth. These artifacts, and the undestimation of the magnitude of the high velocity
164 anomalies in the uppermost 10 km of the crust, have been considered during the subsequent phases
165 of tomography interpretation. The real data tomographic model is about 700×700 km wide, and was
166 obtained after 6 iterations on a 12 layers model of 36×36 nodes each. In the central part of the
167 model, spacing between nodes is equal to 15 km.

168 **4. Results**

169 Figure 3 shows the V_p and V_p/V_s cross-sections along the CIFALPS profile. The lighter areas
170 are those where the diagonal elements of the resolution matrix are < 0.1 . This threshold was chosen
171 as the divider between resolved and non-resolved areas based on a comprehensive comparisons
172 between different resolution indicators (Paul et al., 2001). As expected, the maximum depth of the
173 resolved area is limited by the depth of occurrence of most of the deepest events (Eva et al., 2015;
174 Malusà et al., 2017). Beneath the Dora-Maira (U)HP dome, the tomography model is well resolved
175 down to 50-60 km depth, whereas the two extremes of the CIFALPS cross section are poorly

176 resolved. Letters “a” to “k” indicate the relevant velocity features highlighted by the tomography
177 model. The main tectonic structures previously inferred from receiver function analysis (Zhao et al.
178 2015) and surface geology (Lardeaux et al., 2006; Malusà et al., 2015) are also indicated for
179 comparison (black lines in Fig. 3).

180 The most prominent feature of the tomography model is represented by the high velocity body
181 ($V_p \sim 7.5$ km/s; $V_p/V_s = 1.70$ - 1.72), labelled with “a”, which is located right below the Dora-Maira
182 (U)HP dome, at depths as shallow as ~ 10 km. Such a high-velocity body was already imaged with
183 similar velocities by previous works ($V_p \sim 7.4$ - 7.7 km/s; Paul et al., 2001; Béthoux et al., 2007), but
184 was only resolved down to depths of 15-20 km. It is still observed to the south of the CIFALPS
185 profile (Fig. 4D,E), but progressively vanishing towards the north (Fig. 4A,B). A series of N-S
186 cross sections, ranging from the Western Alps to the Po Plain (Fig. 5), shows that this high-velocity
187 anomaly is exclusively found beneath the Dora-Maira (U)HP dome (Fig. 5A), and disappears
188 farther east.

189 The mantle-wedge region labelled with “b” is located at depth of 20-45 km, in correspondence
190 with a cluster of intermediate depth earthquakes and with a seismically active fault in the mantle
191 described in previous studies (Rivoli-Marene deep fault; Eva et al., 2015; Malusà et al., 2017). This
192 region shows higher V_p values (~ 8.0 km/s) compared to region “a”, and anomalously high V_p/V_s
193 ratios (> 1.74) that are supportive of low shear wave velocities. This cluster of intermediate depth
194 earthquakes in region “b” is not only observed along the CIFALPS profile, but also in cross sections
195 located more to the north or to the south (Fig. 4). The deepest mantle wedge region resolved by the
196 tomographic model is labelled with “c”. This region, located at depth of ~ 40 - 50 km atop the
197 European slab, shows lower V_p and V_p/V_s values compared to region “b” ($V_p \sim 7.0$ - 7.5 km/s;
198 $V_p/V_s < 1.70$), but the V_p/V_s ratio is locally higher ($V_p/V_s \sim 1.74$).

199 The well-resolved regions of the model also include some subducted European lower crust. This
200 shows a progressive increase in V_p from the region labelled with “d” ($V_p \sim 6.7$ km/s) to the region
201 labelled with “e” ($V_p \sim 7.7$ km/s), under a rather constant V_p/V_s ratio of 1.70-1.72. Such variations

202 are detected in all of the analyzed WSW-ENE transects of Figure 4. No seismic event was recorded
203 in regions “d” and “e” since 1990 (installation of permanent seismic networks) and during the
204 CIFALPS experiment (Eva et al., 2015; Malusà et al., 2017).

205 On the eastern side of the transect, the region labelled with “f” is located below the Adriatic
206 Moho as determined by receiver function analysis combined with gravity modelling. It shows V_p
207 values ~ 8.0 km/s and $V_p/V_s = 1.70$ - 1.72 . This region is affected by intermediate depth earthquakes
208 that are also observed to the north and to the south of the CIFALPS transect (Fig. 4). The vertical
209 and horizontal periodic stripes of yellow color observed at 50 km depth in this region are artifacts,
210 as confirmed by the reconstruction test of Fig. 2D. Above the Adriatic Moho, measured V_p values
211 are much lower, but very high V_p/V_s values (>1.8) are locally observed at ~ 30 km depth at the base
212 of the Adriatic crust. This region, labelled with “g”, is also characterized by a cluster of seismic
213 events that are only observed in the vicinity of the main CIFALPS transect.

214 In the uppermost part of the Alpine orogenic wedge (regions “h” to “k”), V_p values are
215 invariably <6.5 km/s, but major variations in V_p/V_s ratios are locally observed. For example, the
216 region to the east of the Dora-Maira (U)HP dome (labelled with “h”) shows V_p/V_s values >1.72 ,
217 whereas the region corresponding to the western flank of the Dora-Maira dome (labelled with “j”)
218 shows much lower V_p/V_s ratios, even <1.66 . V_p/V_s ratios <1.68 are also observed in the region
219 labelled with “k”, located beneath the Frontal Pennine Fault. The double-vergence accretionary
220 wedge located to the east of the Frontal Pennine Fault, and labelled with “i”, shows instead V_p/V_s
221 values > 1.75 , and includes most of the shallow earthquakes recorded in the Western Alps area.

222 **5. Comparison with receiver function analysis**

223 Results of local earthquake tomography are compared in Figure 6 with published CIFALPS
224 results of receiver function analysis (Zhao et al., 2015). Unlike local earthquake tomography, the
225 receiver function technique is based on the analysis of teleseismic earthquakes, and enhances P-to-S
226 (Ps)-converted waves on velocity interfaces beneath an array. The polarity of the converted signal

227 depends on the sign of the velocity change, and interfaces with velocity increase can be
228 discriminated from interfaces with velocity decrease. Assumptions and arbitrary choices of the
229 receiver function approach applied to the CIFALPS transect (e.g., magnitude threshold, epicentral
230 distance, seismograms filtering, velocity model, choice of the direction of back azimuths) are
231 described in full in Zhao et al. (2015).

232 The image of Figure 6B is based on radial receiver functions from teleseismic events with
233 magnitude ≥ 5.5 , epicentral distance of 30-90°, and ENE back-azimuths (see Zhao et al., 2015). This
234 image shows two major interfaces marked by positive-polarity Ps-conversions (red-to-yellow
235 regions), which attest the downward velocity increase corresponding to the European and Adriatic
236 Mohos (thick dashed lines). The eastward-dipping European Moho is recognized from ~40 km
237 depth beneath the Frontal Pennine Fault to ~75 km depth beneath the Po Plain. The Adriatic Moho
238 is recognized from 20-30 km depth, to the east, to 10-15 km depth, to the west. The red spots
239 located at 40-55 km depth beneath the Adriatic Moho are multiples, as confirmed by synthetic tests
240 (Zhao et al., 2015). A shallow positive-polarity converted phase is also observed beneath the Dora-
241 Maira massif, between regions “a” and “h”, whereas a spot of negative-polarity Ps-conversions
242 marking a downward velocity decrease is located above region “c”, at 20-40 km depth (blue
243 region).

244 On the eastern side of the CIFALPS transect, the sharp velocity increase from $V_p < 6.5$ km/s to
245 $V_p > 8$ km/s evidenced by local earthquake tomography faithfully matches the location of the
246 downward velocity increase highlighted by receiver function analysis. Localized anomalies in
247 V_p/V_s ratios, e.g., in region “g”, match with major breaks in the alignment of positive-polarity Ps-
248 conversions. Beneath the Dora-Maira (U)HP dome, the downward increase in V_p values from
249 region “h” ($V_p < 6.5$ km/s) to region “a” ($V_p \sim 7.5$ km/s) is consistent with the observed positive-
250 polarity Ps-conversions, whereas the downward velocity decrease from regions “a” and “b” (V_p
251 ~ 7.5 km/s and > 8 km/s) to region “c” ($V_p \sim 7.0-7.5$ km/s) is consistent with the spot of negative-
252 polarity Ps-conversions located at 20-40 km depth in Figure 6B. The shape of the high-velocity

253 region labelled with “a” is also mirrored by the distribution of seismic events recorded since 1990.
254 Region “a” is virtually aseismic (Malusà et al., 2017), and earthquakes are chiefly located along its
255 external boundaries or in the surrounding regions (Fig. 6B). On the western side of the CIFALPS
256 transect, the alignment of positive-polarity Ps-conversions generated along the European Moho is
257 partly included within the resolved area of the local earthquake tomography model, and fits with a
258 downward velocity increase from ~6.7 km/s (region “d”) to >8 km/s. The velocity structure
259 unravelled by the analysis of local earthquakes is thus independently confirmed by the analysis of
260 teleseismic earthquakes (Zhao et al., 2015) and by the distribution of seismic events (Eva et al.,
261 2015; Malusà et al., 2017).

262 **6. Geologic interpretation**

263 The geologic cross section of Figure 6C shows the main features of the orogenic wedge of the
264 Western Alps, and of the mantle wedge between the European and the Adriatic plates as inferred
265 from the velocity structure derived from local earthquake tomography along the CIFALPS profile.
266 In the European plate, the Vp increase evidenced at ~40 km depth by local earthquake tomography,
267 from <6.5 km/s in region “d” to >8 km/s in region “e”, is consistent with a progressive
268 eclogitization of the lower crust and consequent density increase by metamorphic phase changes
269 (e.g., Hacker et al., 2003). This interpretation also explains the progressive weakening of the
270 positive-polarity converted phases observed along the European Moho (from red to yellow
271 background colours in Fig. 6B) previously described by Zhao et al. (2015). In more detail, the
272 velocity of the European lower crust is symptomatic of an intermediate granulitic composition in its
273 western part (e.g. Weiss et al., 1999; Wang et al., 2005), and of a progressive increase in garnet
274 content to the east during transition from granulite to eclogite (Christensen, 1989; Hacker et al.,
275 2003). Observed Vp values in this region are far too low for a pure mafic eclogite (Bezacier et al.
276 2010; Reynard, 2013), thus suggesting no compositional change from west to east in the European
277 lower crust, but only an increase in metamorphic grade. This interpretation is consistent with the

278 progressive increase in V_s values from west to east described along the same transect (Lyu et al.,
279 2017).

280 On the eastern side of the Western Alps, V_p values >8 km/s confirm the presence of Adriatic
281 mantle at shallow depth beneath the western Po Plain (10-15 km), just in correspondence with the
282 positive gravimetric anomaly classically referred to as the Ivrea body (Closs and Labrouste, 1963;
283 Nicolas et al., 1990) and in line with results of previous tomographic models (e.g., Solarino et al.
284 1997; Paul et al., 2001; Scafidi et al., 2006; 2009; Diehl et al., 2009; Wagner et al., 2012). East of
285 the Ivrea body gravimetric anomaly, the Adriatic Moho is located at 30-35 km depth, which is a
286 much more reliable estimate of the Moho depth beneath the Po Plain compared to previous
287 estimates based on receiver function alone (Zhao et al., 2015). The locally high V_p/V_s ratios >1.8 ,
288 associated to V_p of 7.0-7.5 km/s (region “g”), may be supportive of gabbro (Weiss et al., 1999)
289 underplated at the base of the Adriatic lower crust. Noteworthy, Permian gabbros are indeed
290 exposed north of the Po plain, where they are intruded into lower crust rocks belonging to the
291 Adriatic (Southalpine) basement (Quick et al., 1994; Schaltegger and Brack, 2007).

292 In the uppermost part of the Alpine wedge, the structural variability of stacked rocks is largely
293 mirrored by their variability in V_p/V_s ratios. The V_p/V_s values >1.75 observed in the double-
294 vergence accretionary wedge, chiefly including Briançonnais and Schistes lustrés units (Lardeaux et
295 al., 2006; Malusà et al., 2015), may reflect low V_s values, possibly associated to the widespread
296 network of mesoscale faults developed in these rocks since the Neogene (Tricart et al., 2004; Sue et
297 al., 2007; Malusà et al., 2009). To the east, low V_p/V_s values even <1.66 observed on the western
298 flank of the Dora-Maira dome (region “j”) may instead reflect high V_s velocities, suggesting that
299 the poorly fractured granitic gneisses exposed at the surface (Brossasco granite; Paquette et al.,
300 1999; Lenze and Stöckhert, 2007) may be also present at depth. Fracturing may be also invoked to
301 explain the low V_s values observed along the eastern boundary of the Dora-Maira dome, where
302 (U)HP continental rocks are juxtaposed against the eclogitized mantle rocks of the Lanzo massif
303 (Kienast and Pognante, 1988; Piccardo et al., 2007) along the Lis-Trana deformation zone (Perrone

304 et al., 2010). To the west of the Frontal Pennine Fault, V_p/V_s values <1.68 suggest instead that the
305 European upper crust in the External zones is poorly deformed, consistent with minor seismicity
306 recorded in that area (Fig. 6B).

307 But the most relevant information provided by the tomography model presented in this article is
308 related to the velocity structure beneath the Dora-Maira (U)HP dome. Two end-member tectonic
309 reconstructions were recently proposed for this region in the light of available geophysical data
310 (Zhao et al., 2015; Malusà et al., 2017): one invoking a thick complex of (U)HP continental slivers,
311 in line with the predictions of numerical models of syn-convergent exhumation (e.g., Butler et al.,
312 2013; Jamieson and Beaumont, 2013), and another one invoking a larger volume of mantle rocks
313 exhumed at shallow depth during divergence between the upper plate and the accretionary wedge
314 (e.g., Malusà et al., 2011, 2015). Our tomographic model can be used to discriminate between these
315 end-member hypotheses.

316 The velocity structure of the mantle wedge region “a”, showing V_p velocity of ~ 7.5 km/s from
317 depths as shallow as ~ 10 km down to ~ 30 km, is largely inconsistent with the presence of dry
318 mantle peridotite beneath the Dora-Maira (U)HP continental rocks, and suggests a complex history
319 for these rocks in terms of P-T conditions and fluid-rock interaction. Such V_p values point in fact to
320 widespread serpentinization of mantle rocks ($\sim 60\%$ according to Reynard, 2013), that may locally
321 exceed 90% both in the uppermost part of anomaly “a” and in the Lanzo massif, although velocity
322 values in the uppermost crustal levels may be slightly underestimated as unravelled by the
323 reconstruction tests of Fig. 2D. The degree of serpentinization at 30-40 km depth is instead much
324 lower ($<30\%$), and consistent with the occurrence of intermediate-depth earthquakes (Fig. 6B).
325 V_p/V_s ratios are in the range of 1.70-1.72 in region “a”, but sharply increase to values >1.74 in
326 region “b”, where V_p values (~ 8.0 km/s) are consistent with dry mantle peridotite. The high V_p/V_s
327 ratios in region “b” point to low shear wave velocities, which are in line with the presence of a
328 major active fault in the upper mantle as underlined by the alignment of anomalously deep
329 earthquakes (Eva et al., 2015). According to previous work, the deepest part of the mantle wedge

330 beneath the thick blue spot of negative polarity conversions (region “c” in Fig. 6B) may either
331 include serpentinites, or slivers of (U)HP rocks. Our results indicate that the velocity values
332 observed in region “c” ($V_p \sim 7.0\text{-}7.5$ km/s; $V_p/V_s < 1.70$) are neither consistent with eclogitic
333 metasediments ($V_p \sim 7.0$ km/s; $V_p/V_s \sim 1.75$) nor with mafic eclogite ($V_p > 8.0$ $V_p/V_s \sim 1.73$), but
334 are instead supportive of ultramafic rocks with a degree of serpentinization ranging between 50%
335 and 75% (Reynard, 2013). Minor slivers of eclogitic metasediments may be present beneath the
336 mantle wedge, in regions showing the highest V_p/V_s ratios at the top of the European slab. These
337 results demonstrate that recent geologic cross-sections postulating a thick wedge of Briançonnais
338 eclogites beneath the Dora-Maira (Schmid et al., 2017) are likely incorrect, and that the palinspastic
339 reconstruction of the Alps-Apennines transition zone derived from such geologic cross-sections
340 should be reconsidered.

341 **7. Implication for (U)HP rock exhumation**

342 In the southern Western Alps, the positive gravimetric anomaly ascribed to the Ivrea body is
343 classically interpreted in terms of upper mantle indentation (e.g., Lardeaux et al., 2006; Béthoux et
344 al., 2007), in line with previous tectonic interpretations proposed for the Central Alps and for the
345 northern Western Alps (Schmid and Kissling, 2000). According to these interpretations, the
346 uppermost part of the Adriatic mantle would act as an indenter beneath the Alpine accretionary
347 wedge, and would transfer compression towards the European foreland. The main geologic
348 implications of this model include: (i) major crustal shortening in the upper plate, and (ii) fast
349 erosion focused above the indenter (Fig. 7A). These features are indeed observed in the Central
350 Alps, where upper mantle indentation accommodated by back-folding of (U)HP domes (Keller et
351 al., 2005) and by backthrusting of Adriatic units (Zanchetta et al., 2015) triggered the fast erosional
352 exhumation of the amphibolite-facies Lepontine dome (Malusà et al., 2016). However, these
353 features are not common to the southern Western Alps. Indeed, during and after Eocene (U)HP rock
354 exhumation, shortening recorded in the accretionary wedge of the southern Western Alps was minor

355 (Malusà et al., 2009; Dumont et al., 2012) and erosion was much slower compared to the Lepontine
356 dome, as attested by low-temperature thermochronometers (Vernon et al., 2008; Fox et al., 2015)
357 and by preserved Oligocene corals unconformably lying on top of Eocene eclogites (Quaranta et al.
358 2009; Malusà et al., 2015). A tectonic scenario exclusively invoking upper-plate mantle indentation
359 beneath the accretionary wedge would also imply that seismic velocities in the mantle of the upper
360 plate should be quite similar beneath the orogenic wedge and in the hinterland (Fig. 7A). Major
361 seismic velocity changes, e.g., by metamorphic phase changes triggered by fluids released by the
362 downgoing slab, would remain probably undetected in the local earthquake tomography model,
363 because they would take place at much greater depth (Deschamps et al., 2013; Abers et al., 2017).

364 Noteworthy, our study points to a complex velocity structure in the upper plate mantle of the
365 southern Western Alps. The region beneath the Dora-Maira (U)HP dome is dominated by
366 serpentinitized peridotites, documented from ~10 km depth down to the top of the European slab. To
367 the east, these rocks are juxtaposed against dry mantle peridotites of the Adriatic upper plate along a
368 steeply dipping fault in the lithospheric mantle. In between, mantle rocks of the Lanzo massif
369 underwent subduction during the Alpine orogeny, and were later exhumed and accreted against the
370 Adriatic upper plate when the Dora-Maira (U)HP rocks were still buried at mantle depths (Rubatto
371 and Hermann, 2001). This scenario may suggest (U)HP rock and mantle-wedge exhumation mainly
372 triggered by upper plate divergent motion (Fig. 7B). Serpentinitized peridotites with V_p ~7.5 km/s
373 that are found beneath the Dora-Maira dome may have favoured the exhumation of (U)HP rocks
374 across the upper crust, where these rocks become neutrally buoyant (e.g., Schwartz et al., 2001). No
375 exhumed mantle-wedge serpentinites are recognized so far at outcrop in the southern Western Alps
376 (Scambelluri et al., 1995; Piccardo et al., 2004; Hattori and Guillot, 2007; Deschamps et al., 2013).
377 However strong fluid-rock interactions are recognized in subducted serpentinites and associated
378 ophiolitic rocks (Scambelluri and Tonarini, 2012; Lafay et al., 2013; Plüumber et al., 2017),
379 suggesting that fluid release may have occurred during oceanic and even during continental
380 subduction (e.g., Castelli et al., 2007; Ferrando et al., 2009) possibly triggering the partial

381 serpentization of the Adriatic mantle wedge. Part of the Adriatic mantle wedge was then exhumed
382 at shallow crustal levels during late Eocene transtension along the Western Alps subduction zone
383 (Malusà et al. 2015), and was finally indented beneath the Alpine wedge during the early Oligocene
384 tectonic shortening (Dumont et al., 2012; Jourdan et al., 2012, 2013). Along the Adria-Europe plate
385 boundary, the divergent component of Eocene transtension progressively decreased towards the
386 north to become negligible in the Central Alps, where Adria was indented more deeply beneath the
387 accretionary wedge compared to the Western Alps (Malusà et al., 2015). We speculate that, north of
388 the Dora-Maira dome, upper plate divergence was probably insufficient to allow an effective
389 exhumation of the mantle wedge. However, testing this hypothesis would require a high resolution
390 tomographic image of the northern Western Alps, which may be precluded by the lack of deep
391 earthquakes.

392 The occurrence of mantle-wedge serpentinites exhumed at shallow depth within a continental
393 subduction zone is not specific of the southern Western Alps. Mantle wedge serpentinites associated
394 with (U)HP rock are described, for example, in the Indus Suture Zone in the Himalaya, in the
395 Carribbean (Guillot et al., 2001; Deschamps et al., 2012), in the Western Gneiss Region in Norway
396 (Scambelluri et al., 2010), and are inferred by geophysical evidence under the Dabie-Sulu (Liu et
397 al., 2015). Our findings suggest that orogen-scale exhumation of the mantle wedge may represent a
398 prominent, but still underestimated feature of the deep structure of many orogenic belts. As such, it
399 should be integrated in more advanced theoretical models of subduction and exhumation. Moreover,
400 widespread mantle-wedge exhumation may explain the common occurrence of boudinaged mantle-
401 wedge rocks within continental UHP rocks in the roots of old orogenic belts now unroofed by
402 erosion. In pre-Cenozoic orogenic belts such as the Dabie-Sulu or the Western Gneiss Region,
403 where the evidence of minor erosion during UHP exhumation, if any, is no longer preserved, the
404 occurrence of mantle wedge rocks at shallow depth may represent the only evidence supporting
405 (U)HP rock exhumation triggered by divergent motion between upper plate and accretionary
406 wedge.

407 **8. Conclusions**

408 The new local earthquake tomography model of the southern Western Alps, independently
409 validated by receiver function analysis, unravels a complex seismic velocity pattern consistent with
410 a composite structure of the mantle wedge above the subducted European lithosphere. Seismic
411 velocities indicate that the Dora-Maira (U)HP dome lays directly above serpentinized peridotites,
412 documented from ~10 km depth down to the top of the eclogitized lower crust of the European
413 plate. We propose that peridotite serpentinization was the result of fluids released to the Adriatic
414 mantle wedge during Alpine subduction. The minor amount of eclogitic metasediments above the
415 European slab indicates that the subduction wedge was largely exhumed during Eocene
416 transtension, when part of the mantle wedge was also exhumed at shallow crustal levels to be
417 finally indented under the Alpine metamorphic units. The serpentinized peridotites imaged beneath
418 the Dora-Maira (U)HP dome are not exposed at Earth's surface, and are juxtaposed against dry
419 mantle peridotites of the Adriatic upper plate along an active fault rooted in the lithospheric mantle.
420 Our results suggest that mantle wedge exhumation may represent an important feature of the deep
421 structure of exhumed continental subduction zones. Deep orogenic levels as those imaged by local
422 earthquake tomography in the southern Western Alps may be exposed today in older continental
423 subduction zones, where mantle wedge serpentinites are commonly associated to continental (U)HP
424 metamorphic rocks.

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429 The CIFALPS seismic data are archived at the data center of the Seismic Array Laboratory, Institute of Geology and
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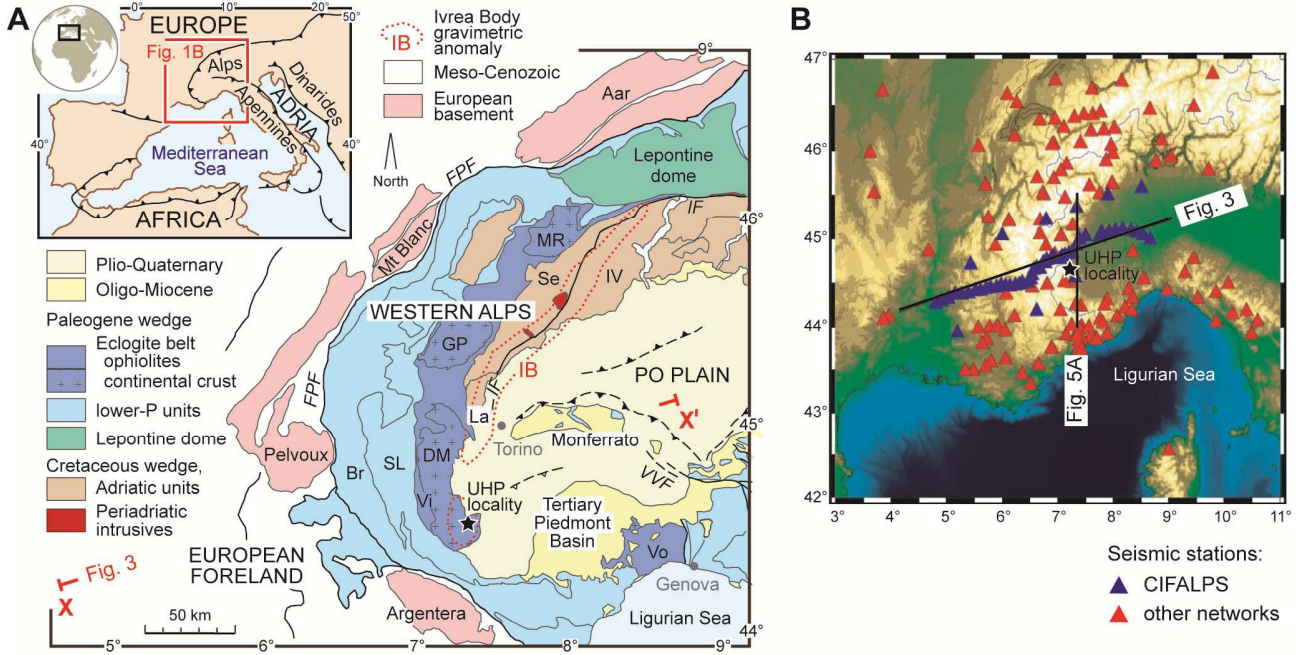
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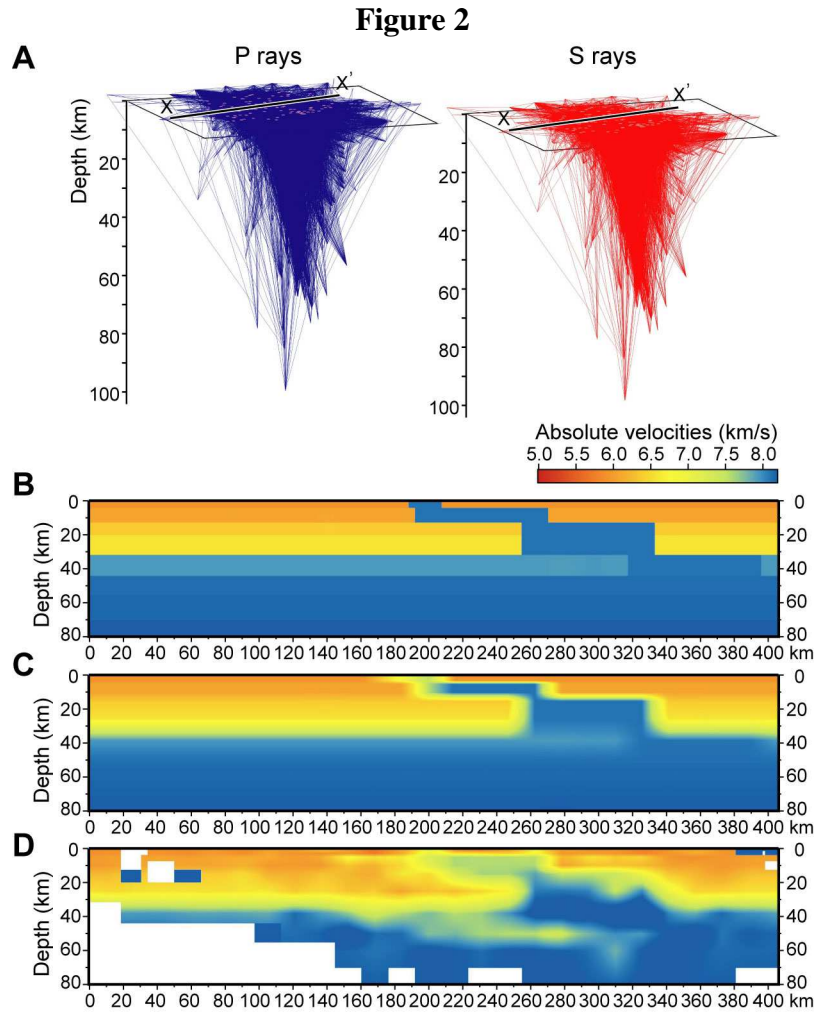
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Figure 1



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Figure 1: A) Tectonic sketch map showing the (U)HP domes of the Western Alps (dark blue), the gravimetric anomaly of the Ivrea body (0 mGal isoline in red), and the location of the CIFALPS transect (X-X'). Acronyms: Br, Briançonnais; DM, Dora-Maira; FPF, Frontal Pennine Fault; GP, Gran Paradiso; IF, Insubric Fault; IV, Ivrea-Verbano; La, Lanzo; MR, Monte Rosa; Se, Sesia-Lanzo; SL, Schistes lustrés; Vi, Viso; Vo, Voltri; VVF, Villalvernia-Varzi Fault. The black star marks the Brossasco-Isasca UHP locality. B) Seismic stations utilized in this work (blue = CIFALPS; red = other networks) and location of tomographic cross sections (black lines).



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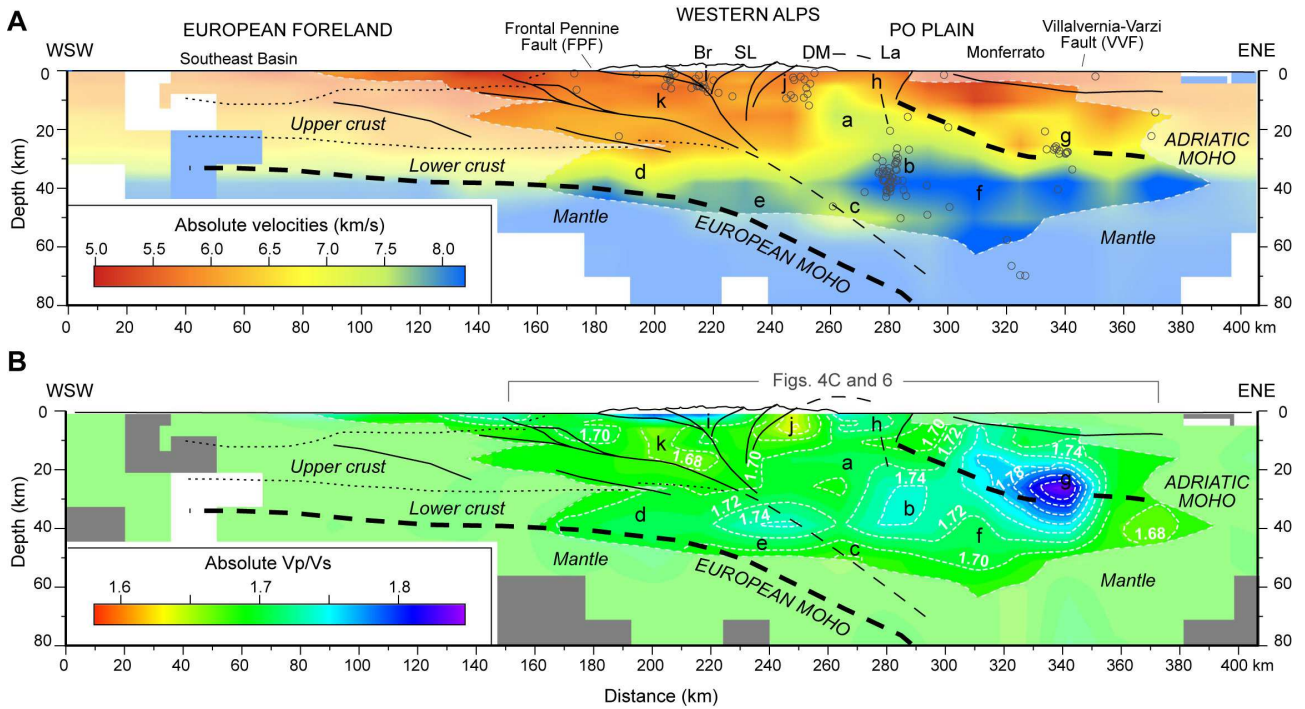
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Figure 2. A) Three-dimensional P and S ray coverage based on the seismic events considered in this study (X-X' indicates the CIFALPS transect, see Fig. 1). B) Imposed stairwell geometry along the CIFALPS transect for testing the resolution capability of the coupling between seismic dataset and inversion setup. C) Same geometry after interpolation by the algorithm used in SIMULPS, which introduces a smoothing and a thin band of fake colors around the anomalies. D) Reconstruction test showing that the shape of the imposed stairwell structure is well reproduced using our dataset, but the high velocities in the uppermost 10 km are converted to lower values (as less as 0.5 km/s); the weak vertical and horizontal periodic stripes of yellow color at 50 km depth within the blue area are artifacts; white areas are not sampled.

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Figure 3

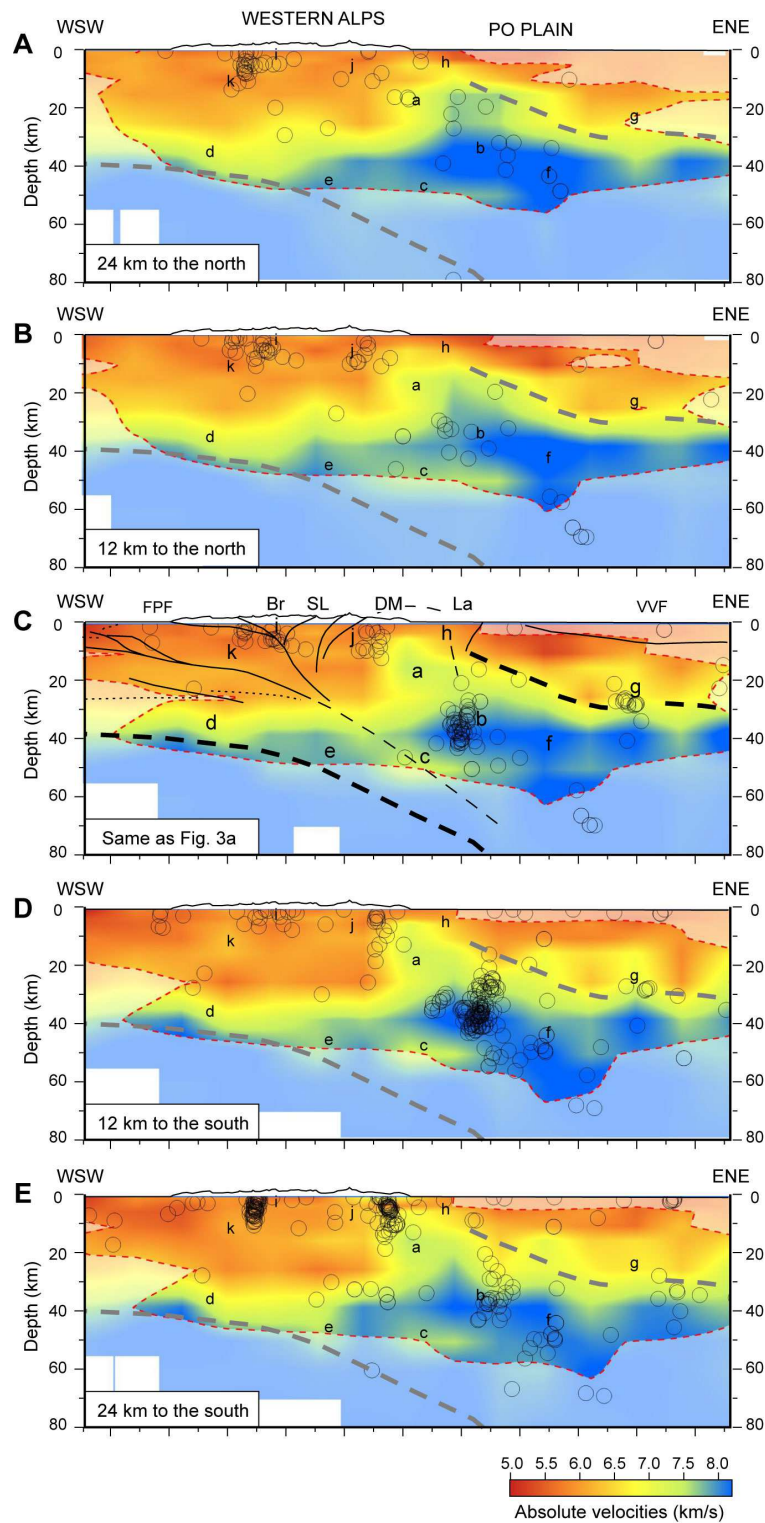


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Figure 3: Tomographic cross sections along the CIFALPS transect. A) Absolute Vp velocity. The velocity structure beneath the Dora-Maira (U)HP dome is well resolved down to 50-60 km depth (acronyms as in Fig. 1A); areas with resolution diagonal elements <0.1 are masked, white areas are not sampled; letters a to k indicate regions of the model discussed in the main text; black circles indicate earthquakes as located with the 3D model; black lines and text in italics indicate the main tectonic features previously inferred from receiver function analysis (Zhao et al., 2015; Malusà et al., 2017, see Fig. 6B). Note the prominent high velocity body (labelled with “a”) located right below the Dora-Maira (U)HP dome. The vertical and horizontal periodic stripes of yellow color at 50 km depth are artifacts, as attested by the reconstruction test of Fig. 2D. B) Vp/Vs ratios. White dashed lines are isolines of equal Vp/Vs, grey areas are not sampled (other keys as in frame A).

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Figure 4

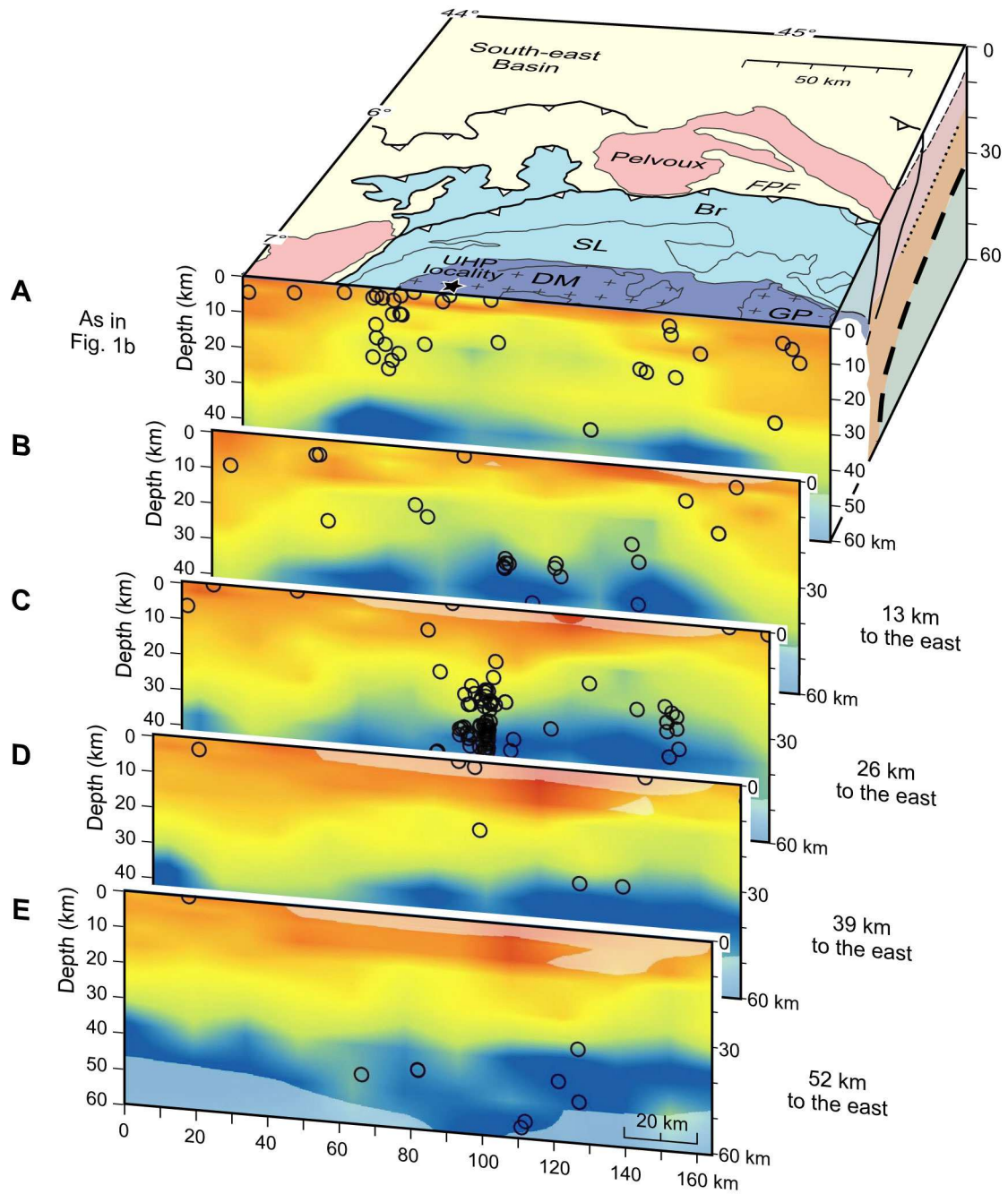


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Figure 4: Lateral variations in V_p velocity in the mantle wedge as shown in a series of WSW-ENE cross-sections lying to the north (A, B) and to the south (D, E) of the main CIFALPS transect (C). The high velocity body labelled with “a” progressively disappears moving to the north. Black circles are projected hypocentres located within ± 5 km distance off the profiles. The thick dashed lines, reported in all sections for comparison, indicate the European and Adriatic Mohos inferred from receiver function analysis (cf. Fig. 6B). Other keys as in Fig. 3.

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Figure 5



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Figure 5: Lateral variations in V_p velocity beneath the Dora-Maira (U)HP dome, as shown in a series of N-S cross-sections from the mountain range to the Po Plain. Black circles are projected hypocentres located within ± 3 km distance off the profiles. The high-velocity body labelled with “a” in Figs. 3 and 4 is exclusively found beneath the Dora-Maira dome (see cross section A) and progressively disappears towards the east. Acronyms as in Fig. 1.

Figure 6

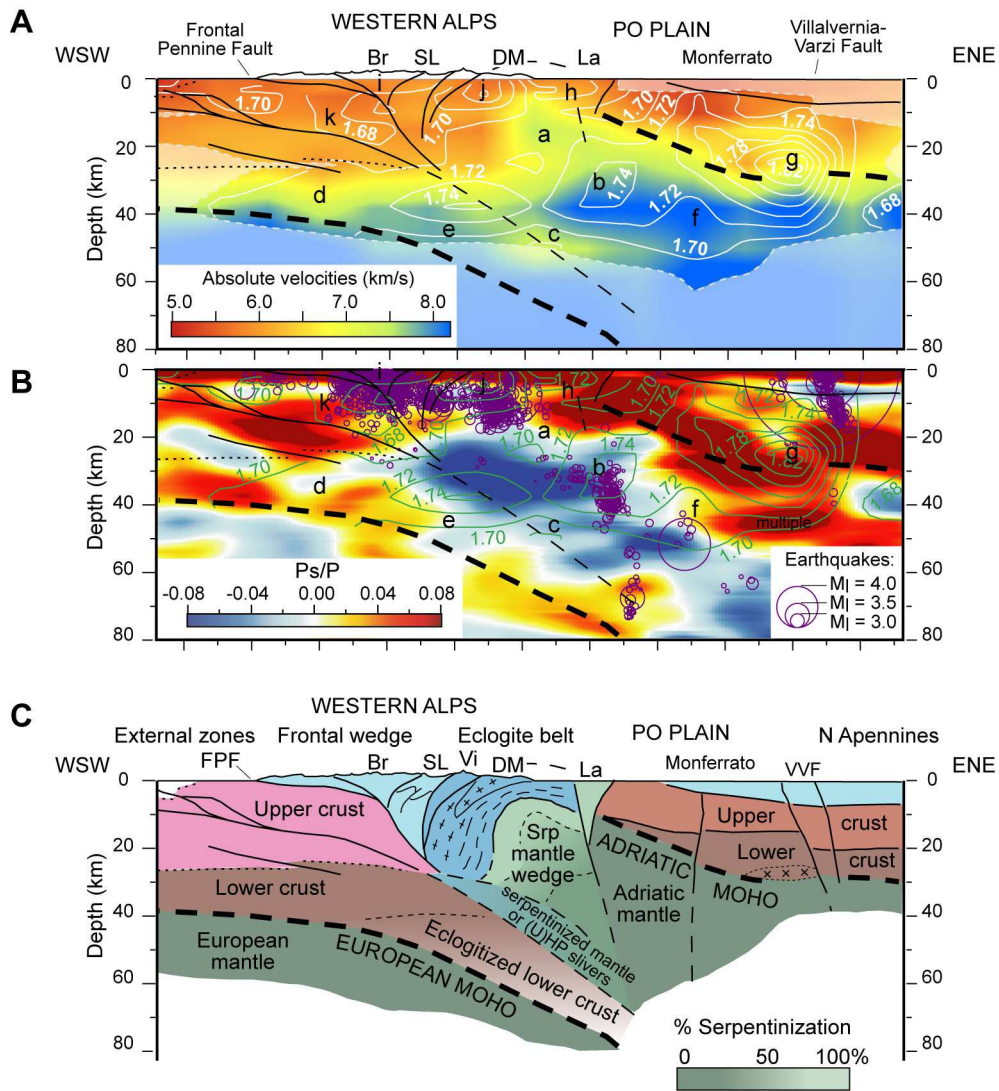
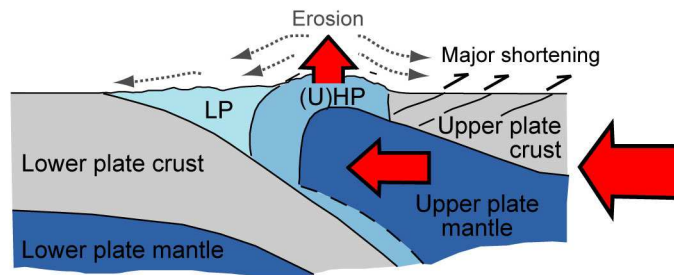


Figure 6: Synthesis of geophysical data (A, B) and inferred mantle wedge structure (C). Black lines in A and B are tectonic features based on receiver function analysis (colors in B indicate positive- and negative-polarity Ps-converted phases, Zhao et al., 2015); contours are isolines of equal V_p/V_s ; purple circles in B are earthquakes recorded since 1990 (Malusà et al., 2017). The amount of serpentinitization in C, in the mantle wedge underlying the Dora-Maira (U)HP dome, is inferred from seismic velocities (Reynard, 2013). Note the consistency between structures unravelled by local (A) and teleseismic (B) events. Acronyms as in Fig. 1, letters a to k as in Fig. 3.

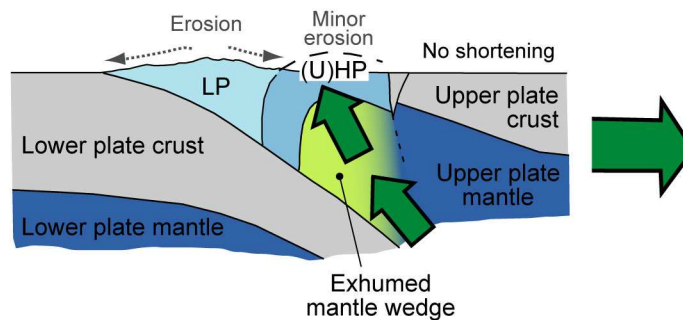
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Figure 7

A SYNCONVERGENT (U)HP ROCK EXHUMATION FOLLOWED BY UPPER-PLATE MANTLE INDENTATION



B (U)HP ROCK AND MANTLE WEDGE EXHUMATION TRIGGERED BY UPPER PLATE DIVERGENT MOTION



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Figure 7: Alternative scenarios of mantle involvement in (U)HP orogenic belts. **A)** Synconvergent exhumation of (U)HP rocks (e.g., Butler et al., 2013) is followed by indentation of the upper-plate mantle (dark blue) beneath the accretionary wedge, with consequent fast erosion of the (U)HP dome and major tectonic shortening in the upper plate (e.g., Béthoux et al., 2007). Seismic velocities in the upper-plate mantle are similar beneath the orogenic belt and in the hinterland. **B)** Divergence between upper plate and accretionary wedge triggers the exhumation of (U)HP rocks and the emplacement of serpentinitized mantle-wedge rocks at shallow depth. Erosion on top of the (U)HP dome is minor at this stage, shortening is negligible. Because of widespread serpentinitization of the mantle wedge during subduction, seismic velocities will be lower in the mantle-wedge rocks beneath the (U)HP dome, and higher in the adjoining dry mantle rocks of the upper plate.