

# 1 **MrLavaLoba: a new probabilistic model for the simulation of lava** 2 **flows as a *settling* process**

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## 6 **Abstract**

7 A new code to simulate lava flow spread, MrLavaLoba, is presented. In the  
8 code, erupted lava is itemized in parcels having an elliptical shape and  
9 prescribed volume. New parcels bud from existing ones according to a  
10 probabilistic law influenced by the local steepest slope direction and by tunable  
11 input settings. MrLavaLoba must be accounted among the probabilistic codes  
12 for the simulation of lava flows, because it is not intended to mimic the actual  
13 process of flowing, but rather to guess the most probable inundated area and  
14 thickness. The code does not provide directly the progression with time of a  
15 flow field. The code's flexibility allows it to produce variable lava flow spread  
16 and emplacement according to different dynamics (e.g. pahoehoe or  
17 channelized-'a'ā). The code does not directly provide the progression of a flow  
18 field with time. For a given scenario, it is shown that model outputs converge,  
19 in probabilistic terms, towards a single solution. The code is finally applied and  
20 compared to real cases in Hawaii and Mt Etna, and the obtained maps are  
21 shown.

22 The model is written in Python and the source code is available at  
23 <http://demichie.github.io/MrLavaLoba/>.

## 24 25 26 **1. Introduction**

27  
28 Forecasting lava flow emplacement is a critical issue at many basaltic volcanoes  
29 where the propagation of flowing lavas threaten surrounding properties and  
30 infrastructures. Lava flow simulation is becoming an important tool for hazard  
31 and risk assessment before and during ongoing eruptions (Favalli et al., 2009a,  
32 2009b; Cappello et al., 2016; Cordonnier et al., 2016; Richter et al., 2016;  
33 Dietterich et al., 2017).

34 Existing numerical codes for the simulation of lava flow emplacement are  
35 either so-called deterministic codes based on the solution of some

36 simplification of the physical governing equations of this phenomenon ([Hidaka](#)  
37 [et al. 2005](#); [Del Negro et al. 2008](#); [Crisci et al. 2010](#)), or, instead, so-called  
38 *probabilistic codes* based on the simple evidence that lava flows tend to follow  
39 the steepest descent path downhill ([Macedonio et al. 1990](#), [Felpeto et al. 2001](#);  
40 [Favalli et al. 2005](#)). Deterministic codes attempt to mimic, with a sufficiently  
41 accurate approximation, the thermo- and/or fluid-dynamics of the natural  
42 system, so that the emplacement of the erupted volume over time can be  
43 derived as a single solution if input parameters are set to the correct values.  
44 Probabilistic codes instead give more emphasis to the uncertain character of  
45 the solution ([Gilbert and Sparks 1998](#)) because the parameters and the  
46 processes under examination are known only imprecisely (e.g. with respect to  
47 the effusion rate, [Harris et al. 2007](#)).

48 Deterministic simulations, in principle, provide a final coverage and  
49 thickness of lava flow field as well as detailed information of the lava flow  
50 emplacement over time, including the velocity of advancement of the front and  
51 the overall shape and thickness of the final flow field ([Vicari et al. 2011](#)). The  
52 output of a probabilistic simulation typically provides the most likely path of the  
53 flow and the associated inundated area, without constrain on the thickness of  
54 the final flow field nor temporal insight into the planimetric evolution of the  
55 lava flow ([Tarquini and Favalli 2011, 2013, 2016](#); [Mossoux et al., 2016](#)).

56 [Glaze and Baloga \(2013\)](#) recently developed and [Hamilton et al. \(2013\)](#)  
57 tested a new code which was specifically focused on the simulation of late-  
58 stage pahoehoe lobes where random processes dominate emplacement ([Baloga](#)  
59 [and Glaze 2003](#); [Glaze and Baloga 2013](#)). This code can be considered  
60 probabilistic because the approach for selecting the transfer direction (North,  
61 South, East or West) is based on the generation of a random number. In  
62 contrast with earlier probabilistic codes, this code can provide as output the  
63 volume of the emplaced lava over time (i.e. the supply rate). This approach  
64 utilizes erupted volumes, one of the most critical input parameters considered  
65 by deterministic codes, within the stochastic approach of probabilistic codes.  
66 This allows for properties of the lava field, such as final thickness, to be  
67 addressed through a probabilistic approach.

68 In this paper, a new probabilistic code to simulate lava flows is presented.  
69 This code, named *MrLavaLoba*, rests on an approach similar to the one  
70 introduced by [Glaze and Baloga \(2013\)](#). Substantial differences are introduced,  
71 such as taking into account the topography and removing the limitation of only  
72 four transport directions allowed. We tested MrLavaLoba for the simulation of  
73 the emplacement of pahoehoe and channel-fed ‘a’ā lava flows, and we show  
74 that, through the tuning of input parameters, the model can provide solutions  
75 for coverage area and thickness.

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## 78 **2. Growth and thickening of basaltic lava flows**

79

80 The emplacement of basaltic lava flows can occur according to  
81 substantially different dynamics, ranging from pahoehoe lobes, which can later  
82 evolve into very thick sheet flows through progressive inflation ([Hon et al.,  
83 1994](#); [Self et al., 1998](#); [Hoblitt et al., 2012](#)), to the massive advancement of ‘a’ā  
84 or blocky flows ([Kilburn 2000](#)), which usually progress as large channelized flows  
85 ([Guest et al., 1987](#); [Lipman and Banks 1987](#); [Harris et al., 2002](#); [Tarquini et al.,  
86 2012a](#); [Wadge et al., 2012](#); [Wantim et al., 2013](#)).

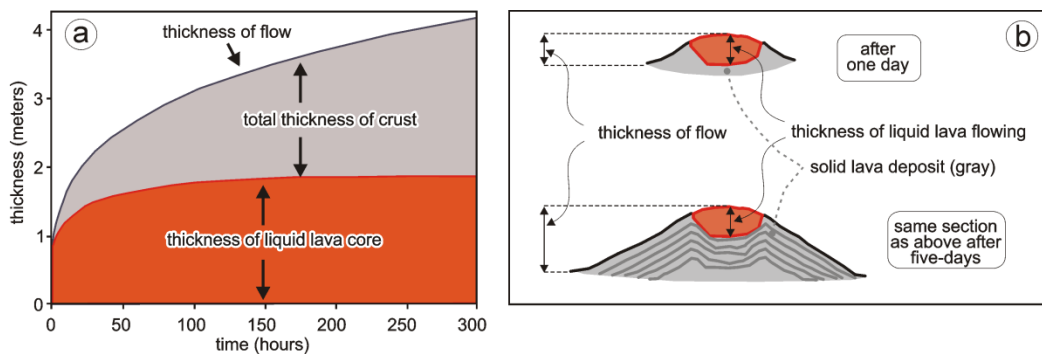
87 Although they show different dynamics and growth mechanisms, both  
88 small pahoehoe lobes and large channelized flows tend to create a thickening  
89 deposit over time ([Tarquini 2017](#)), which is as to say that, after being  
90 transported for a given time, lava parcels become part of a static deposit. As a  
91 rule of thumb, single lobes thicken as a result of injection of lava beneath a  
92 thickening crust, i.e. through an endogenous growth ([Jurado-Chichay and  
93 Rowland 1995](#); [Kauahikaua et al., 1998](#); [Thordarson and Self 1998](#); [Walker 1991,  
94 2009](#)). In contrast, individual open channels usually thicken through an iterative  
95 stacking process, i.e. according to a surface-fed growth mechanism ([Sparks et  
96 al. 1976](#); [James et al. 2007, 2012](#); [Glaze et al. 2009](#); [Applegarth et al. 2010a](#);  
97 [Tarquini and de’ Michieli Vitturi 2014](#)).

98 The thickening of individual flow units (*sensu* [Walker 1971](#)) is not the only  
99 mechanism which increases the overall lava deposit thickness. The formation of

100 multiple flow units also influences the overall thickness (Guest et al. 1987;  
 101 Kilburn and Lopes 1991). New flow units can substitute or accompany former  
 102 ones, flowing at their sides or on top of them, producing a compound flow field  
 103 (Walker 1971, 1973; Wadge 1978; Kilburn and Lopes 1988; Applegarth et al.,  
 104 2010b). The typical shorter runout of the active flow units during the waning  
 105 phase of basaltic eruptions (due to a decreasing supply, Wadge 1981), further  
 106 promotes the piling up of late flow units over the earlier ones (Guest et al.  
 107 1987; Applegarth et al. 2010a; James et al., 2009; Tarquini 2017).

108 In summary, although spreading and lengthening are also very important,  
 109 the emplacement of basaltic lava flows is always characterized by a progressive  
 110 thickening. As a rule, an increasing thickness of the lava flow field means an  
 111 increasing thickness of the solid portion (essentially a static lava deposit - Figure  
 112 1), while the amount of liquid, hot lava transported downhill tends to decrease  
 113 with time as the lava supply decreases (Tarquini 2017).

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118 **Figure 1.** Thickening of basaltic lava flows through time. (a) Simplified  
 119 progression of the inflation process of a sheet flow in Hawaii (13 April 1990)  
 120 based on measurements and modeling presented by Hon et al. (1994). (b)  
 121 Transect section showing the progressive stacking of lava blankets incrising the  
 122 total height of a lava channel measured at Mt Etna (Italy) from a Lidar survey  
 123 carried out during the early phase of the 2004-2005 eruption (modified after  
 124 Tarquini and de' Michieli Vitturi 2014). Static lava deposit is gray in figure,  
 125 molten lava is orange.

126

127 As a result, the progressive growth of a lava flow can be viewed as a  
128 process through which the supplied lava is settled to a given, permanent  
129 position located downhill from the vent, and the transport process (the actively  
130 flowing lava), determines the path along which lava parcels are settled ([Harris  
131 and Rowland 2001](#); [Wright et al., 2008](#)).

132 Parcels are not settled only at the flow front ([Borgia et al. 1983](#); [Young  
133 and Wadge 1990](#); [Castruccio et al., 2014](#)), but also behind it, where batches of  
134 lava are settled along extensive sectors of the generic lava flow unit ([James et  
135 al. 2007](#); [Glaze and Baloga 2006](#); [Baloga and Glaze 2008](#); [Glaze et al., 2009](#);  
136 [Favalli et al., 2010](#)). Thus, the lava transport process is closely intermingled with  
137 lava settling. This concept is the basis of MrLavaLoba.

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### 140 **3. The code**

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142 The main inputs of the code are topography and erupted volume. There are no  
143 special requirements as for the resolution of the topography to be used as  
144 computational domain, the code was successfully tested using DEMs ranging  
145 from 5 to 30 m in resolution. The lava supplied from the vent is itemized in  
146 parcels (lobes). The code combines in a single probabilistic procedure two  
147 elements: (1) the path followed downhill by the flowing lava, and (2) the  
148 settling of lava parcels along the path. A spectrum of tunable parameters  
149 determines the progression downhill of the flow (i.e. the extension of the flow  
150 path), which is obtained by iteratively positioning new lava parcels in their final,  
151 permanent position. In practice, the distribution of lava parcels over the pre-  
152 emplacement topography is obtained without tackling directly the complexity  
153 of the transport dynamics but through much simpler probabilistic rules: one  
154 giving the direction of propagation of new lobes, and one determining where  
155 new lobes bud.

156 The parameters determining the probability distributions of both rules  
157 have to be tuned to a given scenario through comparison with selected real lava  
158 flows.

159

### 160 3.1 The geometry of lava parcels

161 Lava parcels are represented by elliptical lobes having a constant area and  
 162 a prescribed thickness (i.e. the volume is fixed). Parcel thickness can be of the  
 163 order of centimeters down to millimeters. This low thickness of individual  
 164 parcels is, on its own, not a realistic representation, but the individual parcels  
 165 pile up to produce realistic lava flow thicknesses. If the volume of erupted lava  
 166 and the area of a lobe are held constant, as computational lobes become  
 167 thinner, the number of lobes becomes greater. The greater the number of  
 168 lobes, the closer the spreading of the lava to the expectation associated with  
 169 the probabilistic elements of the code. For example, considering a flat  
 170 topography, all directions of lava flows have the same probability, so we expect  
 171 the flow to spread radially. If the entire erupted volume is assigned to a single  
 172 parcel, only one direction will be represented in the output. In contrast, if the  
 173 volume is partitioned into a sufficiently large number of thin parcels, the  
 174 spreading will approach a uniform circle with the vents at its center.

175 The geometry of a new elliptical parcel  $l_i$  (the lobe) is defined by the  
 176 eccentricity  $e_i$  and by the azimuthal orientation of the major axis of the ellipse  
 177  $az_i''$ . Lobe position, instead, is defined by the geographic coordinates  $(x_i, y_i)$  of  
 178 the center of the ellipse ( $c_i$ ) representing the parcel (Figure 2). With the  
 179 notation just introduced, the equation of the elliptical lobe write as

$$180 \frac{[(x-x_i)\cos(az_i'')+(y-y_i)\sin(az_i'')]^2}{a_i^2} + \frac{[(x-x_i)\sin(az_i'')-(y-y_i)\cos(az_i'')]^2}{b_i^2} = 1$$

181 with the semimajor  $a$  and semiminor axis  $b$  satisfying

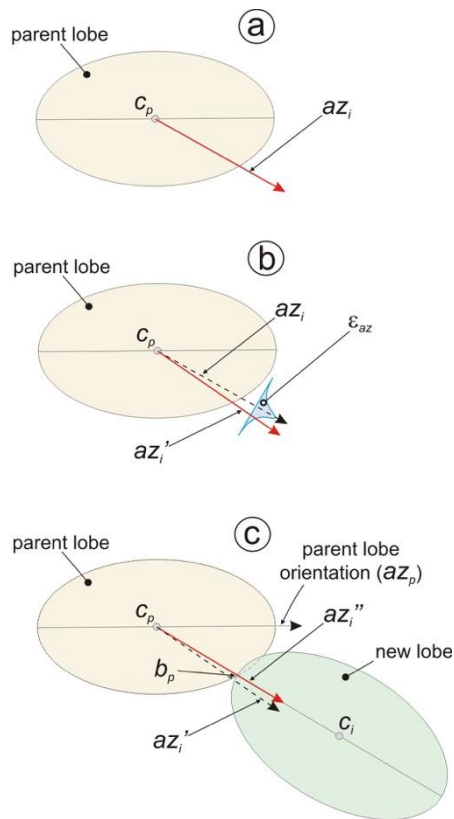
$$182 e_i = \sqrt{1 - \frac{b_i^2}{a_i^2}}; \quad \pi a_i b_i = \text{lobe area.}$$

183

184 The first parcel is set on the vent, and each subsequent parcel buds from  
 185 the edge of an older parcel (denoted as parent parcel), which has been already  
 186 set. The point where the new parcel will bud (budding point,  $b_p$ ) is determined  
 187 according to the following steps:

188

- 189 • STEP 1. A preliminary direction  $az_i$  is given by the direction from the center  
 190 of the parent lobe ( $c_p$ ) to the point of minimum elevation on the boundary  
 191 of the elliptical parcel (Figure 2a);
- 192 • STEP 2. A random perturbation  $\varepsilon_{az}$ , with a probability distribution defined as  
 193 a function of the local slope, is added to  $az_i$  resulting in a new preliminary  
 194 orientation  $az_i'$  (Figure 2b).
- 195 • STEP 3. An inertial factor accounting for the orientation of the parent lobe  
 196 ( $az_p$ ) is considered, resulting in the final orientation  $az_i''=f(az_i',az_p)$  (Figure  
 197 2c). This final direction from the center of the parent lobe determines the  
 198 budding point  $b_p$  on its boundary.



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 201  
 202 **Figure 2.** Sketch of the positioning of a new lobe (light green ellipse) with  
 203 respect to the selected parent lobe (light yellow ellipse). The major axes of both  
 204 ellipses are plotted. Panels (a) to (c) illustrate the three steps implemented and  
 205 described in the main text. The relative weight between the three factors (a) to  
 206 (c) depends on the values selected for the relevant parameters. The budding  
 207 point  $b_p$  always lays on the boundary of the ellipse representing the parent

208 lobe, and the line joining  $c_p$  and  $b_p$  identifies the direction of propagation of the  
 209 flow. The center  $c_i$  of the new lobe always lays along this direction  $azi''$ , and is  
 210 defined by the distance  $d$  between  $c_p$  and  $c_i$ , which can vary between a  
 211 minimum ( $c_i = b_p$ ) and a maximum (the new ellipse passes through  $b_p$ ) according  
 212 to the value set in input to the parameter *dist\_fact* within the interval [0;1]. The  
 213 latter case (*dist\_fact* = 1) is illustrated in figure.

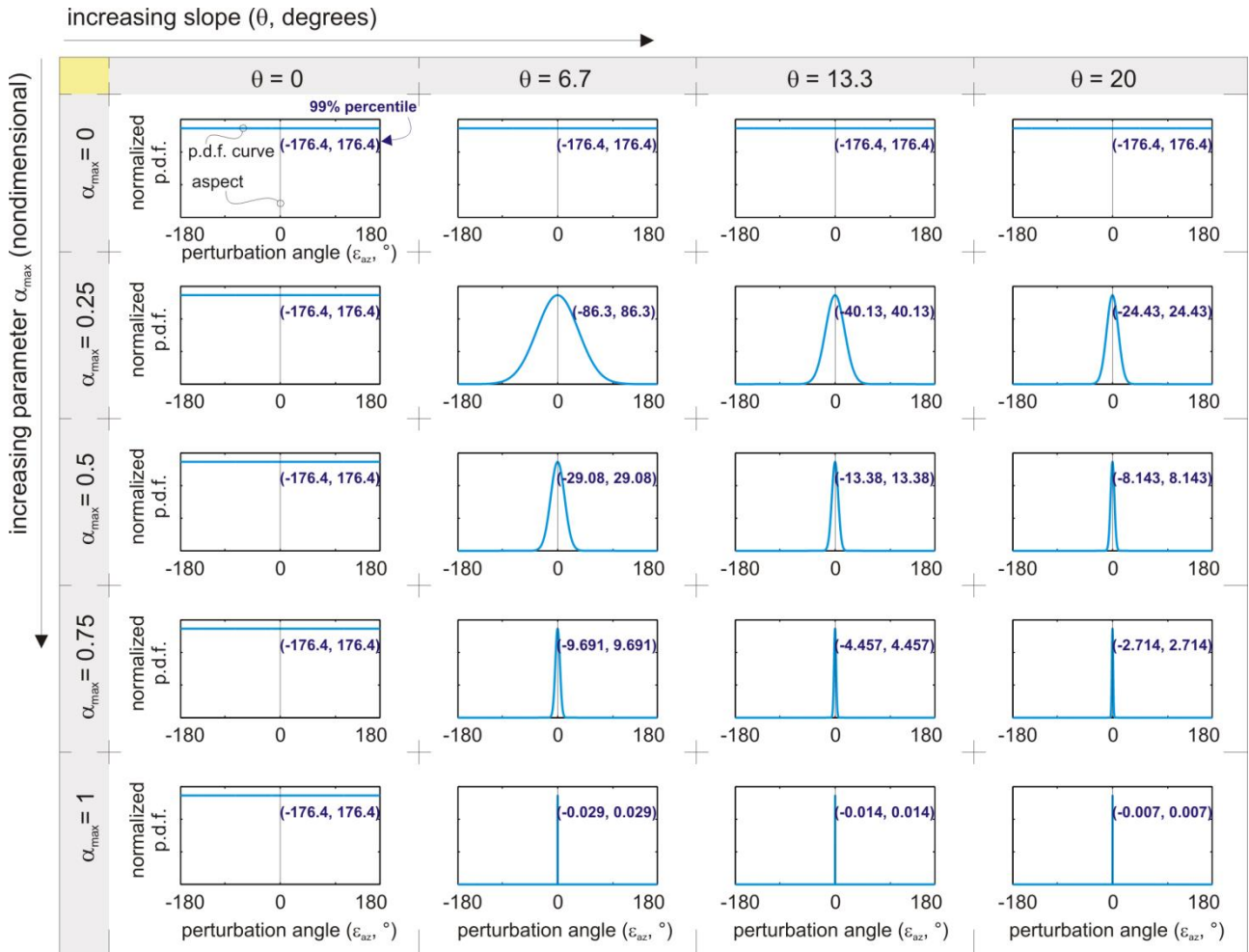
214  
 215 The probability distribution of the random perturbation ( $\varepsilon_{az}$ ) introduced in the  
 216 second step is determined by one of the set input parameters, and is ruled by  
 217 two additional parameters: (i) the local slope ( $\theta$ , degrees); and (ii) a  
 218 dimensionless input parameter  $\alpha_{max}$ , which can vary within the interval [0;1]. As  
 219 stated above, the flatter the topography, the wider the range of possible  
 220 directions of propagation. Consequently, the steeper the slope in  $c_p$ , the lower  
 221 the probability of having a large perturbation ( $\varepsilon_{az}$ ) and the closer  $azi'$  will be to  
 222 the preliminary direction ( $azi$ ). The probability distribution associated with the  
 223 perturbation is a truncated normal distribution in the interval [-180;180]  
 224 (Chopin 2011) with zero as the mean value and with a standard deviation  
 225 defined by the following equation:

$$226 \quad S = \frac{(1 - \alpha_{max})(90 - \theta)}{\alpha_{max}\theta} \quad (1)$$

227 When  $\alpha_{max} = 0^\circ$ , or when the local slope  $\theta = 0$ , all perturbations have the same  
 228 probability, resulting in a random choice of  $azi'$ . When  $\alpha_{max} > 0$ , a null  
 229 perturbation (the maximum slope direction) has a larger probability, and this  
 230 probability increases with increasing value of  $\alpha_{max}$ . If  $\alpha_{max} = 1$  (excluding when  
 231  $\theta = 0$ ) no perturbation is added to  $azi$  resulting in a new preliminary orientation  
 232  $azi' = azi$ . (i.e. the budding process follows strictly the steepest descent path,  
 233 making the choice of the budding point purely deterministic). The probability  
 234 density function for several values of  $\alpha_{max}$  and  $\theta$  are plotted in Figure 3.

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238 **Figure 3.** Probability density function (p.d.f.) of the angular perturbation  $\varepsilon_{az}$   
 239 added to the local aspect  $az_i$ . When local slope ( $\theta$ ) is null or the input parameter  
 240  $\alpha_{\max} = 0$ , the probability distribution from which  $\varepsilon_{az}$  is sampled is uniform; as  $\theta$   
 241 and/or  $\alpha_{\max}$  increase, the probability of large values of  $\varepsilon_{az}$  decreases, so that  $az_i'$   
 242 tends to coincide with the local steepest descent.

243

244 The inertial contribution is a deterministic function of the slope at  $c_p$  (the  
 245 shallower the slope in  $c_p$  the larger the inertial contribution) and of a further  
 246 input parameter *inertial\_exponent* controlling how much the new orientation  
 247  $az_i''$  is inherited from the parent lobe. This negative relation between the  
 248 *inertial\_exponent* and slope may sound counterintuitive, but one has to take  
 249 into account that a higher slope typically promotes flow bifurcation (Dietterich  
 250 and Cashman 2014), which always result in lateral spreading of the flow field  
 251 (Tarquini and Favalli 2011; 2016). In contrast, emplacement on shallow slopes

252 tend to follow a straight path, as seems to have happened during the  
253 emplacement of the 180 km long pahoehoe lava flow described by [Giacomini et](#)  
254 [al. \(2009\)](#) in the almost flat Payen volcanic Complex (Argentina), or as happened  
255 during the 2014-2015 Holuhraun eruption (Iceland), when the initial flow  
256 (slabby pahoehoe to aha in nature, [Pedersen et al., 2017](#)) progressed straight  
257 along the steepest path for about 20 km driven by a very low average slope of  
258 0.2 degrees ([Tarquini et al., 2015](#)).

259 Finally, the eccentricity  $e_i$  of the budding lobe is defined as a function of  
260 the slope from the center of the parent lobe to the budding point  $b_p$ , defined as  
261 the intersection between the boundary of the parent lobe and the ray from the  
262 center of the lobe with direction  $azi''$  ([Figure 2](#)). The steeper the slope the larger  
263 the eccentricity, with circular lobes forming over a flat topography. We remark  
264 here that while the eccentricity of the lobes changes as a function of the slope,  
265 the area and the thickness are fixed (unless otherwise stated in the input  
266 settings).

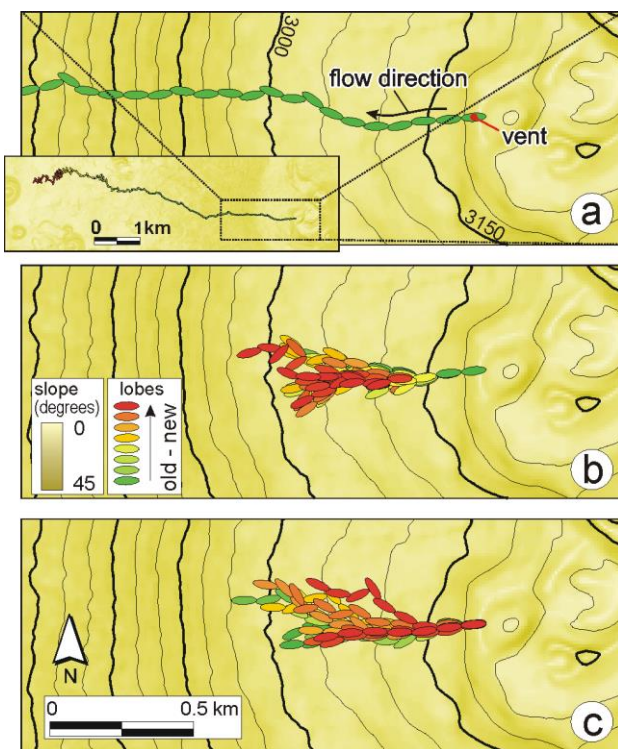
267

### 268 3.2. Budding of parcels

269 The code proceeds by iteratively setting new parcels on the topography  
270 until their total volume equals that prescribed for the simulation. The first  
271 parcel is set on the vent. The second parcel will bud from the first one, which is  
272 the only one already set on the topography. From the third parcel onward it is  
273 necessary to make a choice to select which will be the parent parcel among all  
274 the parcels already positioned. This choice, input parameter *lobe\_exponent*, has  
275 a critical importance to the propagation of the flow, and is defined with a value  
276 within the interval [0;1]. The closer *lobe\_exponent* is to 0, the higher the  
277 probability that younger lobes will be the parent of a new lobe. If  
278 *lobe\_exponent* = 0, the youngest lobe will be the parent lobe of the next, and a  
279 single chain of lobes trending downhill originates (figure 4a). At the other  
280 extreme, when *lobe\_exponent* = 1, all existing lobes have the same probability  
281 to be the parent, and the code generates more complex features, with frequent  
282 overlapping between lobes and widening of the flow through multiple  
283 branching (figure 4b). When *lobe\_exponent* = 0 (budding from the last lobe), the

284 user can select a maximum number of lobes constituting a chain, and the  
285 creation of a chain will be reiterated several times generating a series of smaller  
286 chains (Figure 4c) until the total volume is considered. The latter case mimics  
287 the creation of many perturbed steepest descent paths implemented in the  
288 DOWNFLOW code (Favalli et al. 2005). The number of parcels forming the  
289 chains can be fixed as a constant or as variable within a range, and influences  
290 the obtained runout. Several further tuning options are implemented, for  
291 example to account for a given propensity to lengthening, widening or  
292 thickening of the flow field, and details are provided in the largely commented  
293 source code available at the model repository  
294 (<http://demichie.github.io/MrLavaLoba/>).

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299 **Figure 4.** Simplified MrLavaLoba outputs illustrating the result of the budding of  
300 150 parcels ( $2000 \text{ m}^2$  in area and  $0.5 \text{ m}$  in thickness each for a cumulative  $150 \times$   
301  $10^3 \text{ m}^3$ ) with different input settings. All the other input parameters remained  
302 fixed. (a) shows the output with *lobe\_exponent* = 0 (budding is always from the  
303 youngest lobe) and without restarts from the vent (all the lobes are visible in

304 the inset); (b) shows the output with *lobe\_exponent* = 1 (budding is from a  
305 random parent lobe) and without restarts from the vent; (c) shows the output  
306 when *lobe\_exponent* = 0 as in (a), but with multiple restarts from the vent.

307

308 In addition to being at a given position and possessing a given geometry,  
309 each parcel is also assigned with two additional attributes: (i) its distance  
310 downflow from the vent, and (ii) its number of descendant parcels. The distance  
311 from the vent is an integer value that sums the number of parcels (i.e. the  
312 number of budding processes) separating each parcel from the vent. This  
313 parameter tracks the length of the path (ideally) followed by each lava parcel to  
314 attain its final position, and is computed for each new parcel as the distance  
315 value of its parent lobe increased by one. The number of descendants is an  
316 integer value, and is updated each time a new parcel is set by adding one to all  
317 the ancestor parcels, from the parent lobe backward to the vent. If we assume  
318 that a newly generated lobe has travelled from the vent to its actual settling  
319 position passing through all ancestor parcel positions, then the number of  
320 descendants represents the number of lava parcels that passed through a  
321 particular lobe, and this information can be used to describe the relative  
322 probability of a given model cell being inundated.

323

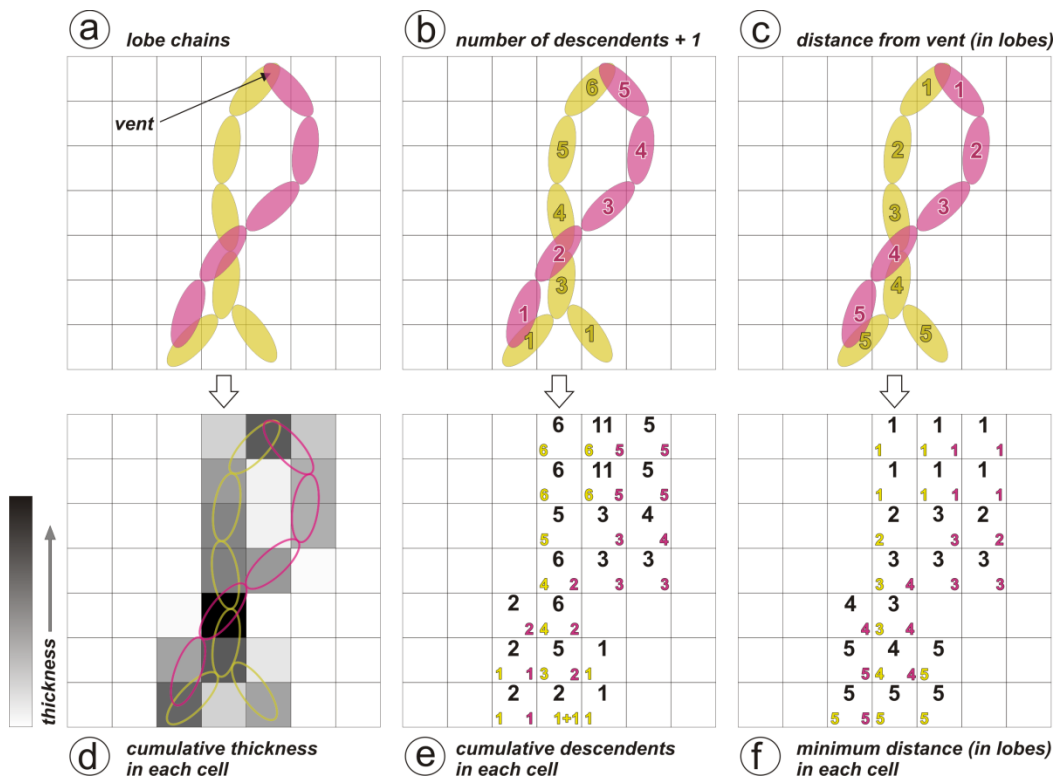
324

### 325 *3.3. The output of a simulation*

326 Once the selected number of parcels has been settled as vector lobes on  
327 the topography (Figure 4), this result is processed to derive an output in grid  
328 format (Figure 5). For convenience, any output grid has the same cell size and  
329 the same extent as the input topography used as computational domain. It is  
330 important to stress that, to improve the readability, the lobe and the grid cell in  
331 Figure 4 have a rather similar size, but in performing real simulations the lobe  
332 should be about 5-10 times bigger than the grid cell in order to optimize the  
333 computational time (see also section 3.6).

334 Up to three grids can be saved as output: (i) the grid representing, at each  
335 cell, the thickness of the lava deposit, obtained as the cumulative volume of the

336 portions of the parcels covering each cell divided by the cell area; (ii) the grid  
 337 representing, at each cell, the sum of descendant parcels for all the parcels  
 338 touching that cell, indicating the total number of transits in that cell  
 339 (proportional to the local probability of inundation, or the hazard); and (iii) the  
 340 grid expressing, at each cell, the minimum along-path distance from the vent  
 341 among all the parcels covering that cell (indicative of the minimum time  
 342 required to reach the cell). For the latter grid, the distance is computed as  
 343 described in the previous section.



344 **Figure 5.** Examples of simulation outputs: from vector (a-c) to raster  
 345 format (d-f). (a) Two lobe-chains generated from the vent with different  
 346 budding rules: budding strictly from the last lobe (red chain) or not (yellow  
 347 chain); (b) each lobe is assigned with the number of descendants (+1); (c) each  
 348 lobe is assigned with the distance from vent measured as number of lobes (by  
 349 following backwards the budding process); (d) output grid where the thickness  
 350 in each cell is assigned according to the portion(s) of the lobe(s) covering it; (e)  
 351 output grid of the hazard in which each cell is assigned with the sum of the  
 352 descendants (black number) of all the touching lobes (discarding direct  
 353 budding); (f) output grid of the minimum distance from the vent.

355

356 The code allows for saving *masked* grids obtained by considering only  
357 some of the inundated (or touched) cells. This option is controlled by the input  
358 parameter *masking\_threshold*, expressed in fraction of the total volume (i.e.  
359 varying within the interval [0; 1]). As an example, if *masking\_threshold* is set to  
360 0.95 for the thickness grid, the thinnest portion of the final lava deposit  
361 representing 5% of the total volume emplaced is filtered from the results (see  
362 Figure 6 and compare maps in the first and second rows). When the masking is  
363 applied to the hazard grid, the threshold function removes 5% of the total  
364 volume emplaced in the areas with the lowest probability of being inundated.  
365 The masking threshold is fundamental for the definition of convergence of the  
366 outputs of the code, as described in the next section.

367

### 368 3.4. The convergence of simulations

369 MrLavaLoba is a probabilistic model, and repeated runs always yield  
370 different results even if input parameters are unchanged. If this uncertainty is  
371 not assessed, the latter point can undermine the reliability of a simulation. For a  
372 fixed total emplaced volume and input settings, reducing the thickness of lava  
373 parcels means an increase in the total number of parcels, and hence an increase  
374 in the statistical meaningfulness of the lava spreading process. For each model  
375 realization, the local thickness (i.e. the number of parcels over a cell) can be  
376 seen as the probability of a point of the topography to be inundated. We have  
377 found that using a sufficiently high number of parcels, for a fixed cumulative  
378 volume, fixed input settings and a fixed threshold (*masking\_threshold*), the  
379 masked area of inundation, obtained by removing the low probability tails of  
380 the deposit, converges to a univocal high probability region. We do not find a  
381 convergence using the complete unfiltered results. The convergence of the  
382 masked result demonstrates the dependability of the simulation output.

383 The convergence can be tested by iterating the following procedure a  
384 number,  $N$ , of times. In the first iteration (i.e.  $i = 0$ ), considering a fixed  
385 cumulative volume of lava, a fixed number of lava parcels and fixed input  
386 settings, the simulation is repeated twice, producing two output grids, which  
387 are then masked (e.g. by setting *masking\_threshold* = 0.95) obtaining the two

388 grids  $G_{i-1}$  and  $G_{i-2}$ . We denote with  $C_{i-1}$  and  $C_{i-2}$  the area covered by the masked  
389 grids, and we define the coverage convergence index  $\mu_i$  (quantifying the  
390 similarity between the two masked results) as follows:

391

$$392 \quad \mu_i = \frac{C_{i-1} \cap C_{i-2}}{C_{i-1} \cup C_{i-2}} \quad (2)$$

393

394 At each new iteration (i.e.  $i = 1, 2, 3 \dots N$ ), the number of parcels is multiplied by  
395 a given factor (here set to 2), while the thickness of parcels is divided by the  
396 same factor (all the other setting remaining fixed), thus keeping constant the  
397 cumulative volume. By applying equation (2) at each iteration, a series of  
398 coverage convergence indexes is obtained ( $\mu_1, \mu_2 \dots \mu_N$ ). If the obtained  
399 sequence tends to 1, then the masked inundated area converges towards a  
400 univocal solution.

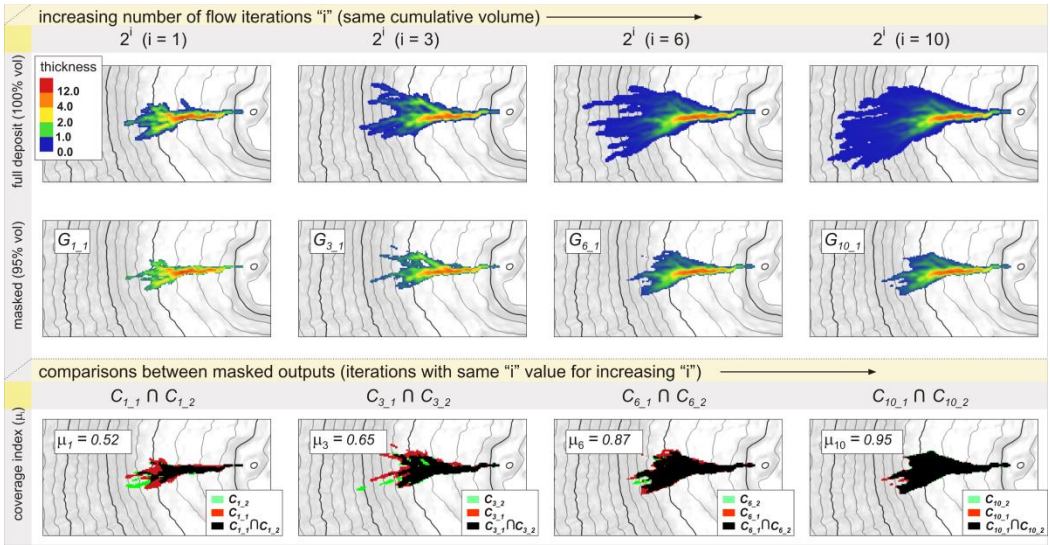
401 In order to illustrate the convergence, we carried out a series of tests with  
402 a cumulative  $150 \times 10^3 \text{ m}^3$  bulk volume on the Etna topography. For the first  
403 iteration of the convergence test, 150 lobes ( $2000 \text{ m}^2$  in area each as in [Figure](#)  
404 [4](#)) are used, resulting in a constant thickness of 0.5 m for each lobe. In the  
405 second iteration the number of lobes is doubled, with 300 parcels having  
406 constant thickness of 0.25 m, and so on up to the 11<sup>th</sup> iteration when 156 300  
407 lobes approximately 0.5 mm thick are considered for the settling process.  
408 [Figure 6](#) shows graphically a few examples of the iterative computation of the  
409 coverage convergence index for  $\alpha_{max} = 0.9$  and  $lobe\_exponent = 0.6$ ,  
410 highlighting the convergence of the masked output to a univocal area (thus  
411 implying that  $\mu_i$  tends to 1 as the number of iteration of the procedure  
412 increases). [Figure 7a-b](#) further illustrates that MrLavaLoba simulation outputs  
413 converge toward a unique solution as the number of parcels increases for a  
414 suite of combinations of different values for input parameters. The convergence  
415 of  $\mu_i$  to 1 is not monotonic because the model is probabilistic.

416 In parallel with the convergence of the coverage convergence index  $\mu_i$  we  
417 also checked the convergence of the thickness of the masked deposit (i.e.  $G_{i-1}$   
418 and  $G_{i-2}$  defined above) by verifying that the discrepancy between the two

419 outputs obtained at each iteration (defined below by the additional thickness  
 420 convergence index  $\tau_i$ ) vanishes to zero as the number of iterations increases  
 421 (Figure 7c-d):

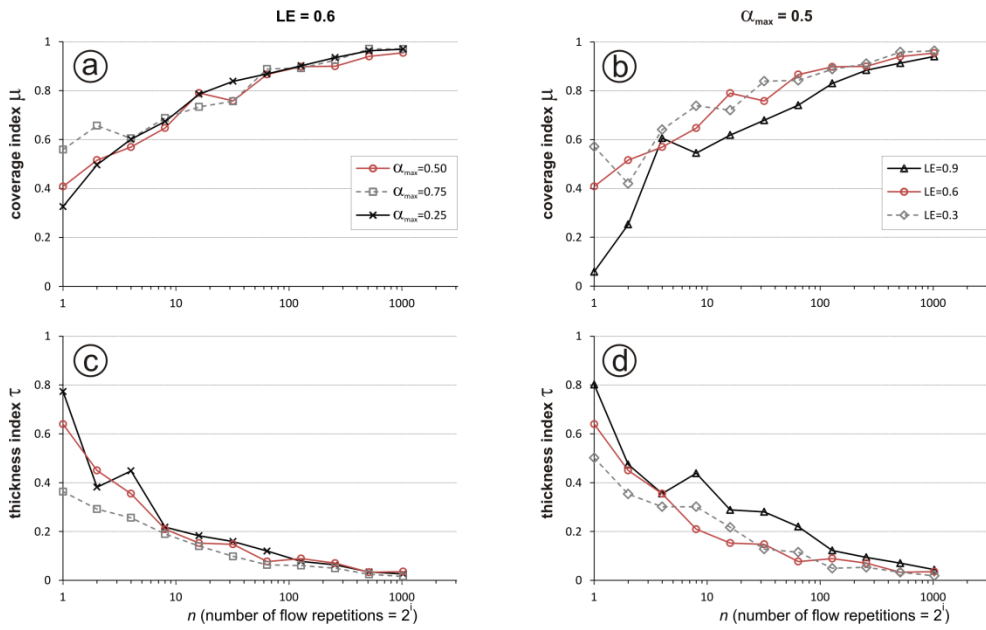
$$422 \quad \tau_i = \frac{V_d}{\max(V_{i-1}, V_{i-2})} \quad (3)$$

423 where  $V_d = \int_{C_{i-1} \cap C_{i-2}} \text{abs}(G_{i-1} - G_{i-2})$ ,  $V_{i-1} = \int_{C_{i-1}} G_{i-1}$  and  $V_{i-2} = \int_{C_{i-1} \cap C_{i-2}} G_{i-2}$ . For an  
 424 increasing number of iterations  $i$ , if  $\tau_i$  tends to zero (as in Figure 7c-d), then the  
 425 output of the simulation tends to an univocal solution (i.e. a given grid of the  
 426 thickness). The performed tests also indicate that for a number of iterations  
 427 close to  $10^3$ , the obtained results appear to converge, as shown by the trends in  
 428 the plots of both convergence indexes (Figure 7).  
 429  
 430  
 431



432 **Figure 6.** Convergence of the outputs of MrLavaLoba simulations. First row:  
 433 unmasked outputs of simulations (thickness grid) obtained with an increasing  
 434 number of parcels. Second row: same as above, but with masked outputs  
 435 ( $\text{masking\_threshold} = 0.95$ ). Third row: intersections between coverage of the  
 436 masked outputs to compute  $\mu_i$ .  
 437  
 438

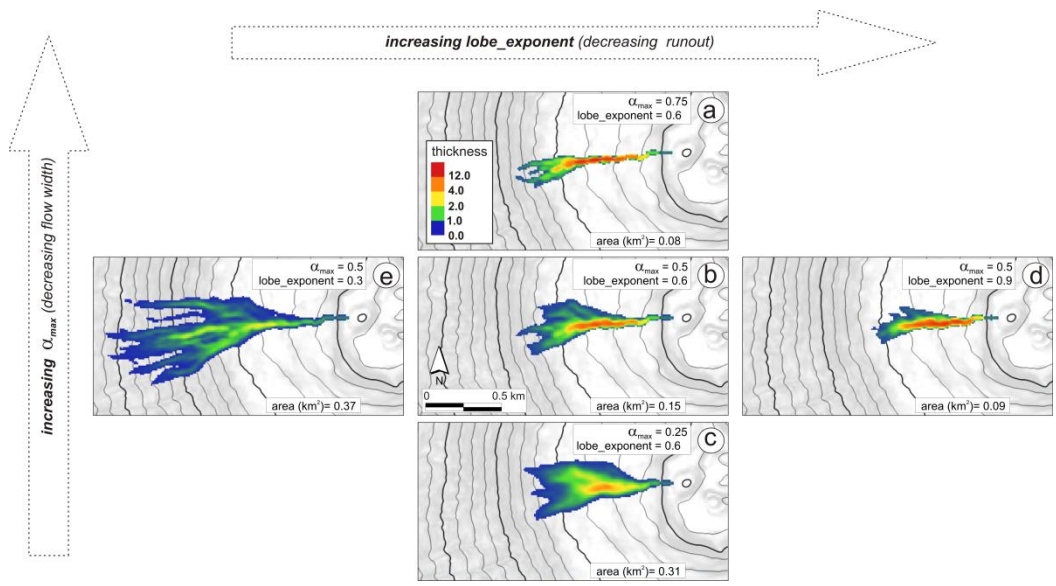




439  
 440 **Figure 7.** Convergence of the coverage and thickness of the masked grids for  
 441 different values of *lobe\_exponent* (LE) and  $\alpha_{max}$ . The coverage convergence  
 442 index  $\mu_i$  tends to 1 (a-b) and thickness convergence index  $\tau_i$  tends to 0 (c-d) as  
 443 the number of iterations increases. The obtained plots support that the  
 444 simulations always converge towards a unique result.

445  
 446 In order to explore graphically how the convergence solution varies according  
 447 to different input settings, the grids of the thickness obtained in performing the  
 448 convergence test for varying input parameters are shown in **Figure 8**.

449  
 450



451  
 452

453 **Figure 8.** Masked grids of the thickness (for 156300 lava parcels) illustrating  
454 how the output of MrLavaLoba simulations change as the main input  
455 parameters *max\_slope\_prob* and *lobe\_exponent* also change (all other factors  
456 remaining constant).

### 457 458 *3.5. Accounting for the effect of syn-emplacement topographic changes*

459 The key characteristic of MrLavaLoba, which marks a difference with  
460 respect to other probabilistic codes such as DOWNFLOW (Favalli et al. 2005), is  
461 the fact that computational parcels are assigned a volume. Users can choose to  
462 account or not for the modification of the topography due to the volume of the  
463 already emplaced parcels.

464 MrLavaLoba accounts for the modification of the topography through two  
465 specific settings: (i) the parameter *topo\_mod\_flag*, identifying how often the  
466 changes in topography are computed, and the parameter  
467 *thickening\_parameter*, determining the weight assigned to these changes when  
468 computing the direction of propagation of new lobes. The more often changes  
469 are computed and the lower the *thickening\_parameter* assigned (which can  
470 vary between 0 and 1), the higher the lateral spreading of the flow and the  
471 lower the thickness of the final lava deposit. It is important to highlight that  
472 even when changes in topography are considered during the emplacement, the  
473 outputs converge as shown in Figures 6-7. In general, this specific setting can be  
474 used to reproduce a given trend of supply rate (and consequent emplacement  
475 style) such as a continuous or episodic supply (Guest et al. 1987; Harris and Neri  
476 2002).

### 477 478 *3.6. The size of the parcel*

479 Within a single simulation, all the parcels have the same area defined by  
480 the user as an input parameter. The optimal area of the parcel depends also on  
481 the cell size of the grid used as computational domain (i.e. the digital elevation  
482 model representing the topography). Considering computational lobes having  
483 an area smaller than the cell size of the computational grid does not improve  
484 the accuracy of simulations. In general, it is useful to keep the area of the parcel

485 as large as possible, without reducing the accuracy of the results (assessed on a  
486 case by case basis). By running the code preliminarily with a spectrum of  
487 different parcel sizes, it is usually easy to select the size which provides the  
488 optimal trade-off, because outputs obtained by considering exceedingly small  
489 or exceedingly large parcels produce too small or too large flow units, while the  
490 right size (i.e. the one similar to the real flow units) is missing. As a general rule,  
491 care should be made to choose a parcel size smaller than a characteristic width  
492 of typical flow units for a given scenario. As an example, in the case of  
493 channelized flow at Mt Etna considered in [Figures 4, 5 and 8](#) (using a 10 m-cell  
494 size DEM – 100 m<sup>2</sup> cell area), we realized that there is a characteristic size of the  
495 features formed during the emplacement (e.g. characteristic dimension of the  
496 flow units), which are better reproduced by a parcel size of 1000-2000 m<sup>2</sup>, while  
497 a significantly smaller parcel size produces narrower features which are  
498 negligible in real cases.

499 The parcel thickness is determined by the cumulative volume of the  
500 effusive episode in question, the parcel area and the total number of parcels. A  
501 higher number of parcels reduces the thickness and consequently the parcel  
502 volume. Thus, for a given cumulative lava volume, the computational time will  
503 increase proportionally to the number of parcels.

504

#### 505 **4. Using the code and case studies**

506

507 MrLavaLoba has been written in Python and the source code is accessible  
508 at the code repository. Python is a widely used general-purpose, high-level  
509 programming language designed to emphasize code readability and conciseness  
510 ([Oliphant 2007](#)). Python interpreters are freely downloadable for installation on  
511 any operating system (e.g. Windows, Apple, Linux, Android), so that Python  
512 codes have great portability. Moreover, several GIS software such as Q-GIS,  
513 Grass and ArcGIS, provide a simple way to embed Python code in their GIS  
514 environment ([Bartolini et al., 2013](#)). It has been decided to release the code  
515 without embedding it in a specific GIS environment or interface for the  
516 following reasons: (i) to leave all future users free to load the simulation output

517 into their own GIS software; and (ii) to force future users to read the largely  
518 commented list of input settings, thus becoming a bit more familiar with the  
519 many nuances of MrLavaLoba.

520 The easiest way to become familiar with the code is to run it for a given  
521 scenario. To support this option, the test cases illustrated in the following can  
522 be downloaded from the code repository.

523

#### 524 *4.1. Input settings*

525 As already highlighted, the code is structured to work with a large spectrum of  
526 input settings. In order to make tuning easier, we arranged input parameters  
527 into two input files: the first one (“input\_data.py”) collects basic settings such  
528 as type and coordinates of the vent(s), the size of the parcels, the cumulative  
529 volume of lava, the reference topography and all the very fundamental tuning  
530 parameters; the second input file (“input\_data\_advanced.py”), instead, includes  
531 secondary settings which can be tuned for a more advanced calibration, but can  
532 be left in their default values during the initial tuning. Examples of input files for  
533 the current version of MrLavaLoba are provided along with the source Python  
534 code. All the input parameters are listed and commented in detail to minimize  
535 misunderstandings.

536

537

#### 538 *4.2. Application to real cases*

539 In this section, we present the application of MrLavaLoba to two real cases of  
540 lava flows emplaced according to very different dynamics at Hawaii and Mt Etna  
541 (i.e. pahoehoe and channelized-‘a‘ā, respectively).

542

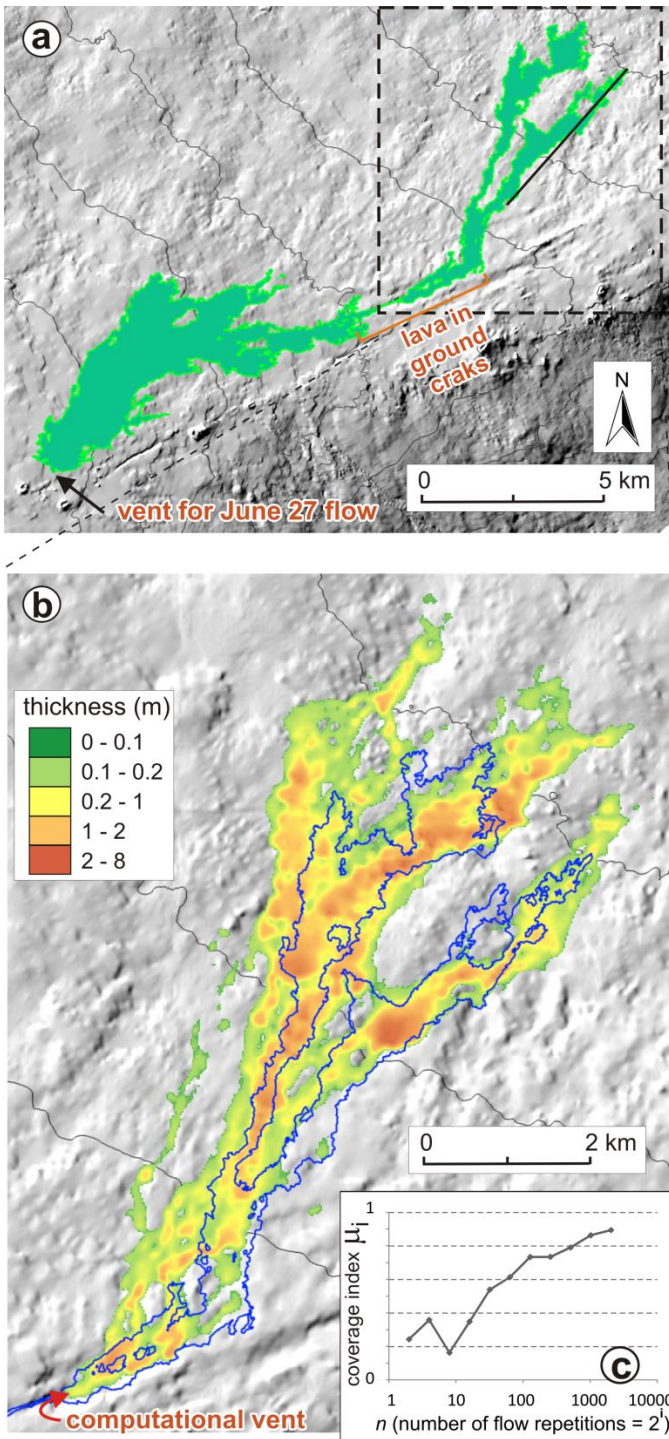
##### 543 *4.2.1. The distal sector of the Kilauea 2014-2016 pahoehoe flow field*

544 As a first test case, we considered the pahoehoe flow field which erupted at  
545 Kilauea volcano (Hawaii) between June 27, 2014 - June 8, 2015 ([Patrick et al.  
546 2016](#)). On June 27, 2014, the opening of new vents on the north flank of Pu'u  
547 'Ō'ō changed the hazard potential of ensuing lava flows dramatically, because  
548 lava advanced in a general northeast direction, posing a threat to the village of

549 Pahoia and other inhabited areas (Patrick et al. 2016). 7-8 km down-flow from  
550 the vent, lavas repeatedly entered pre-existing ground cracks along the East Rift  
551 Zone (ERZ), and continued to advance northeastwards for a few kilometers  
552 (Figure 9a). In early September lava returned to the surface, heading in a more  
553 northward direction, and continued advancing until March 2015. Given the  
554 complexity of the whole episode, and the peculiar presence of the ground  
555 cracks, mostly smaller the 20 m-cell size resolution of the topography used  
556 here, we decided to simulate only the terminal segment, setting a  
557 computational vent along the ERZ where lava exited from the ground cracks to  
558 flow again on the surface (Figure 9b). In this case, our simulations were  
559 supported by USGS colleagues based at the Hawaiian observatory, who  
560 provided a pre-emplacement digital elevation model (a 20 m-resolution DEM  
561 obtained from the 10 m-resolution NED topography, Gesch et al., 2002) and  
562 lava coverage in addition to input settings such as an estimation of the volume  
563 of lava flowed through the computational vent considered. Unfortunately, no  
564 map of the thickness of the final lava deposit was available to validate the  
565 simulation results with respect to the thickness.

566         Considering the complexity of the scenario and the relatively low quality  
567 of the available DEM, the obtained result can be considered promising. Defining  
568 a fitness index for the simulation as  $\phi = (A_R \cap A_S) / (A_R \cup A_S)$ , where  $A_R$  is the  
569 area covered by the real flow and  $A_S$  the area covered by our simulation (Favalli  
570 et al., 2009c), we obtain  $\phi = 0.36$ . It is worthy to note that the masked output  
571 shown in Figure 9b presents isolated pockets. This appears, of course,  
572 unrealistic if we think about a real lava flow, but, when considered from a  
573 probabilistic point of view, those pods represent areas more likely inundated  
574 than the immediate surroundings.

575

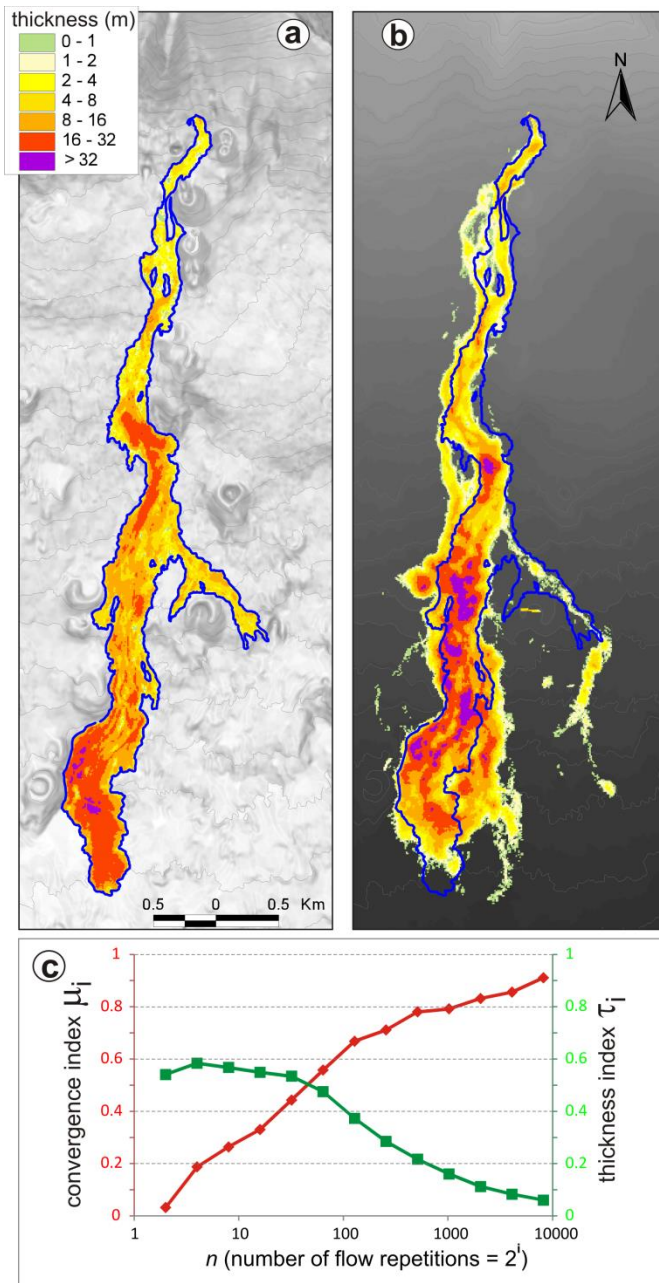


576  
 577 **Figure 9.** Simulation of the distal segment of the Kilauea 2014-2016 pahoehoe  
 578 flow field. (a) map of the total flow field (after Patrick et al. 2016), contour lines  
 579 are spaced at 100 m in elevation. (b) simulation result: masked grid of the  
 580 thickness. (c) plot of the convergence for the selected input settings. The 20 m-  
 581 cell size pre-emplacment topography is courtesy of M Patrick from USGS. Input  
 582 parameters used as well as the DEM are accessible at the model repository.

583  
 584 **4.2.2. The Mt Etna 2001 LFS1 lava flow**

585 The second test case considered is the Mt Etna 2001 LFS1 flow field, which  
586 formed from July 18 - August 9 2001 from a vent located at about 2100 m  
587 elevation near Monti Calcarazzi (Coltelli et al. 2007). This flow is becoming a  
588 kind of a benchmark for lava flow simulation codes, and has been simulated  
589 repeatedly over the years with different codes (Crisci et al. 2004; Proietti et al.  
590 2009; Tarquini and Favalli 2011; Mossoux et al. 2016). The cumulative volume  
591 of lava emplaced is known from the existing literature (Coltelli et al. 2007), and  
592 the map of the lava deposit thickness has been already presented by Tarquini  
593 and Favalli (2011, Figure 10a). Here we used as pre-emplacement topography  
594 the 10 m-resolution TINITALY digital elevation model (Tarquini et al. 2007;  
595 2012b). The flow thickness for a simulation obtained with  $2^{13}$  flow repetitions is  
596 shown in Figure 10b. By comparing this coverage area for this particular run  
597 with the real coverage, we have obtained a fitness index  $\phi = 0.47$ . The  
598 simulation tends to overestimate the inundated area. As for the obtained  
599 thickness, in the first half of the flow the obtained thickness are consistently  
600 low, and both the real and the simulated flow show a peak above 30 m at about  
601 3 km from the vent. In the second half of the flow, instead, the mass  
602 distribution is more distributed towards the front in the real case than in the  
603 simulated one. Figure 10c shows the convergence of the simulation results in  
604 terms of both area coverage and local thickness. Similarly to the Hawaiian case  
605 presented in the previous section, both input data and DEM are provided as  
606 supplementary material.

607



608  
 609 **Figure 10.** Simulation of the Mt Etna 2001 LFS1 ‘a’ā flow field. (a) map of the  
 610 thickness of the final flow field (obtained from DEM difference). (b) simulation  
 611 results: masked grid of the thickness. (c) plot of the convergence for the  
 612 selected input settings (for both coverage and thickness). The 10 m-cell size pre-  
 613 emplacement topography is provided by INGV (Tarquini et al., 2012b). Input  
 614 settings as well as the DEM are accessible at the code repository. The obtained  
 615 fitness index for the simulation is  $\phi = 0.47$ . Time for a simulation on a standard  
 616 PC is about 2’30’’ for the budding of > 200,000 lobes having an area 10 times  
 617 the cellsize of the computational DEM.

618



619 **5. Discussion**

620 The main principle at the foundation of MrLavaLoba is that uncertainty always  
621 affects the modeling of lava flows. This uncertainty is due both to the  
622 uncertainty of observations of critical input parameters such as the effusion  
623 rate ([Harris et al. 2007](#)) or the DEM, and to the extraordinary complexity in the  
624 emplacement processes of pahoehoe and ‘a‘ā lavas ([Baloga and Glaze 2003](#);  
625 [Applegarth et al. 2010a, 2010b](#); [Favalli et al. 2010](#); [James et al. 2012](#); [Glaze and](#)  
626 [Baloga 2013](#)). This concept has been largely validated for channelized flows by  
627 the success of the probabilistic code DOWNFLOW ([Favalli et al. 2005, 2009,](#)  
628 [2011, 2012](#)), which embed a Monte Carlo approach consisting in the repetition  
629 of many, slightly perturbed steepest descent paths. A probabilistic approach is  
630 also followed by [Glaze and Baloga \(2013\)](#) in approaching the simulation of the  
631 emplacement of pahoehoe lavas. MrLavaLoba uses an approach similar to that  
632 of [Glaze and Baloga \(2013\)](#), with the scientific advancement of utilizing a  
633 stochastic perturbation of the main driving factor of lava flow emplacement (i.e.  
634 the steepest descent path), introducing a variety of tunable settings, and the  
635 probabilistic assessment of the convergence of the results.

636 In cellular automata models such as SCIARA ([Crisci et al. 2010](#)), changes to  
637 the pre-emplacement topography due to the emplacing lava is considered at  
638 each computational step. This appears to be a conservative choice. In contrast,  
639 previous probabilistic codes such as DOWNFLOW do not account at all for  
640 changes in the topography due to the emplacing lava. For short-lived or  
641 relatively simple flows, it is not necessary to account for changes in volume as  
642 often as is done in cellular automata codes to achieve realistic results ([Tarquini](#)  
643 [and Favalli 2016](#)). MrLavaLoba provides an intermediate solution in between  
644 these two extremes, and changes to the pre-emplacement topography  
645 determined by the volume of the emplacing lava flows can be calculated at a  
646 desired frequency by tuning specific parameters.

647

648 **6. Conclusion**

649

650 We have presented MrLavaLoba, a new code for the simulation of the area  
651 inundated by lava flows, freely downloadable at  
652 <http://demichie.github.io/MrLavaLoba>. MrLavaLoba is a probabilistic model  
653 that calculates the most probable areas that could be inundated by lava and the  
654 thickness of the deposits. The code does not provide directly the progression of  
655 a flow field with time. The code is written to be versatile. Input files provide  
656 tunable parameters that enable the generation of realistic results for a range of  
657 lava types and characteristics and for different emplacement dynamics.  
658 MrLavaLoba input parameters must be tuned for each new lava flow scenario  
659 based on an appropriate case. Upon proper parameters calibration, the code  
660 produces realistic results in different lava flow scenarios, from pahoehoe flows  
661 to channelized 'a'ā flows. Simulation results must be interpreted with the care  
662 due to the highly tunable and probabilistic nature of the model.

663

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672

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