

1 Temporal variations of helium isotopes in volcanic gases  
2 quantify pre-eruptive refill and pressurization in magma  
3 reservoirs: The Mount Etna case

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8 **ABSTRACT**

9 Two approaches to the challenging aim of forecasting impending eruptions are  
10 searching for empirical precursors and developing suitable interpretative models. Here  
11 we present high-resolution time series of  $^3\text{He}/^4\text{He}$  ratios measured in gases emitted from  
12 peripheral vents around Mount Etna volcano (Italy), which revealed variations with  
13 strong correlations over both time and a broad spatial scale. The main eruptive episodes  
14 are preceded by increases in  $^3\text{He}/^4\text{He}$ , making this ratio a unique tracer for monitoring  
15 volcanic activity. These features strongly reflect pressurization beneath the volcano due  
16 to deep magma influx. We propose a pioneering model that relates the changes in  
17  $^3\text{He}/^4\text{He}$  to the time-dependent outflow of volatiles from a magmatic chamber subjected  
18 to evolution of its internal pressure due to magma injection. At Mount Etna, the model  
19 makes it possible to estimate in near real time key parameters such as the rate of magma  
20 input and volume change in deep chamber preceding eruptions, and to compare them  
21 with geodetic estimations. This represents an unprecedented use of  $^3\text{He}/^4\text{He}$  to obtain  
22 quantitative information on the physics of magmatic systems. Volcanoes showing

23 changes of  $^3\text{He}/^4\text{He}$  ratio in discharged gases due to unrest episodes are widespread in the  
24 world, **and therefore** we envisage extensive future applications of this approach.

## 25 **INTRODUCTION**

26 Episodes of magma injection into subvolcanic reservoirs are recognized as the  
27 main trigger of the eruptive activity of volcanoes (Caricchi et al., 2014). Along with other  
28 signals, the arrival of the deep magma can be detected by means of increases in the He  
29 isotope ratio ( $^3\text{He}/^4\text{He}$ ) measured in outgassed fluids (Sano and Fischer, 2013, and  
30 references therein). This is possible because the ascending primitive magmas outgas  
31 volatiles having  $^3\text{He}/^4\text{He}$  ratios that are higher than those of fluids from more evolved  
32 melts and/or crust. In the recent case of the Mount Ontake eruption in Japan (Sano et al.,  
33 2015), this tracer was the only one capable of providing clues about increasing activity  
34 over a time scale of years. In spite of this, **the**  $^3\text{He}/^4\text{He}$  ratio has never been used to  
35 achieve physical information of magmatic systems.

36 Here we describe a 12-yr-long time series of the  $^3\text{He}/^4\text{He}$  ratio in volcanic gases  
37 emitted from Mount Etna volcano (Italy) that clearly shows that the main eruptive phases  
38 are preceded by increases in this ratio. An unprecedented use of this tracer allows us to  
39 quantify, in near real time, the flow of magma entering a reservoir beneath the volcano  
40 and the related pressure buildup at depth.

## 41 **DATA AND OBSERVATIONS AT MOUNT ETNA**

42 The  $^3\text{He}/^4\text{He}$  ratios of the gases emitted from 4 vents located along the flanks of  
43 Mount Etna (Fig. 1), 15–20 km from the summit, have been monitored for two decades at  
44 a **nearly** bimonthly sampling rate, yielding a matchless time series. The  $^3\text{He}/^4\text{He}$  ratio of  
45 each vent site fluctuates around its specific average value (Fig. 1), and the decreases in

46 the average  $^3\text{He}/^4\text{He}$  values from **the** P39 to Fd vents cannot be caused by an increasing  
47 contribution of crustal/hydrothermal fluids, as clearly indicated by the geochemical  
48 features of the gases (Caracausi et al., 2003). Instead, the  $^3\text{He}/^4\text{He}$  ratios of the sites  
49 indicate that magmas with isotope ratios lower than that of pristine Etnean magma ( $\sim 7.6$   
50  $R_a$ , where  $R_a = ^3\text{He}/^4\text{He}_{\text{air}}$ ; Correale et al., 2014) are stored in the plumbing system and  
51 feed the vents. En route processes (chiefly,  $^4\text{He}$  production by U and Th radioactive  
52 decay) **reduce** the isotope ratios of Mount Etna resident magmas (Correale et al., 2014).  
53 Three of the monitored sites are separated by only 2–3 km (Fig. 1), making it highly  
54 unlikely that each of them is fed by a distinct magmatic body with its own  $^3\text{He}/^4\text{He}$   
55 marker. More probably, the gas end member from the high  $^3\text{He}/^4\text{He}$  pristine magmas ( $7.6$   
56  $R_a$ ) mixes with a low  $^3\text{He}/^4\text{He}$  end member from resident magmas before feeding the  
57 vents (Fig. 2). The differences in average values among the vents hence reflect different  
58 mixing proportions of these two end members.

59 Figure 1 indisputably shows that increases of 0.2–0.7  $R_a$  occur in all the vents  
60 simultaneously, followed by similar decreases. Considering that one vent is up to 35 km  
61 from the others, the synchronous changes on such a wide scale testify **to** a common deep  
62 process occurring beneath the volcano (Fig. 2). A relative increase in the contribution of  
63 the deep-**origin** **[[SU: correct meaning, vs. “coming”?]]** high  $^3\text{He}/^4\text{He}$  end member with  
64 respect to the low  $^3\text{He}/^4\text{He}$  one explains the growth of isotope ratio at all sites very well.

65 The investigated period represents a phase of considerable activity of Mount Etna  
66 during which the volcano exhibited at least six **major** eruptions **that** occurred in 2001,  
67 2002–2003, 2004–2005, 2006, 2008, and 2011–2012. All of these eruptions were  
68 characterized by considerable volumes of emitted magma ( $>10^7 \text{ m}^3$ ), and most were gas

69 rich and highly explosive, with remarkable fire-fountaining activity (Bonaccorso and  
70 Calvari, 2013). The 2004–2005 eruption was the only one triggered by a passive response  
71 to eastward flank movements, with magmatic overpressure playing a minor role (Burton  
72 et al., 2005). An impressive feature in Figure 1 is that the main eruptive phases are  
73 systematically preceded by increases in  $^3\text{He}/^4\text{He}$  lasting from a few weeks to several  
74 months, with the sole exception of the 2004–2005 event. This emphasizes the potential of  
75 this tracer in forecasting eruptive periods, as already highlighted at Mount Etna (Paonita  
76 et al., 2012, and references therein). Moreover, it strongly suggests that the increases in  
77  $^3\text{He}/^4\text{He}$  (and consequently the increase in the contribution of the high  $^3\text{He}/^4\text{He}$  end  
78 member) are associated with input of high  $^3\text{He}/^4\text{He}$ , gas-rich pristine magma into the  
79 system and consequent pressurization that may lead to eruptions.

## 80 **PHYSICAL MODEL**

81 Each recharge episode thus introduces a volume of the high  $^3\text{He}/^4\text{He}$  magma into  
82 the reservoir that is connected more directly to the source (i.e., the deep reservoir in Fig.  
83 2). Due to this direct connection, such an entering magma will have a very similar He  
84 isotope signature with respect to the magma already stored in the deep reservoir (i.e., 7.6  
85 Ra). The massive entrainment of melts in this confined zone increases the internal  
86 pressure and hence the amount of outgassed fluids, namely an increase in the contribution  
87 of the high  $^3\text{He}/^4\text{He}$  gas end member. Conversely, magma batches stored more marginally  
88 in the system are poorly hydraulically connected, if at all, to the deep reservoir, as  
89 supported by models of plumbing systems operating on time scales of months to years  
90 that include magma compartments (Gudmundsson, 2012). Both the internal pressure and  
91 the low  $^3\text{He}/^4\text{He}$  signature of these distal batches of magmas are not therefore affected by

92 the refilling event and, for the same reason, the output of fluids from these magmas does  
93 not undergo any change. After the end of the magma input, when the overpressure  
94 exhausts, the excess contribution of the high  $^3\text{He}/^4\text{He}$  end member terminates, and the  
95  $^3\text{He}/^4\text{He}$  ratios of the sites return to their background values.

96 We modeled this complex process by coupling a mass balance between the two  
97 gas end members to a physical model of a magma chamber, which is **[[SU: correct?]]** the  
98 part of the deep reservoir that **pressurizes** (Fig. 2; see the Appendix; details are provided  
99 in the GSA Data Repository<sup>1</sup>). **During** periods without input of new magmas  
100 (background), a site **S** **[[SU: italicize variables only; and not subscripts]]** receives  
101 magmatic gases with a measured  $^3\text{He}/^4\text{He}$  ratio ( $R_{S,B}$ ) and He concentration ( $\text{He}_{S,B}$ ). This  
102 already constitutes a mixture of high and low  $^3\text{He}/^4\text{He}$  end members, with the total flow  
103 ( $Q_B$ ) of this mixture being constrained by measurements of the total output of volatiles  
104 from the volcano during quiescent degassing. When a deep input pressurizes the  
105 chamber, the latter releases large amounts of gases with flow rate  $Q_{g,o}(t)$ , **[[SU: define t**  
106 **(time?) here]]** known He concentration  $\text{He}_o$  (Paonita et al., 2012), and having the marker  
107 of the high  $^3\text{He}/^4\text{He}$  end member ( $R_o$ ). **The gases** combine with the previous mixture  
108 according to a mixing fraction  $x_S$ , so as to change  $^3\text{He}/^4\text{He}$  at site S [ $R_S(t)$ ].  $Q_{g,o}(t)$  is fed  
109 by a foam layer that grows in the chamber, and depends on the rock permeability of the  
110 chamber roof, which in turn is controlled by overpressure (Zencher et al., 2006). The  
111 chamber overpressure ( $\Delta P$ ) changes over time as a function of (1) parameters  
112 constrainable by data in **the** literature or **those** having only a minor effect (e.g., chamber  
113 and rock properties; **for explanation**, see the Data Repository) and (2) mix fraction  $x_S$  (not

114 constrainable from external information) and melt input  $Q_{m,i}$  in the chamber (to be  
115 estimated for each recharge event), which are discussed **in the following**.

## 116 **RESULTS AND DISCUSSION**

117 We apply the model to two particular phases of Mount Etna activity: the 2010–  
118 2012 and 2001 events. During the 2010–2012 period, the eruption started while  $^3\text{He}/^4\text{He}$   
119 was increasing, and ended in early 2012 when the ratios were at the minimum values  
120 (Fig. 1). This suggests that the **entire** erupted volume was probably supplied by a single  
121 recharge event. Because Mount Etna has displayed a tendency to maintain an equilibrium  
122 between magma input and output over the past 40 years (Bonaccorso and Calvari, 2013),  
123 the volume that erupted during 2010–2012 ( $50 \pm 10 \times 10^6 \text{ m}^3$ ; mean from Bonaccorso  
124 and Calvari, 2013; Behncke et al., 2014) is a good proxy of the recharge volume from  
125 depth. Divided by the time span of the  $^3\text{He}/^4\text{He}$  variation, we achieve a deep input of melt  
126 of  $\sim 0.95 \text{ m}^3/\text{s}$  (plus the gas input computed from the gas/melt ratio; Table 1; see the Data  
127 Repository). This input **in** the chamber results in the overpressure increasing toward an  
128 asymptotic value (Fig. 3A). Due to the overpressure-dependent rock permeability, the  
129 increase in the gas flow rate from the chamber [ $Q_{g,o}(t)$ ] lags both the overpressure  
130 increase and the gas flow from melt into the foam ( $F_g$ ), thereby producing a sigmoid  
131 curve. As soon as the deep input ends, the chamber elastically returns to its original  
132 volume while the overpressure decays exponentially (Woods and Huppert, 2003) (Fig.  
133 3A).  $Q_{g,o}$  decays even more quickly due to the overpressure-driven decrease in  
134 permeability. Figure 3B shows the calculated  $^3\text{He}/^4\text{He}$  ratio for two sites [ $R_S(t)$ ]. The  
135 simulated and observed signals are surprisingly similar in both shape and time scale. The  
136 model predicts the sharp increase in  $^3\text{He}/^4\text{He}$ , followed by a smooth trend toward an

137 asymptote. When the deep input of magma terminates, the model correctly predicts the  
138  $^3\text{He}/^4\text{He}$  decrease and the restoration of background values. We note that the selected  
139 mixing fraction  $x_S$  (the only parameter we are constraining based on  $^3\text{He}/^4\text{He}$  data) affects  
140 the highest  $^3\text{He}/^4\text{He}$  value reached during the simulation, but not the time scale over  
141 which this is reached. Instead, the time scale is affected by the deep input of magma  
142 entering the chamber. Curves computed with markedly differing values of deep input are  
143 unable to reproduce the observables irrespective of  $x_S$  (see the Data Repository).

144         The model suggests a striking sequence of events during the 2010–2012 period  
145 (Fig. 3B). First, a melt plus gas input started to feed the deep plumbing system in  
146 February 2010, causing the observed increase in  $^3\text{He}/^4\text{He}$  that peaked  $\sim 1$  yr later close to  
147 the maximum possible value under those conditions. This suggests that the overpressure,  
148 and then the ascent of melt toward shallow levels, reached the highest values, and  
149 accordingly the eruptive activity (fire fountaining) started. At this stage the model  
150 predicts a volume increase of  $\sim 2 \times 10^7 \text{ m}^3$  for the chamber. At the end of 2011, while  
151 Mount Etna was still erupting, the  $^3\text{He}/^4\text{He}$  ratio started to decrease dramatically. Our  
152 model suggests that the deep input ceased and, at depth, both the overpressure and  
153 magma ascent to shallow levels were decreasing rapidly. This resulted in the supply of  
154 eruption being extinguished, and accordingly the fire fountaining ended in April 2012.  
155 Supporting evidence comes from areal dilatation signals calculated from data collected at  
156 three summit GPS stations and three intermediate altitude ones, which are considered as  
157 representative of shallow and deep volume changes in Mount Etna plumbing system,  
158 respectively (Aloisi et al., 2011). The input that occurred during February 2010 matches a  
159 change in the rate of dilatation at intermediate altitude, whereas there was negligible

160 deformation of the summit area (Patanè et al., 2013). This pattern of observations was  
161 attributed to an active inflation source located 8 km below sea level (bsl) under the  
162 summit craters (Patanè et al., 2013), in agreement with the depth of our modeled chamber  
163 (see Table 1). The observed deformation was generated by a volume change of  $\sim 5 \times 10^7$   
164  $\text{m}^3$  (Patanè et al., 2013), which is amazingly close to the volume increase estimated by  
165 our model. The input started some weeks after the deep (20–30 km) earthquake swarm  
166 affecting the northwestern sector of the volcano on 19 December 2009, which was the  
167 most intense event during that period (Patanè et al., 2013). Volcano-tectonic swarms that  
168 occur at these levels are thought to be predictive of magmatic replenishments feeding  
169 future volcanic activities (Patanè et al., 2013).

170         The rate of  $^3\text{He}/^4\text{He}$  increase and its absolute value were highest during a violent  
171 eruption that occurred in July–August 2001 (Fig. 1). This sharp peak followed a smoother  
172 increase that started several months earlier, suggesting that the paroxysmal recharge  
173 occurred during a longer period of slower injection of magma. This means that we cannot  
174 use the erupted volume as a proxy of the recharge volume to constrain the deep input of  
175 magma. Instead, we utilize the constrained  $x_s$  value during 2010–2012, and adjust the  
176 deep melt input  $Q_{m,i}$  to fit the  $^3\text{He}/^4\text{He}$  increase that occurred in July–August 2001. As a  
177 result, the  $^3\text{He}/^4\text{He}$  increase was due to a deep magmatic input much larger ( $\sim 2.7 \text{ m}^3/\text{s}$ )  
178 than that predicted for 2010–2012 (Fig. 3D). It further agrees with the magma emission  
179 rates observed during the 2001 eruption, which were the highest of the entire investigated  
180 period (Behncke and Neri, 2003). The input started in May (Figs. 3C and 3D),  $\sim 2$  weeks  
181 after the seismic swarm of 22 April 2001, interpreted as an intrusive process occurring at  
182  $\sim 8$  km bsl (Bonaccorso et al., 2004). The onset of the 2001 eruption on 17 July was



183 anticipated by an impressive seismic swarm (12–17 July; Figs. 3C and 3D) and abrupt  
184 ground deformation **that** indicated the propagation of a dike from the shallow reservoir at  
185 3 km **bsl** (Neri et al., 2005), which overtops our modeled chamber (Fig. 2). Our  
186 simulation predicts that, at the moment of the seismic swarm, the overpressure in the  
187 deep chamber exceeded the tensile strength of the local wall rocks (Fig. 3C; see the  
188 Appendix). The connection between the deep and shallow reservoirs also suggests that  
189 the latter was seriously overpressurized.

190 Modeling the effect of dike opening on overpressure and gas release from the  
191 deep chamber is outside the aim of this work. The high  $^3\text{He}/^4\text{He}$  of olivine-hosted fluid  
192 inclusions that erupted in 2001 suggests that the new dike carried high  $^3\text{He}/^4\text{He}$  magmas  
193 **directly** to the surface (Correale et al., 2014), explaining the amazingly high  $^3\text{He}/^4\text{He}$  ratio  
194 at vent P39. **In addition**, the abrupt outpouring of magma after dike opening guaranteed a  
195 drop in the overpressure. The consequent decrease in gas output  $Q_{g,o}(t)$  from the chamber  
196 could explain the very high rate of  $^3\text{He}/^4\text{He}$  decrease (higher than model prediction; Fig.  
197 3D) observed between 7 and 10 August, when the eruption was **ending**.

## 198 CONCLUSION

199 This work highlights that temporal variation of  $^3\text{He}/^4\text{He}$  **is** a valuable indicator for  
200 assessing the level of volcanic activity and can provide key physical parameters such as  
201 the magma input rate and overpressure in the chamber, which until now have only been  
202 available from geodetic approaches. The conditions for chamber stability and wall-rock  
203 failure can be determined, **and** will improve the accuracy of eruption forecasting. Our  
204 approach is predicted to be widely applicable, **because** time-dependent He isotope mixing  
205 between primitive and more radiogenic end members appears to be common in many

206 active volcanoes (Sano and Fischer, 2013). Long-lasting time series with sufficiently  
207 frequent samplings and high-precision  $^3\text{He}/^4\text{He}$  measurements in air-free volcanic gases  
208 are required, since even small isotope variations (fractions of 1 Ra unit) can reflect  
209 important volcanic processes.

## 210 **APPENDIX: MODEL AND PARAMETERS**

211 Here we recap the main equations of the model (for full details, including  
212 parameterization and sensitivity study, see the Data Repository). The  $^3\text{He}/^4\text{He}$  ratio at site  
213 S, after mathematical manipulation, can be computed using[[SU: Note, the only  
214 characters that should be italicized in Equation 1 are variables R, Q, t, and x;  
215 subscripts and He (helium) are not italicized]]

$$216 \quad R_S(t) = \frac{0.5Q_B \text{He}_{S,B} R_{S,B} + Q_{g,o}(t) \text{He}_o R_o x_S}{0.5Q_B \text{He}_{S,B} + Q_{g,o}(t) \text{He}_o x_S} \quad (1)$$

217 (all symbols are defined in the text). Calculation of  $Q_{g,o}(t)$ , i.e., the gas outflow from the  
218 chamber, requires a physical model of the reservoir. We modified the approach of Woods  
219 and Huppert (2003) to account for the input and output of gases, as well the formation of  
220 a foam layer. This approach has no pretensions of being complete, but it does capture the  
221 key processes governing chamber overpressure (Woods and Huppert, 2003). Pressure and  
222 volume changes in the chamber can be described by:

$$223 \quad f \frac{d\Delta P}{dt} = \frac{d\Delta V}{dt} = Q_{g,i} + Q_{m,i} - Q_{g,o} - Q_{m,o}, \quad (2)$$

224 where subscripts m, g, i, and o refer to melt, gas, input, and output volumetric flows into  
225 and out from the chamber, respectively, and  $f$  is the chamber volume divided by the  
226 effective compressibility of the melt-gas-rock system, which depends on gas fraction  $n_g$

227 in the chamber (Woods and Huppert, 2003). The value of  $n_g$  depends on the difference  
228 between the input and output of gas:

$$229 \quad \frac{dn_g}{dt} = \frac{\rho_g}{M_m} (Q_{g,i} - Q_{g,o}), \quad (3)$$

230 where  $M_m$  is the mass of magma (with density  $\rho_m$ ) in chamber volume  $V_{ch}$ , and  $\rho_g$  is the  
231 gas density. At any time  $n_g$  will be shared between dispersed bubbles in the melt plus a  
232 fraction in the foam. If  $n_b$  is the amount of gas in the dispersed bubbles, it evolves as  
233 (Woods and Cardoso, 1997):

$$234 \quad \frac{dn_b}{dt} = \frac{Q_{g,i}\rho_g}{M_m} - \frac{v_b}{H} n_b, \quad (4)$$

235 where  $v_b$  is speed of bubble ascent (from Stokes's law). Output flows  $Q_{g,o}$  and  $Q_{m,o}$  are  
236 functions of  $\Delta P$ :

$$237 \quad Q_{g,o} = \frac{k(\Delta P) \cdot A_{ch}}{\mu_g} \left[ \frac{\Delta P}{h} + (\rho_m - \rho_g)g \right], \quad (5)$$

$$238 \quad Q_{m,o} = \frac{\pi r^4}{8\mu_m Z} \Delta P, \quad (6)$$

239 where  $k(\Delta P)$  is the overpressure-dependent permeability (Zencher et al., 2006),  $h$  is a  
240 characteristic vertical distance from the chamber roof at which the effect of  $\Delta P$   
241 disappears,  $A_{ch}$  is the roof area,  $\mu_m$  and  $\mu_g$  are the melt and gas viscosities, respectively,  $r$   
242 and  $Z$  are the radius and height of the output conduit of melt, respectively, and  $g$  is  
243 gravity.

244 All of the required parameter values are listed in Table 1.

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248 **REFERENCES CITED**

- 249 Aiuppa, A., et al., 2008, Total volatile flux from Mount Etna: Geophysical Research  
250 Letters, v. 35, L24302, doi:10.1029/2008GL035871.
- 251 Aloisi, M., Mattia, M., Monaco, C., and Pulvirenti, F., 2011, Magma, faults, and  
252 gravitational loading at Mount Etna: The 2002–2003 eruptive period: Journal of  
253 Geophysical Research, v. 116, B05203, doi:10.1029/2010JB007909.
- 254 Behncke, B., and Neri, M., 2003, The July–August 2001 eruption of Mt. Etna (Sicily):  
255 Bulletin of Volcanology, v. 65, p. 461–476, doi:10.1007/s00445-003-0274-1.
- 256 Behncke, B., Branca, S., Corsaro, R.A., De Beni, E., Miraglia, L., and Proietti, C., 2014,  
257 The 2011–2012 summit activity of Mount Etna: Birth, growth and products of the  
258 new SE crater: Journal of Volcanology and Geothermal Research, v. 270, p. 10–21,  
259 doi:10.1016/j.jvolgeores.2013.11.012.
- 260 Bonaccorso, A., and Calvari, S., 2013, Major effusive eruptions and recent lava  
261 fountains: Balance between expected and erupted magma volumes at Etna volcano:  
262 Geophysical Research Letters, v. 40, p. 6069–6073, doi:10.1002/2013GL058291.
- 263 Bonaccorso, A., D’Amico, S., Mattia, M., and Patanè, D., 2004, Intrusive mechanisms at  
264 Mt. Etna forerunning the July–August 2001 eruption from seismic and ground  
265 deformation data: Pure and Applied Geophysics, v. 161, p. 1469–1487,  
266 doi:10.1007/s00024-004-2515-4.
- 267 Bonaccorso, A., Cianetti, S., Giunchi, C., Trasatti, E., Bonafede, M., and Boschi, E.,  
268 2005, Analytical and 3-D numerical modelling of Mt. Etna (Italy) volcano inflation:  
269 Geophysical Journal International, v. 163, p. 852–862, doi:10.1111/j.1365-  
270 246X.2005.02777.x.

- 271 Burton, M., et al., 2005, Etna 2004–2005: An archetype for geodynamically-controlled  
272 effusive eruptions: *Geophysical Research Letters*, v. 32, L09303,  
273 doi:10.1029/2005GL022527.
- 274 Caracausi, A., Italiano, F., Nuccio, P.M., Paonita, A., and Rizzo, A., 2003, Evidence of  
275 deep magma degassing and ascent by geochemistry of peripheral gas emissions at  
276 Mt. Etna (Italy): Assessment of the magmatic reservoir pressure: *Journal of*  
277 *Geophysical Research*, v. 108, no. B10, 2463, doi:10.1029/2002JB002095.
- 278 Caricchi, L., Annen, C., Blundy, J., Simpson, G., and Pinel, V., 2014, Frequency and  
279 magnitude of volcanic eruptions controlled by magma injection and buoyancy:  
280 *Nature Geoscience*, v. 7, p. 126–130, doi:10.1038/ngeo2041.
- 281 Correale, A., Paonita, A., Martelli, M., Rizzo, A., Rotolo, S.G., Corsaro, R.A., and Di  
282 Renzo, V., 2014, A two-component mantle source feeding Mt. Etna magmatism:  
283 Insights from the geochemistry of primitive magmas: *Lithos*, v. 184, p. 243–258,  
284 doi:10.1016/j.lithos.2013.10.038.
- 285 Gudmundsson, A., 2012, Magma chambers: Formation, local stresses, excess pressures,  
286 and compartments: *Journal of Volcanology and Geothermal Research*, v. 237–238,  
287 p. 19–41, doi:10.1016/j.jvolgeores.2012.05.015.
- 288 Jamtveit, B., and Yardley, B., 1997, Fluid flow and transport in rocks: Mechanisms and  
289 effects: London, UK, Chapman & Hall, 315 p.
- 290 Johnson, S.E., Jin, Z.-H., Naus-Thijssen, F.M.J., and Koons, P.O., 2011, Coupled  
291 deformation and metamorphism in the roof of a tabular midcrustal igneous complex:  
292 *Geological Society of America Bulletin*, v. 123, p. 1016–1032,  
293 doi:10.1130/B30269.1.

- 294 Neri, M., Acocella, V., Behncke, B., Maiolino, V., Ursino, A., and Velardita, R., 2005,  
295 Contrasting triggering mechanisms of the 2001 and 2002–2003 eruptions of Mount  
296 Etna (Italy): *Journal of Volcanology and Geothermal Research*, v. 144, p. 235–255,  
297 doi:10.1016/j.jvolgeores.2004.11.025.
- 298 Paonita, A., Caracausi, A., Iacono-Marziano, G., Martelli, M., and Rizzo, A., 2012,  
299 Geochemical evidence for mixing between fluids exsolved at different depths in the  
300 magmatic system of Mt Etna (Italy): *Geochimica et Cosmochimica Acta*, v. 84,  
301 p. 380–394, doi:10.1016/j.gca.2012.01.028.
- 302 Patanè, D., et al., 2013, Insights into magma and fluid transfer at Mount Etna by a  
303 multiparametric approach: A model of the events leading to the 2011 eruptive cycle:  
304 *Journal of Geophysical Research*, v. 118, p. 1–21, doi:10.1002/jgrb.50248.
- 305 Sano, Y., and Fischer, P.T., 2013, The analysis and interpretation of noble gases in  
306 modern hydrothermal systems, *in* Burnard, P., ed., *The noble gases as geochemical*  
307 *tracers*: Berlin, Springer-Verlag, p. 249–317, doi:10.1007/978-3-642-28836-4\_10.
- 308 Sano, Y., Kagoshima, T., Takahata, N., Nishio, Y., Roulleau, E., Pinti, D.L., and Fischer,  
309 P., 2015, Ten-year helium anomaly prior to the 2014 Mt Ontake eruption: *Scientific*  
310 *Reports*, v. 5, 13069, doi:10.1038/srep13069.
- 311 Woods, A.W., and Cardoso, S.S.S., 1997, Bubble magma separation as a trigger of  
312 basaltic volcanic eruptions: *Nature*, v. 385, p. 518–520, doi:10.1038/385518a0.
- 313 Woods, A.W., and Huppert, H.E., 2003, On magma chamber evolution during slow  
314 effusive eruptions: *Journal of Geophysical Research*, v. 108, 2403,  
315 doi:10.1029/2002JB002019.

316 Zencher, F., Bonafede, M., and Stefansson, R., 2006, Near-lithostatic pore pressure at  
317 seismogenic depths: A thermoporoelastic model: *Geophysical Journal International*,  
318 v. 166, p. 1318–1334, doi:10.1111/j.1365-246X.2006.03069.x.

319 **FIGURE CAPTIONS**

320 Figure 1. Time variations of  $^3\text{He}/^4\text{He}$  measured at P39, St, VS, and Fd gas vents of Mount  
321 Etna (Italy) during the period 2001–2012 (locations shown in inset; see Paonita et al.,  
322 2012, for sampling and analytical methods). Sampling times are the same for all sites and  
323 are shown only for VS. Isotope values are in units of  $R/R_a = (^3\text{He}/^4\text{He}_{\text{sample}})/(^3\text{He}/^4\text{He}_{\text{air}})$ ,  
324 and are corrected for (very small) air contamination (Caracausi et al., 2003). The values  
325 of mean  $R/R_a \pm 1\sigma$  at P39, VS, St, and Fd are  $7.05 \pm 0.17$ ,  $6.57 \pm 0.14$ ,  $6.21 \pm 0.15$ , and  
326  $6.17 \pm 0.14$ , respectively. Red lines highlight the eruptive phases; grey areas indicate the  
327 modeled periods.

328

329 Figure 2. Sketch of the plumbing system, with magma-plus-fluids storage zones (yellow)  
330 and the crystalline body (blue; modified from Patanè et al., 2013). The shallow (3 km  
331 below sea level, bsl) and deep (5–12 km bsl) reservoirs (res.) are from Paonita et al.  
332 (2012); dark red to orange indicate magmas with high to low  $^3\text{He}/^4\text{He}$ . Dotted arrows  
333 mark the magma ascent pathway. Brown and orange arrows indicate pathways of high  
334 and low  $^3\text{He}/^4\text{He}$  gaseous end members, respectively. The high  $^3\text{He}/^4\text{He}$  magma chamber  
335 is shown (ellipse). The right diagram provides details of the magma chamber (symbols as  
336 in the text; also see Table 1). Fumar.—fumarole. **[[SU: symbols and variables should  
337 be as in text (subscripts and Greek characters should not be italicized, but P and t  
338 should be).]]**

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340 Figure 3. A: Time variation of overpressure  $\Delta P$ , gas flow into foam  $F_g$ , and gas output  
341 from chamber  $Q_{g,o}$ . B: Time variation of  $^3\text{He}/^4\text{He}$  of gases from vents St and P39,  
342 computed for the 2010–2012 input. C, D: Same as A and B, but computed for the 2001  
343 input (P39 site only). B and D also show the measured  $^3\text{He}/^4\text{He}$  signals. The start and stop  
344 of magma input are indicated. Thick curves for vents St and P39 in B differ by the  
345 selected background and  $x_S$  values (see Table 1). Background for each site is the  
346 minimum  $^3\text{He}/^4\text{He}$  ratio of the time series. Thin dotted curves for St show the effect of  
347 changing  $x_S$ : 0.25 (higher) and 0.15 (lower) versus 0.18 (best fit; see Table 1). Eruptions  
348 are shown as red bars. **[[SU: Need uppercase A–D in figures; “g,o” after Q should be**  
349 **subscripts, Q should be italicized, g after F should be subscript, F should be**  
350 **italicized. On left axes and in legends of A and C, “P” should be italicized; “mar” at**  
351 **bottom of D should be spelled out (March).]]**

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353 <sup>1</sup>GSA Data Repository item 2016xxx, **[[SU: Need DR item names and brief**  
354 **descriptions here]]**, is available online at [www.geosociety.org/pubs/ft2016.htm](http://www.geosociety.org/pubs/ft2016.htm), or on  
355 request from [editing@geosociety.org](mailto:editing@geosociety.org) or Documents Secretary, GSA, P.O. Box 9140,  
356 Boulder, CO 80301, USA.

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TABLE 1. MODEL PARAMETERS

<b>Chamber:</b>	
height ( $H$ ), volume ( $V_{ch}$ )	1 km <sup>†</sup> , 4.0 km <sup>3 [1]</sup>
temperature ( $T_m$ )	1150 °C <sup>[2]</sup>
pressure ( $P_m$ ), depth ( $d$ )	300 MPa <sup>[2]</sup> , 8.0 km bsl <sup>*</sup>
dissolved H <sub>2</sub> O, CO <sub>2</sub>	3.5 wt%, 1800 ppm <sup>[2]</sup>
H <sub>2</sub> O, CO <sub>2</sub> in vapor	35 mol%, 65 mol% <sup>[2]</sup>
radius of bubbles	200 μm <sup>[5]</sup>
<b>Conduit:</b>	
radius ( $r$ ), height ( $Z$ )	0.2 m*, 3.0 km*
<b>Density:</b>	
melt ( $\rho_m$ ), gas ( $\rho_g$ )	2600 kg/m <sup>3*</sup> , 870 kg/m <sup>3*</sup>
<b>Viscosity:</b>	
melt ( $\mu_m$ ), gas ( $\mu_g$ )	3.0 Pa·s <sup>[5]</sup> , 5 × 10 <sup>-6</sup> Pa·s <sup>[5]</sup>
<b>Young's modulus (<math>G</math>)</b>	15 GPa <sup>[3]</sup>
<b>Poisson's ratio (<math>\nu</math>)</b>	0.25 <sup>[3]</sup>
<b>Melt bulk modulus (<math>\beta_m</math>)</b>	10 GPa*
<b>Distance (<math>h</math>)</b>	200 m <sup>[4]</sup>
<b>Deep input:</b>	
melt ( $Q_{m,i}$ ) in 2010–2012	0.95 m <sup>3</sup> /s <sup>[5]</sup>
melt ( $Q_{m,i}$ ) in 2001	2.7 m <sup>3</sup> /s*
gas/melt ratio ( $Q_g/Q_{m,i}$ )	0.5 <sup>[5,6]</sup>
<b>Gas output from chamber:</b>	
He concentration ( $He_c$ )	6.0 ppm <sup>[2]</sup>
<sup>3</sup> He/ <sup>4</sup> He ( $R_c$ )	7.6 R/Ra <sup>[7]</sup>
<b>Rock permeability:</b>	
constant permeability ( $k^o$ )	5 × 10 <sup>-18</sup> m <sup>2 [8]</sup>
fracture half-length ( $l$ )	2.0 mm <sup>[9]</sup>
fracture distance ( $D$ )	0.5 m <sup>[9]</sup>
<b>Background gas flow:</b>	
Gas flow rate ( $Q_B$ )	0.14 m <sup>3</sup> /s <sup>[6]</sup>
He concentration ( $He_{S,B}$ )	6.0 ppm <sup>[2]</sup>
<sup>3</sup> He/ <sup>4</sup> He at vent St ( $R_{S,B}$ )	5.96 R/Ra*
<sup>3</sup> He/ <sup>4</sup> He at vent P39 ( $R_{S,B}$ )	6.76 R/Ra*
mix fraction at St ( $x_S$ )	0.18*
mix fraction at P39 ( $x_S$ )	0.88*

*Note:* bsl—below sea level. References in superscript brackets: 1—Bonaccorso et al., 2005; 2—Paonita et al., 2012; 3—Aloisi et al., 2011; 4—Johnson et al., 2011; 5—Bonaccorso and Calvari, 2013; 6—Aiuppa et al., 2008; 7—Correale et al., 2014; 8—Jamtveit and Yardley, 1997; 9—Zencher et al. 2006; S—Other sources (see text footnote 1[[SU: which DR item?]]).

\*Computed.

†Assumed in this work.

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[[SU: what should highlighted superscript be before 18? What is the superscript o in permeability  $k^o$ ?]]